

Search for BSM physics in di-photon final states at CMS

C.Rovelli, INFN Roma
On behalf of the CMS Collaboration

ICHEP 2016, Chicago Aug 5th

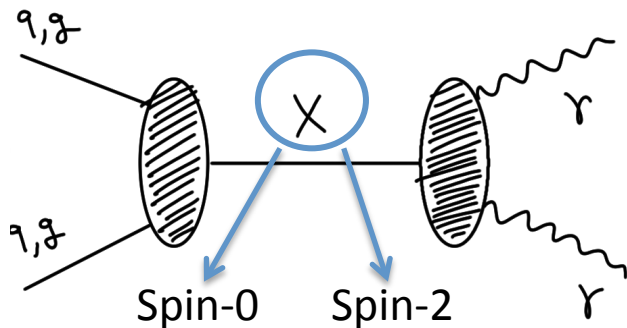
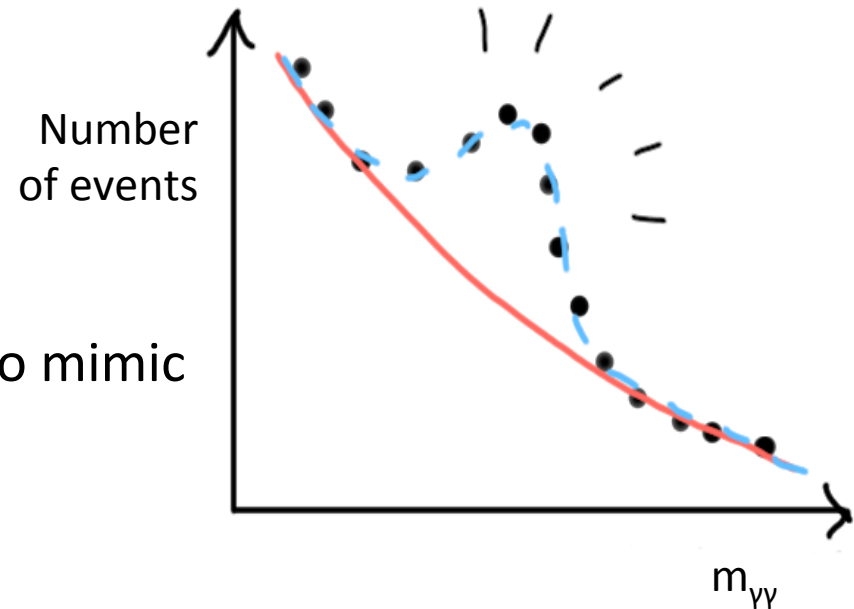
Why di-photon searches

Fully reconstructed resonances: simplest way to discover new particles

Statistically significant peak
over a smooth background

- experimentally robust
- small systematics
- difficult for unknown backgrounds to mimic

⇒ *simple yet striking signature!*

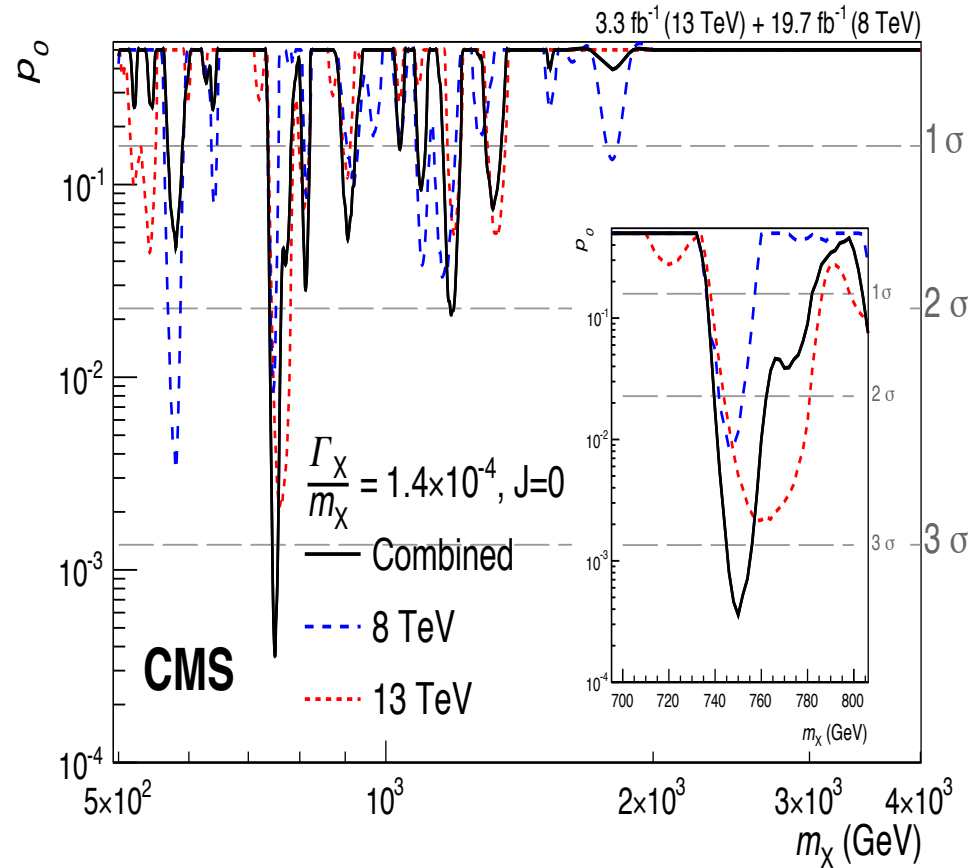
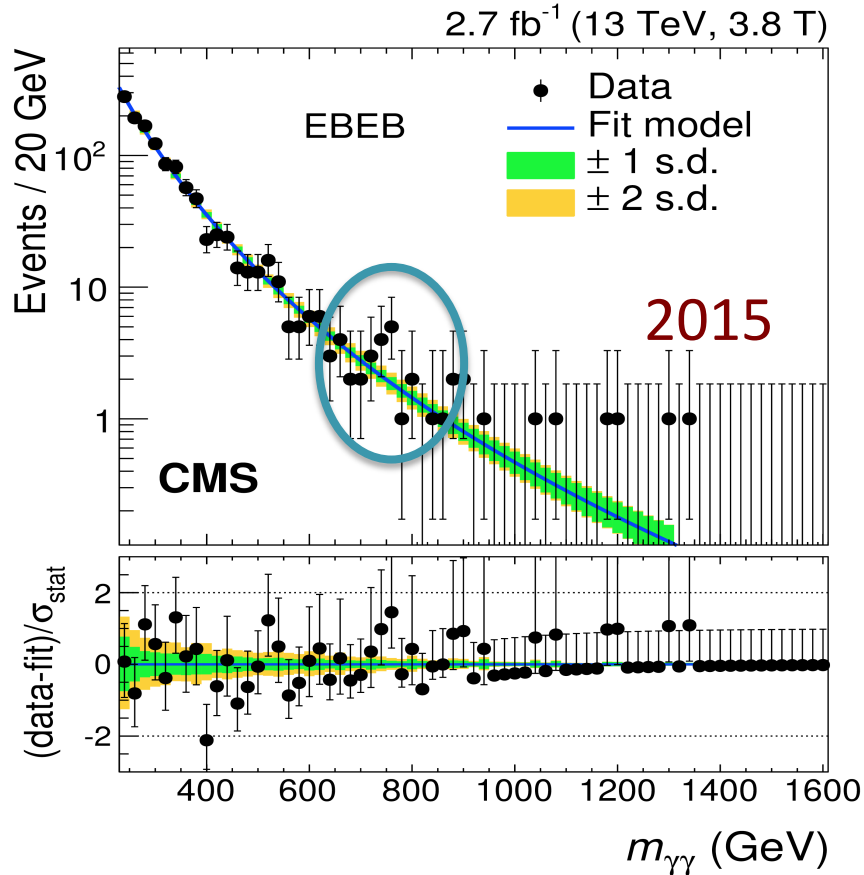


Final states with high p_T photons:

- relatively low background at hadron colliders
- good mass resolution

Recap of 2015 results

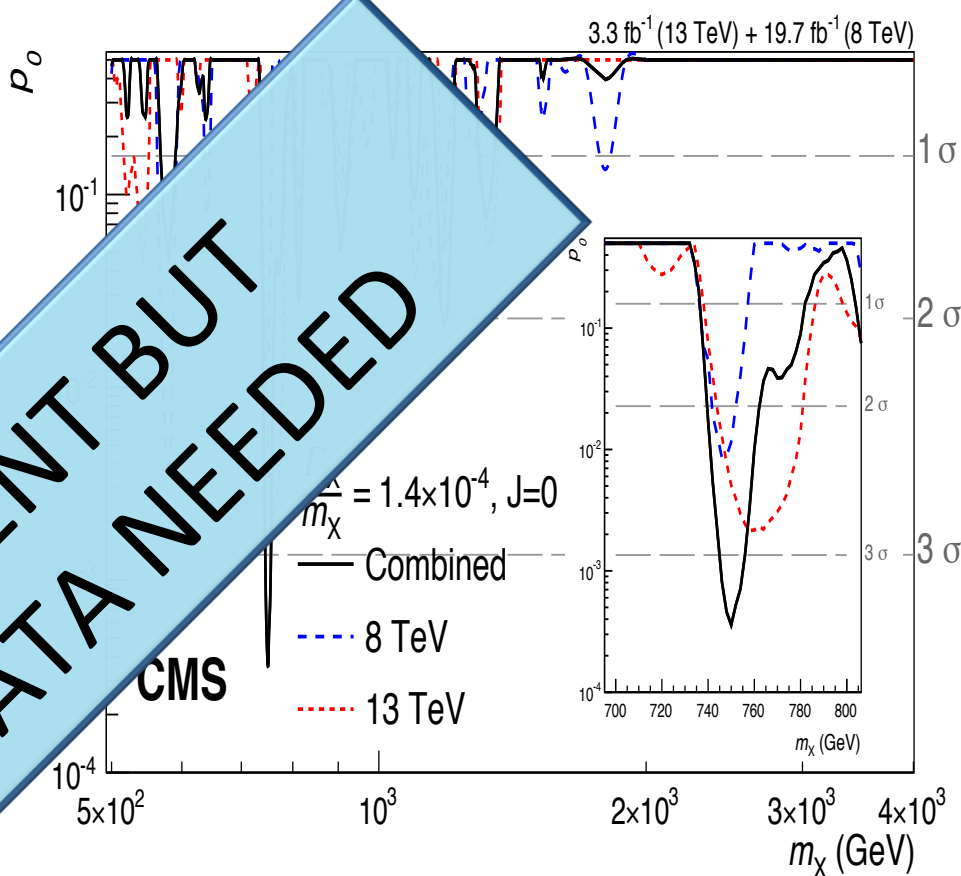
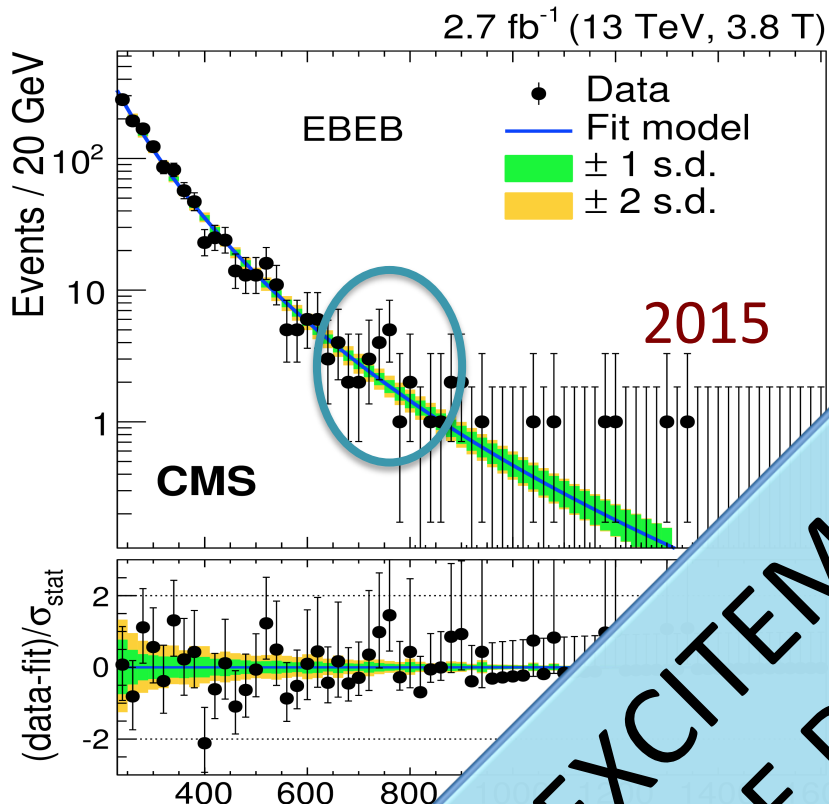
Phys.Rev.Let. 117(2016), no. 5, 051802



	2015	2015 + 8TeV
Mass (GeV)	760	750
Local significance	2.9σ	3.4σ
Global significance	<1σ	1.6σ

Recap of 2015 results

Phys.Rev.Let. 117(2016), no. 5, 051802



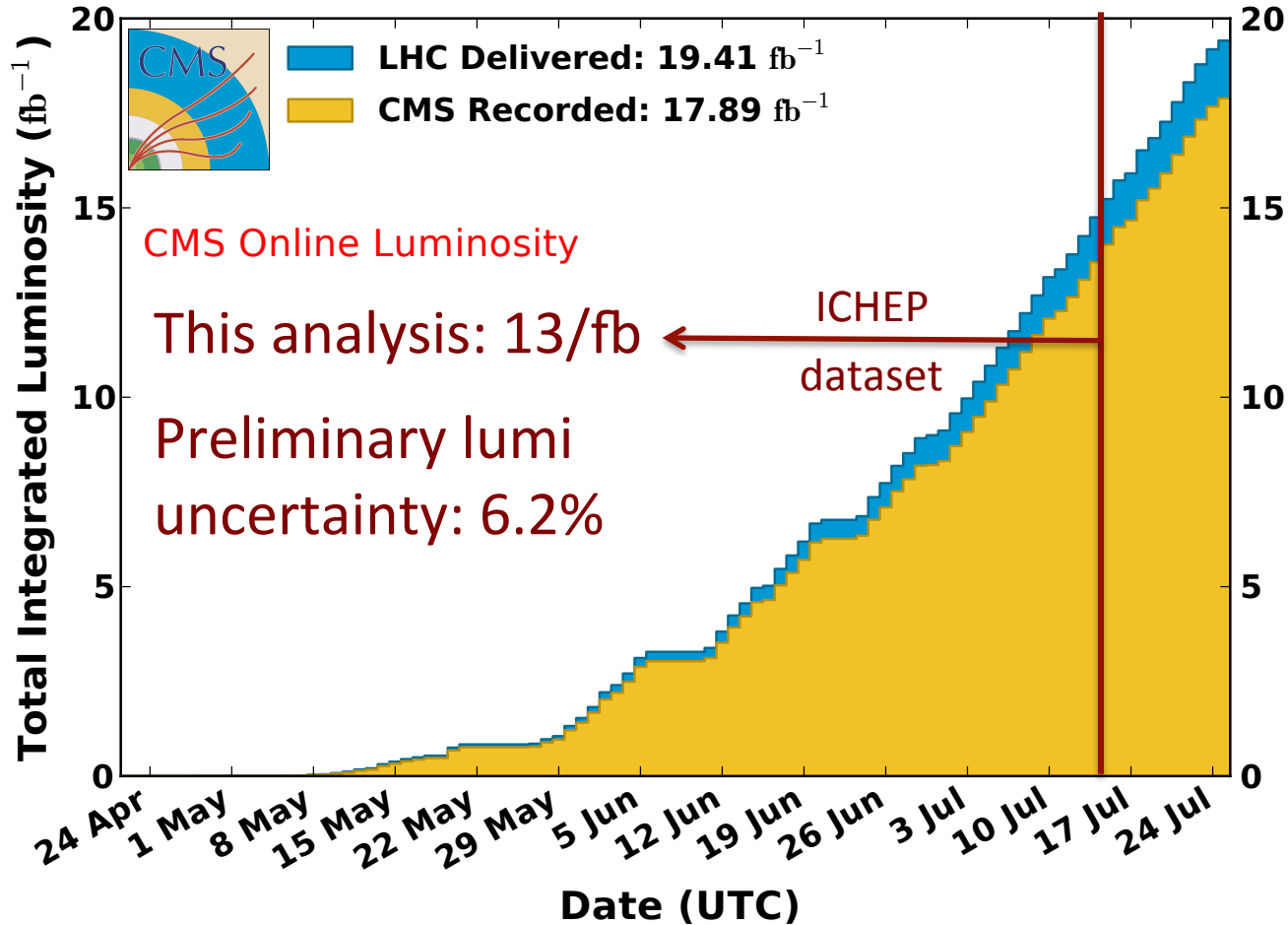
EXCITEMENT BUT
MORE DATA NEEDED

	2015	2015 + 8TeV
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Global significance	<1 σ	1.6 σ

CMS 2016 dataset

CMS Integrated Luminosity, pp, 2016, $\sqrt{s} = 13$ TeV

Data included from 2016-04-22 22:48 to 2016-07-25 21:26 UTC



Thanks to LHC wonderful performance, now we have more data⁵

The analysis in a nutshell

Search for a localized excess in the di-photon invariant mass spectrum

Events selection: 2 high p_T isolated photons

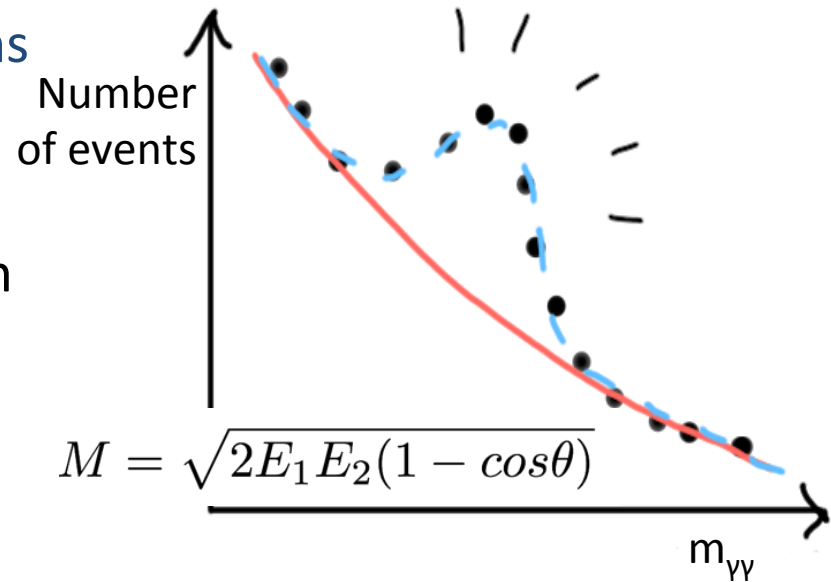
- Robust criteria

Signal modeling with data-driven inputs

- Efficiencies, energy scale and resolution
- Good detector understanding needed

Background modeling

- Parameterization from data



The analysis in a nutshell

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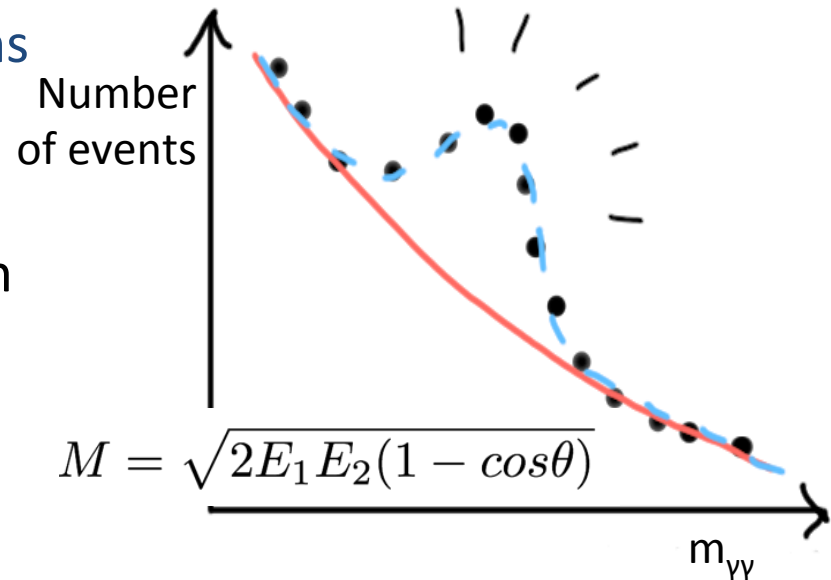
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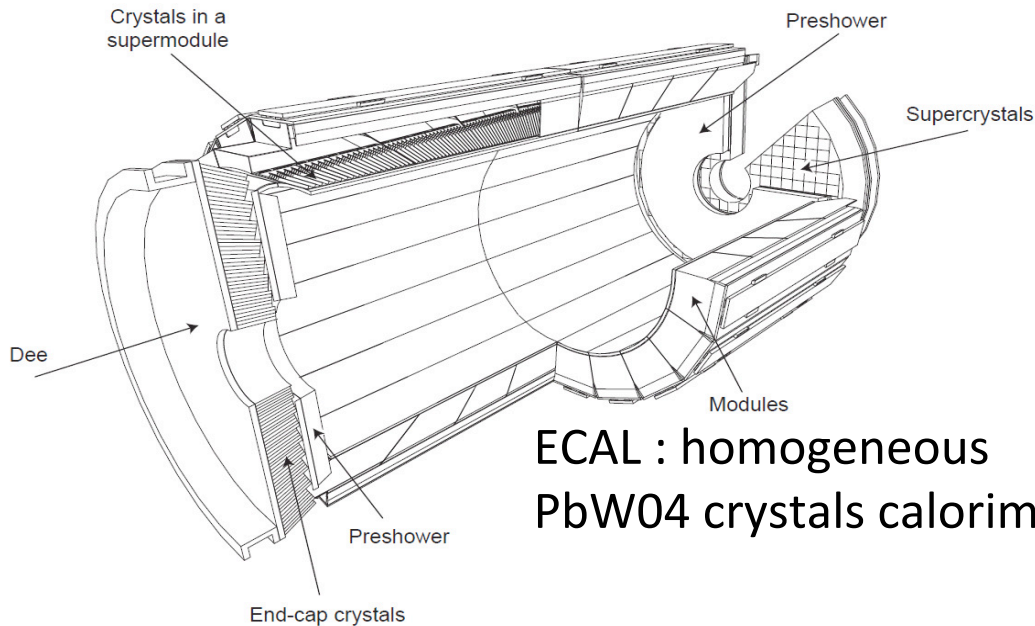


Solid techniques exploited

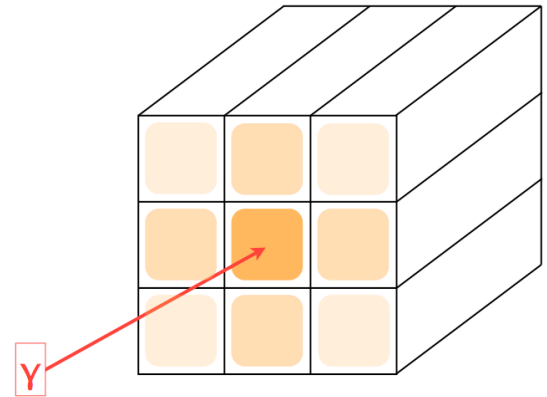
2016 data blinded to investigate potential improvements wrt 2015

- No significant gain found
- No selection change applied: basically same analysis as last year

Photons in CMS



ECAL : homogeneous PbW04 crystals calorimeter



Photons reconstructed from energy deposits in clusters of ECAL crystals

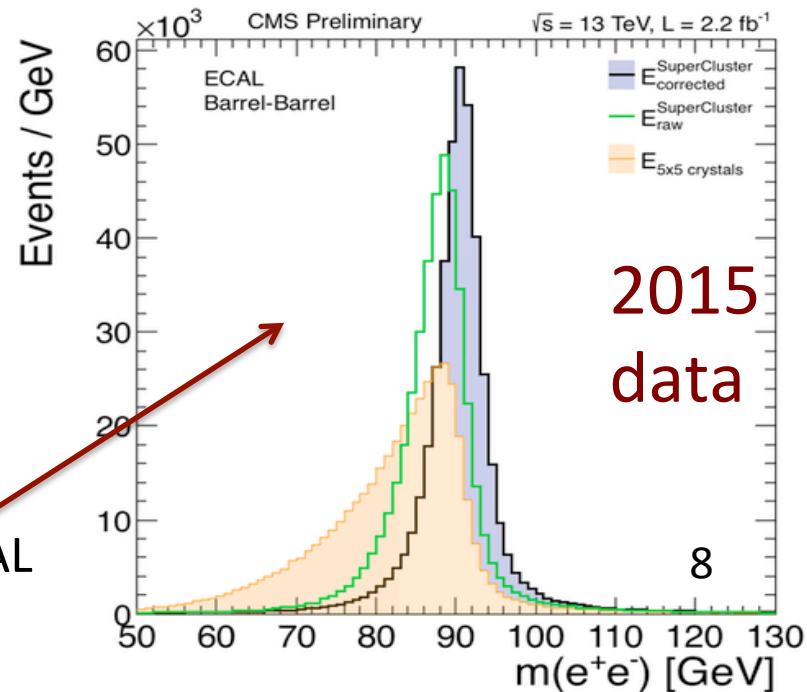
- No associated track

97% energy contained in a 5x5 matrix

Clustering optimized to collect radiated energy

- up to $2X_0$ distributed material in front of ECAL

Energy from multivariate regression



Selection criteria

High Level Trigger

2 photons, $p_T > 60$ GeV

Offline kinematic selection

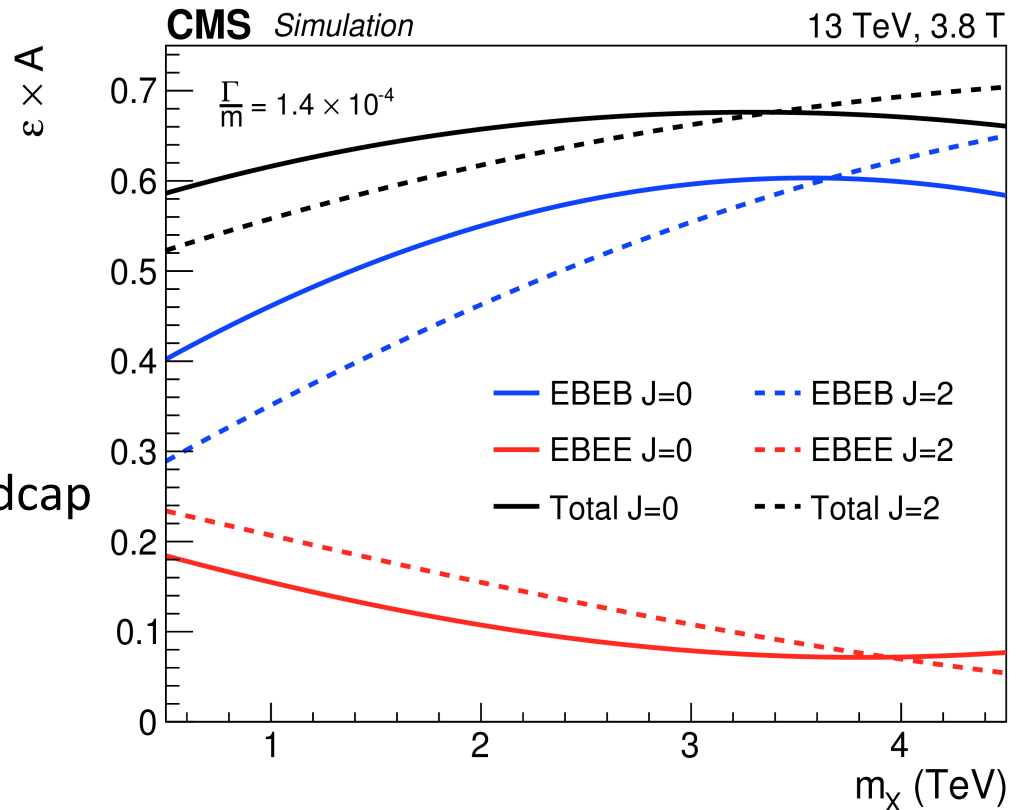
- Fixed p_T cut, $p_T > 75$ GeV
- ECAL fiducial region ($|\eta| < 2.5$)

2 event categories

- EBEB: both γ s in the barrel
- EBEE: one γ in barrel, one in endcap
- Based on S/B ratio

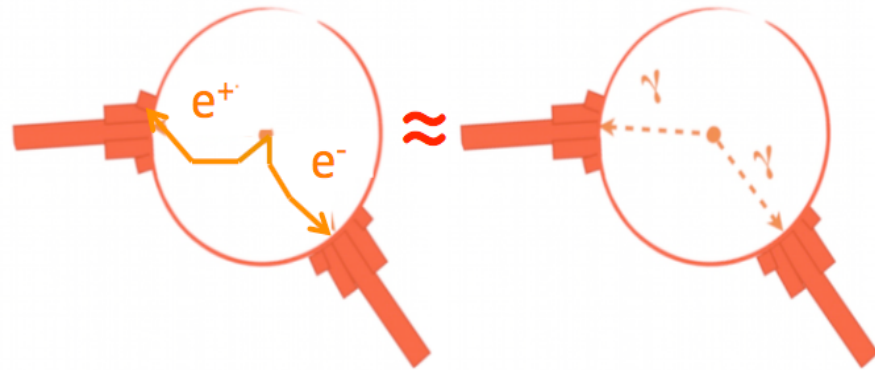
Identification requirements:

- Shower shape and isolation
- No associated track
- $\gamma\gamma$ selection efficiency:
 - $\sim 85\%$ in EBEB
 - $\sim 80\%$ in EBEE



Small differences due to different Kinematics for different spin hypothesis

Efficiency measurement

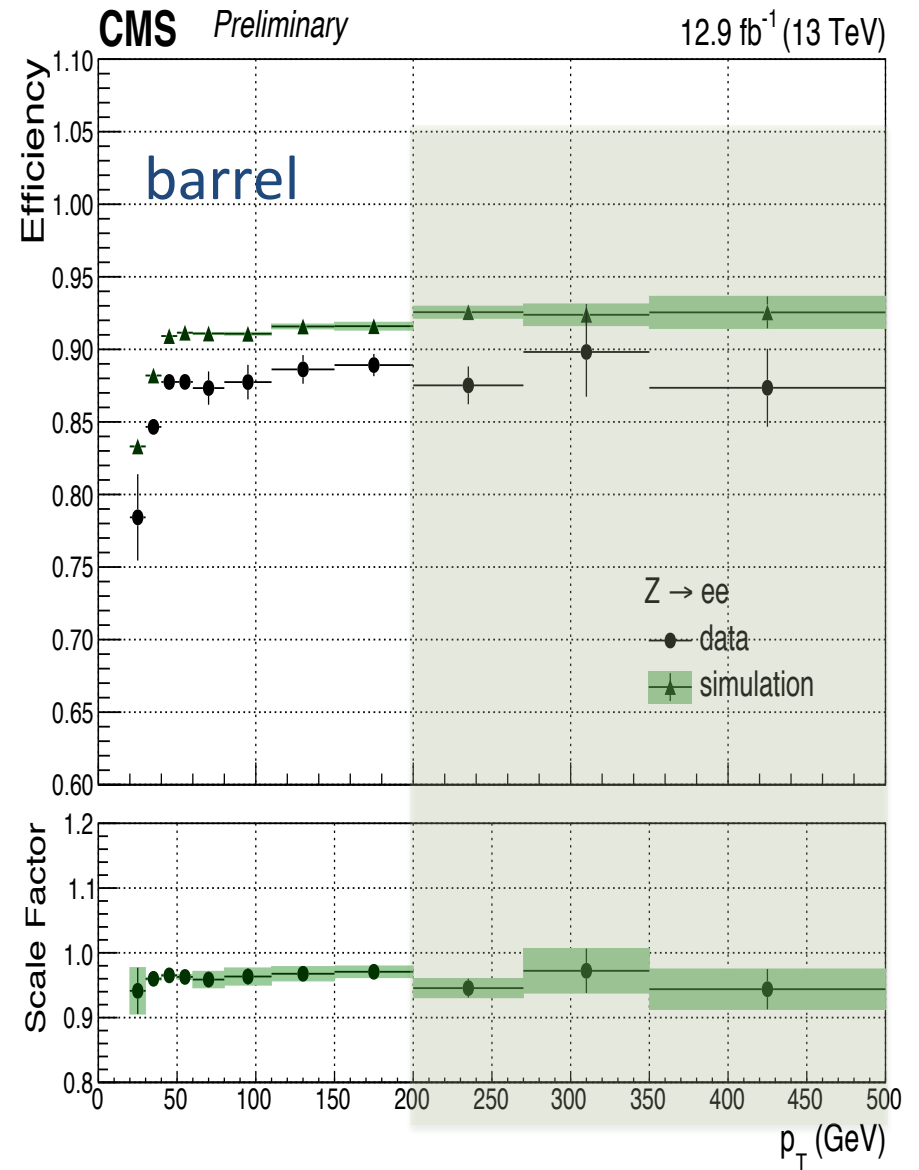


Tuned in MC and validated in data

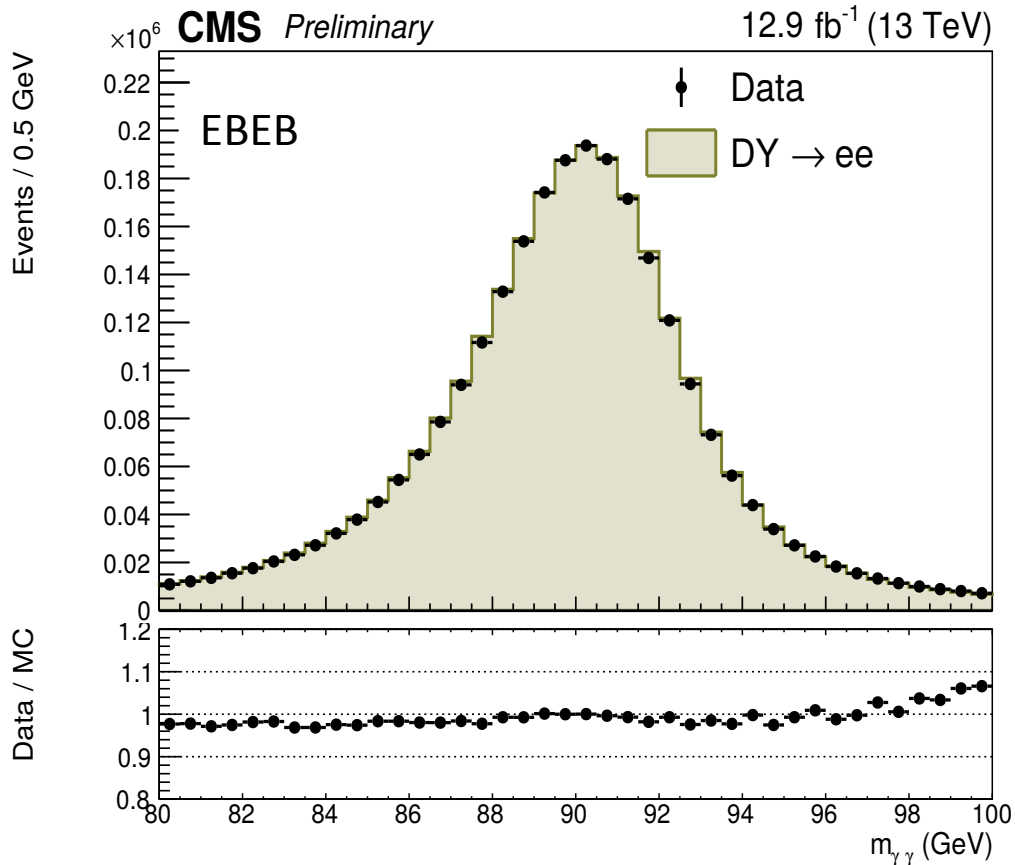
- $Z \rightarrow ee$ events
- $Z \rightarrow \mu\mu\gamma$ events for electron veto requirement

Data/MC scale factors

- computed for EB and EE
- almost constant in p_T
- non compatible with 1
 - included in signal modeling



Photon energy scale and resolution



Measured with Z in data at O(0.1%) level in bins of η and cluster shape

- simultaneously adjust scale and resolution

Good agreement in term of shape and normalization after all corrections

Energy scale stability vs E_T checked with boosted Zs up to ~ 150 GeV

- Deviations within a few permilles
- 1% uncertainty to account for further extrapolation

Signal modeling

Statistical interpretation from simultaneous fit to $m_{\gamma\gamma}$ distribution in the two analysis categories

Signal shape: convolution of intrinsic line-shape (Pythia) and detector resolution

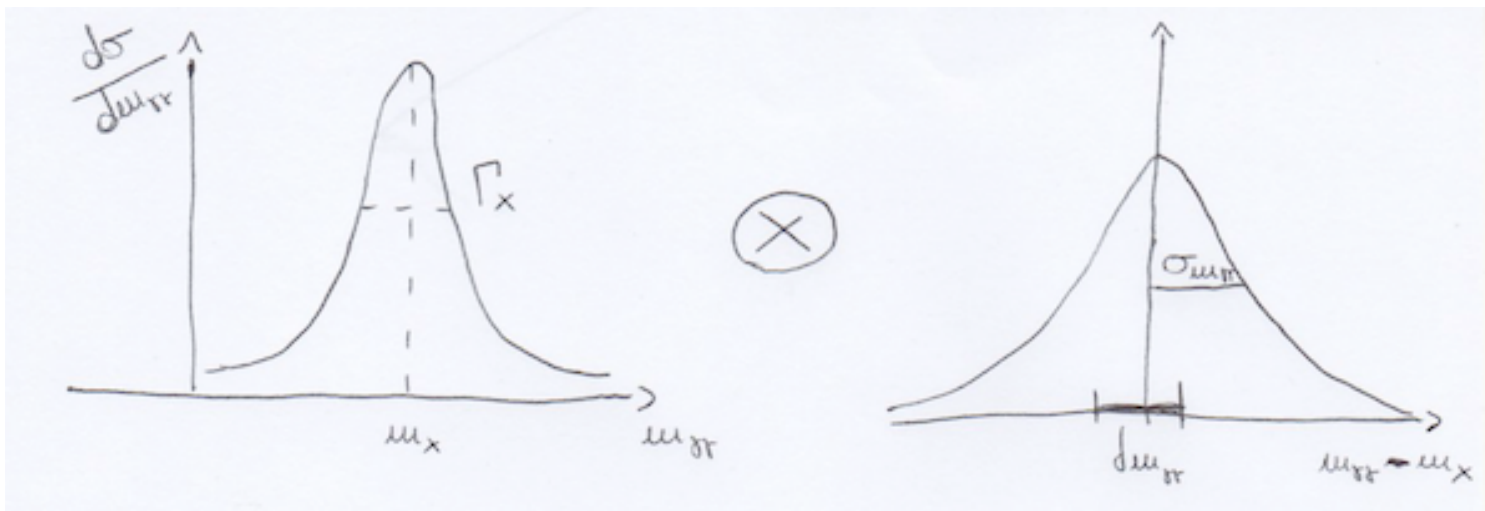
- $\sigma_m/m \sim 1\%$ (EBEB) , 1.5% (EBEE)
- Spin-0 and spin-2
- m_x : 0.5 - 4.5 TeV
- $\Gamma/m = 1.4 \times 10^{-4}$, 1.4×10^{-2} , 5.6×10^{-2} (k=0.01, 0.1, 0.2 for RS Graviton)



Detector resolution dominates

Comparable resolution and width

Resonance width dominates

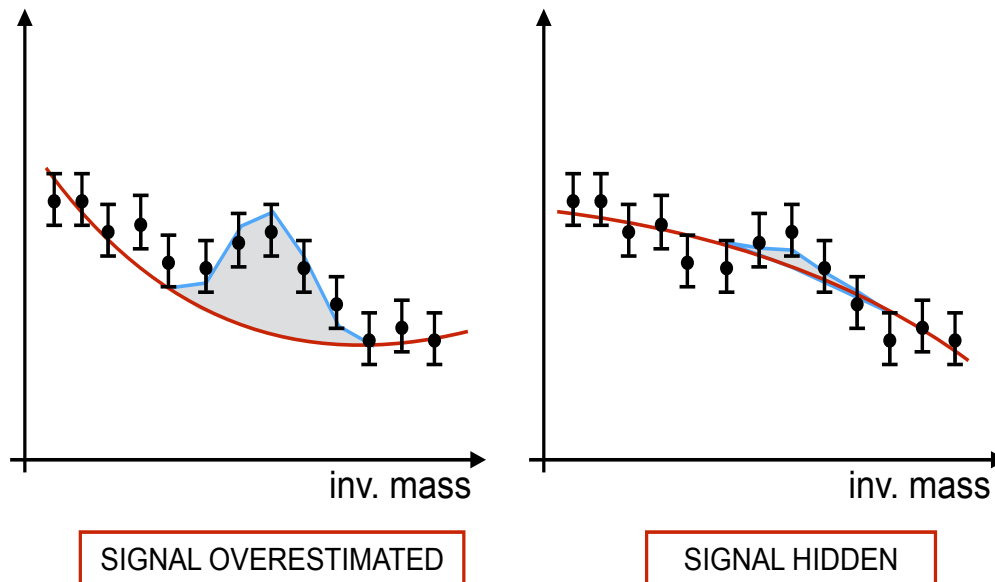


Background modeling

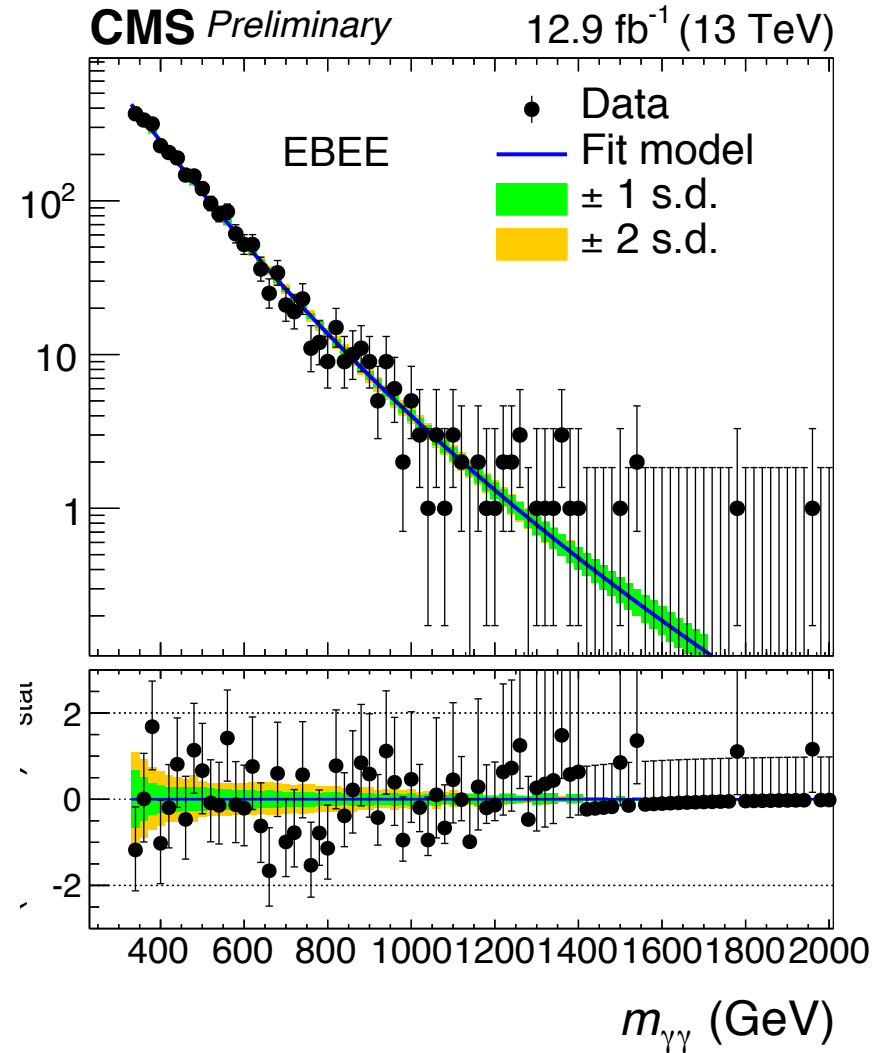
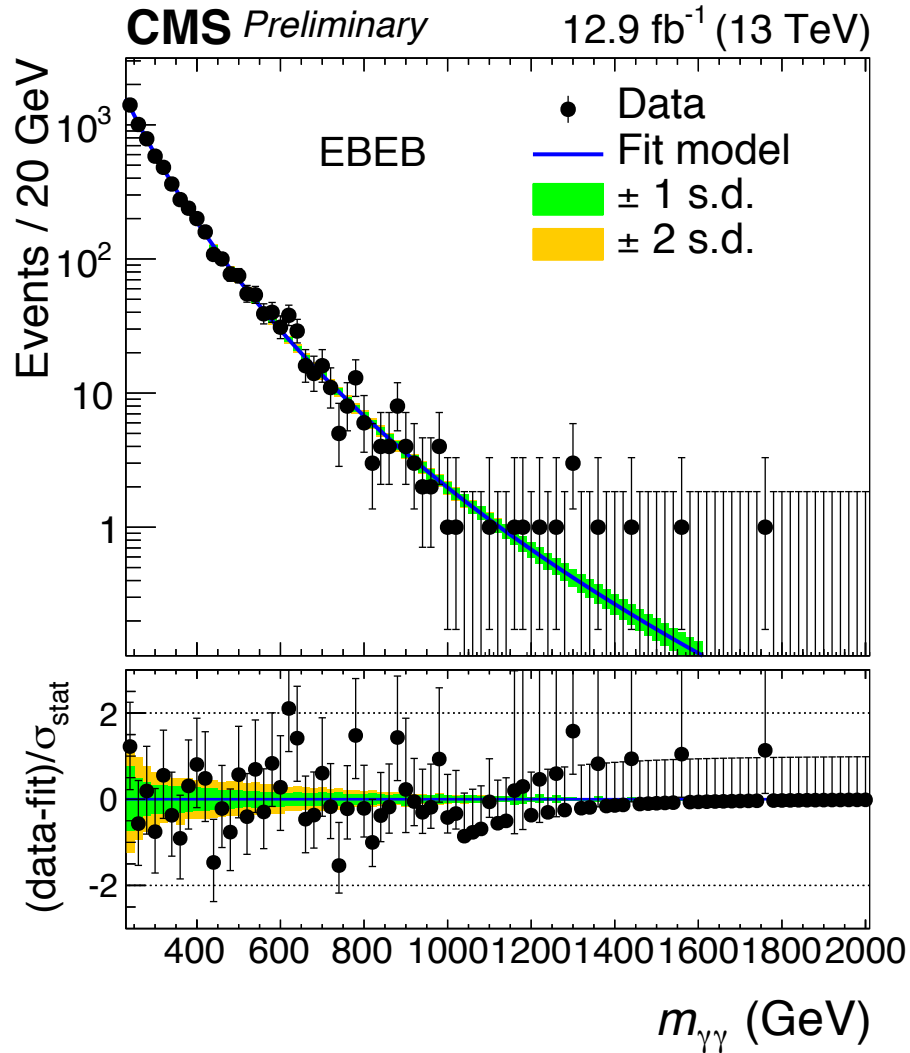
Statistical interpretation from simultaneous fit to $m_{\gamma\gamma}$ distribution in the two analysis categories

Background model

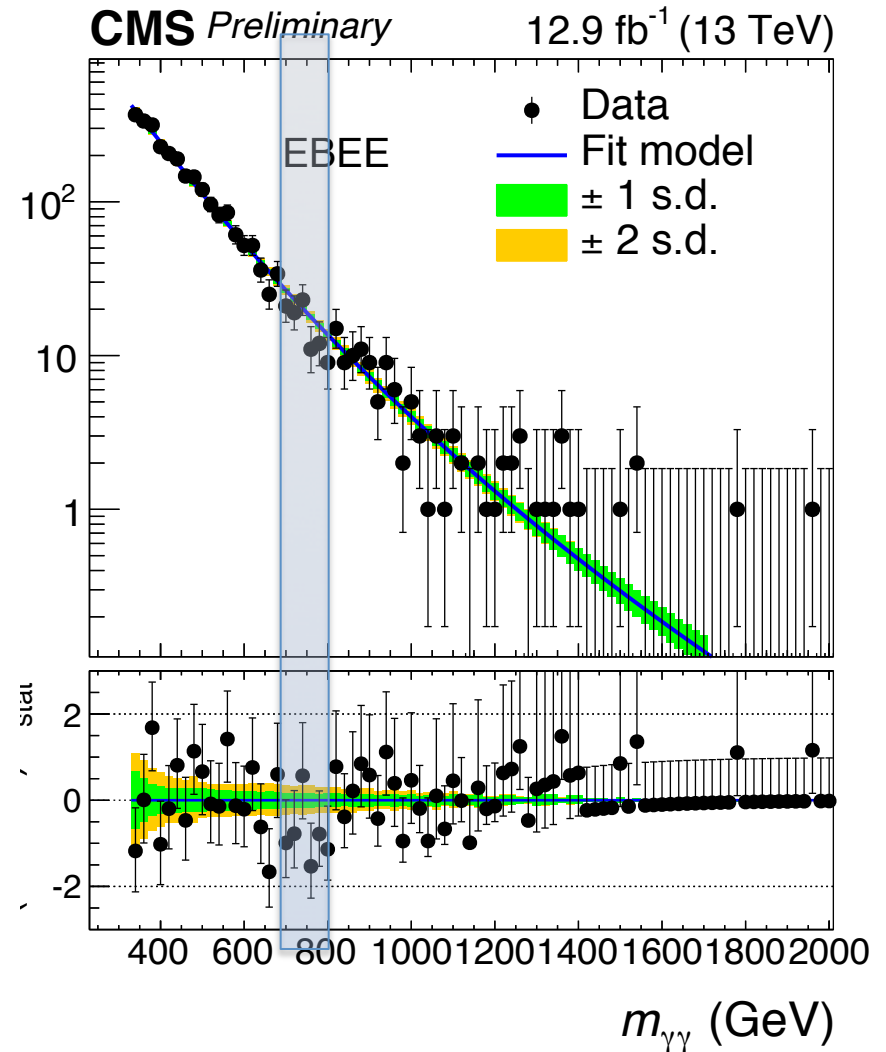
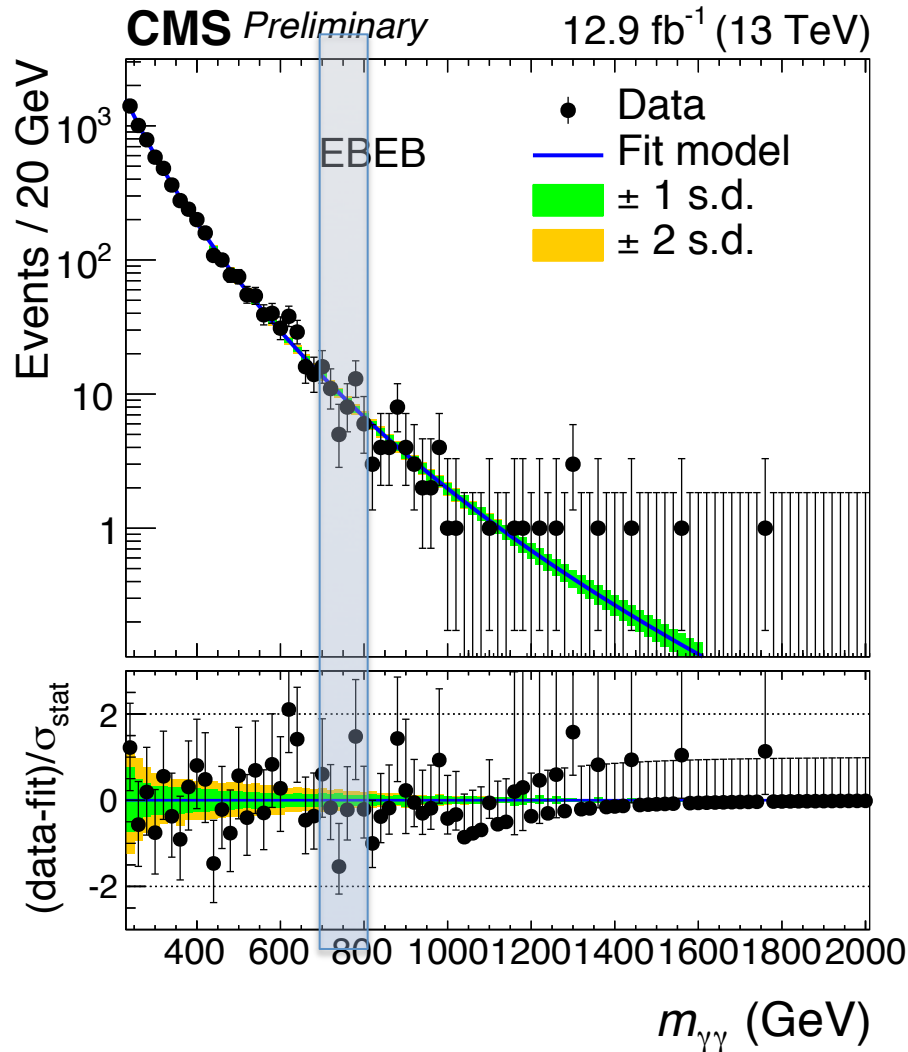
- Parametric fit to data with empirical function $f(m_{\gamma\gamma}) = m_{\gamma\gamma}^{a+b \cdot \log(m_{\gamma\gamma})}$
- Independent shape for each category
- Model coefficients: nuisance parameters in the hypothesis test
- Possible mis-modeling studied on MC and included as a “bias term”



2016 mass spectra

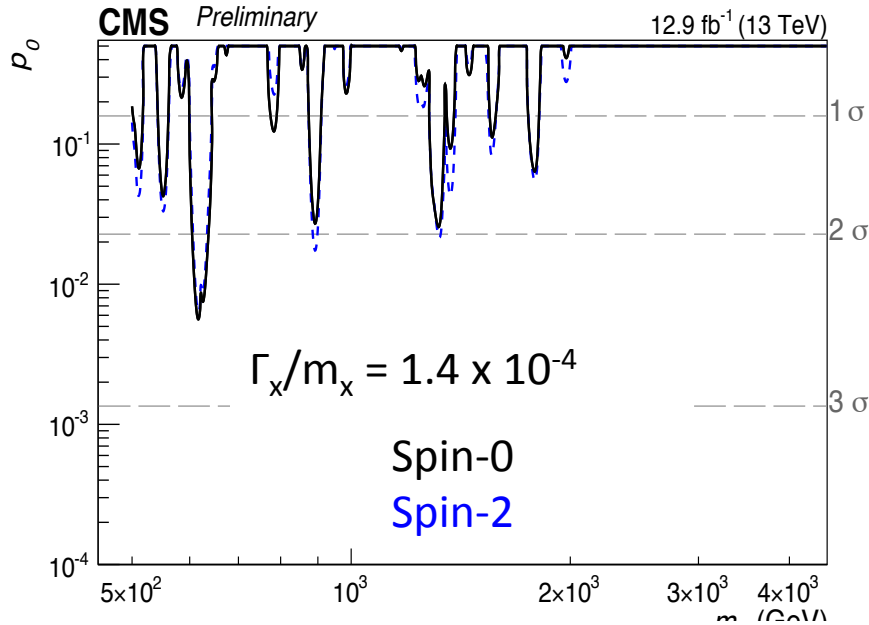


2016 mass spectra



Data consistent with Standard Model expectations

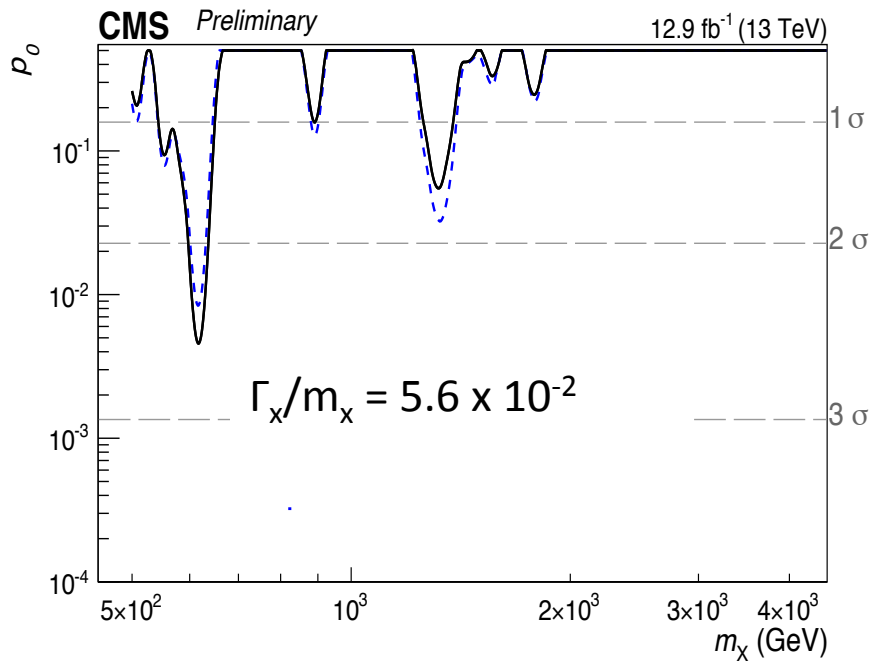
2016 observed p-value



No significant excess in proximity of 750 GeV

Largest excess now observed for $m_\chi \sim 620$ GeV

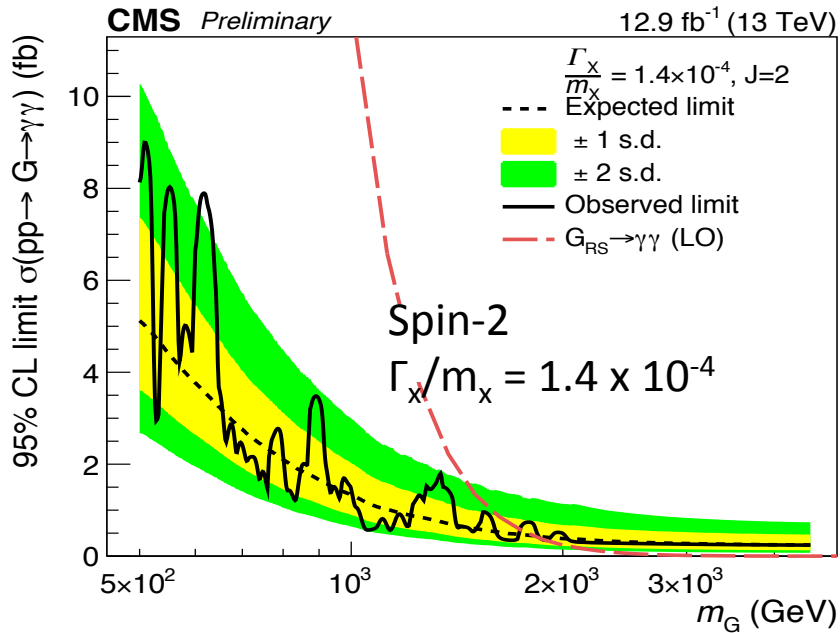
Local significance (narrow width): ~ 2.4 - 2.7σ



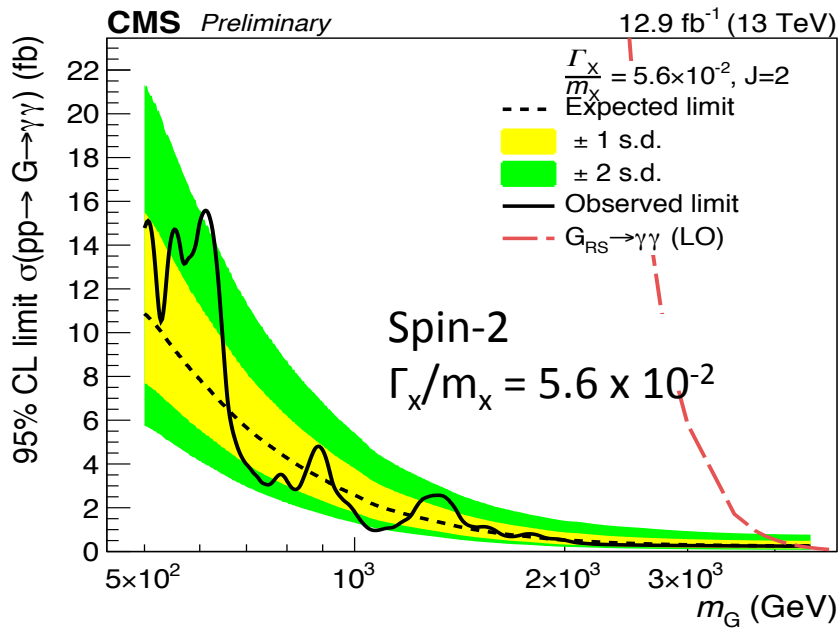
Showing only plots for selected width and spin hypothesis.

All other plots available in backup

2016 exclusion limits



Coupling	Exclusion
0.01	$m_G < 1.75$ TeV
0.1	$m_G < 3.75$ TeV
0.2	$m_G < 4.35$ TeV



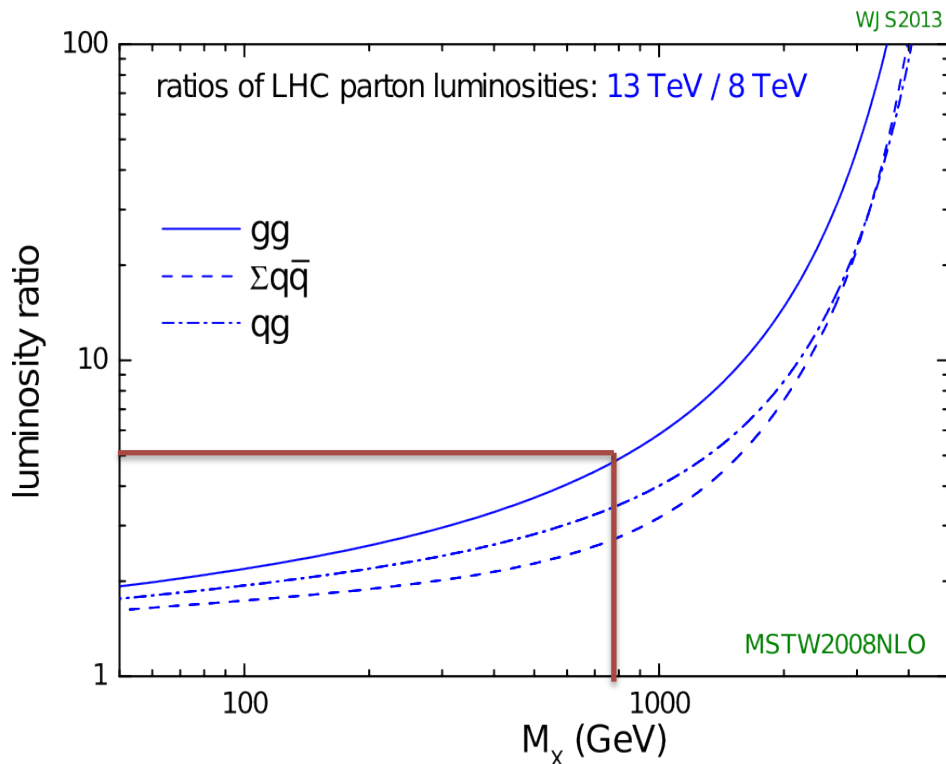
- Exclusion limits for RS Graviton
 (LO cross-sections)
- in agreement with expectations

8TeV + 13TeV

Combination of results from

- 19.7/fb at $\sqrt{s} = 8\text{TeV}$
- 3.3/fb at $\sqrt{s} = 13\text{TeV}$ (2015)
- 12.9/fb at $\sqrt{s} = 13\text{TeV}$ (2016)

CMS publications	\sqrt{s} [TeV]
CMS-PAS-EXO-16-027 (new)	8 / 13
PRL 117, 051802 (2016)	8 / 13
PLB750 (2015) 494–519	8

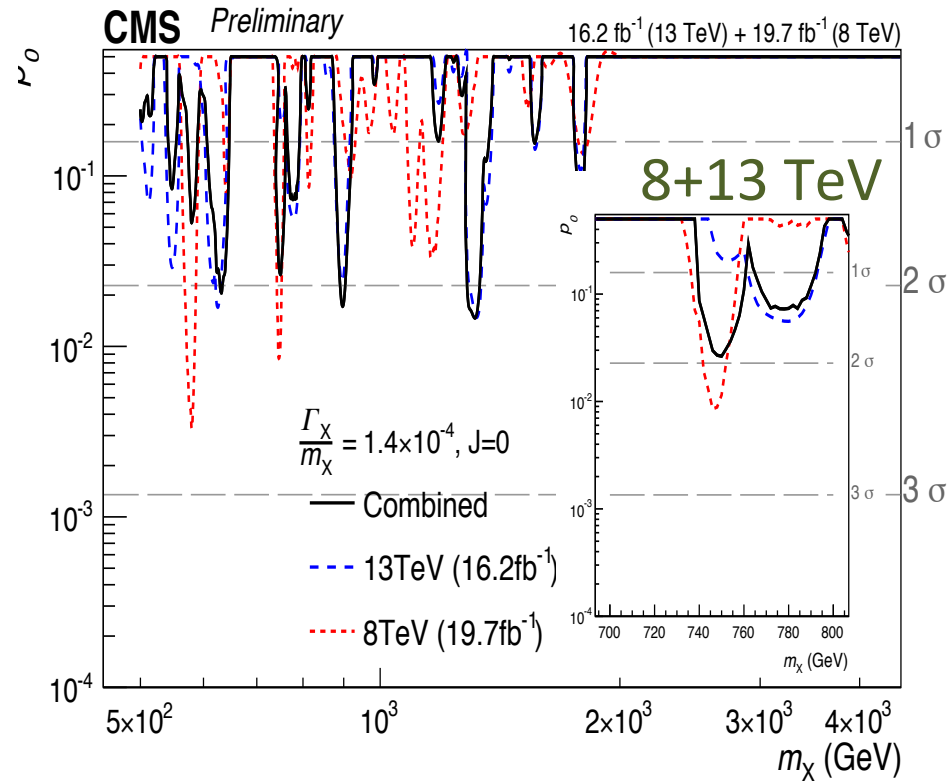
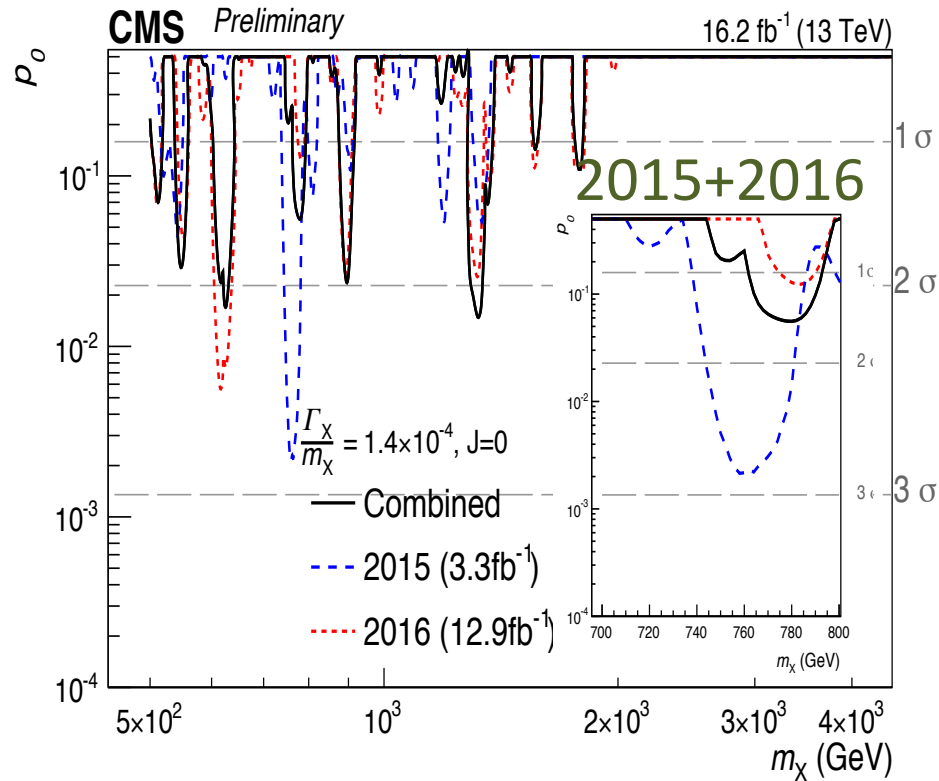


@750GeV:

$$\sigma(13\text{TeV})/\sigma(8\text{TeV}) = 4.5 \text{ spin-0}$$

$$= 4.2 \text{ spin-2}$$

Run1+Run2 significance



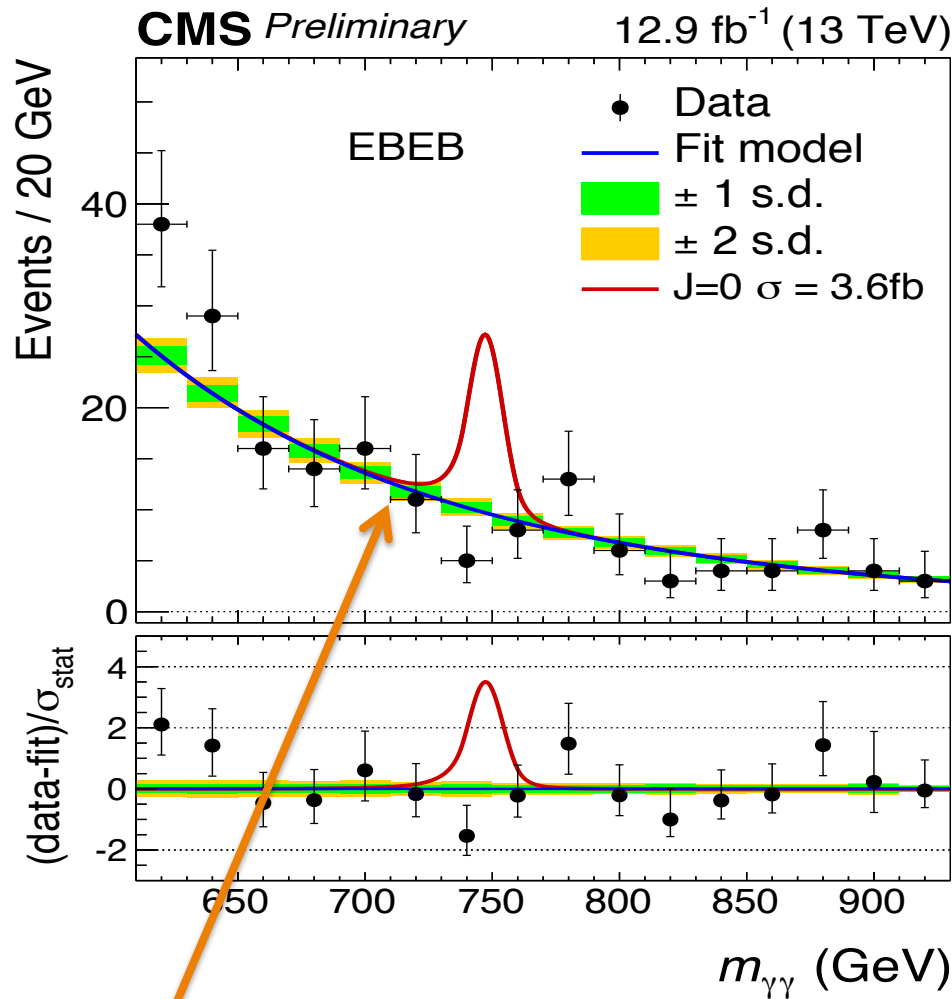
Spin-0, $\Gamma_X/m_X = 1.4 \times 10^{-4}$ hypothesis

Local excesses around 750GeV:

2015 only: 2.9 σ 2015+2016: <1 σ

8TeV+2015: 3.4 σ 8TeV+2015+2016: <2 σ

Compatibility among results

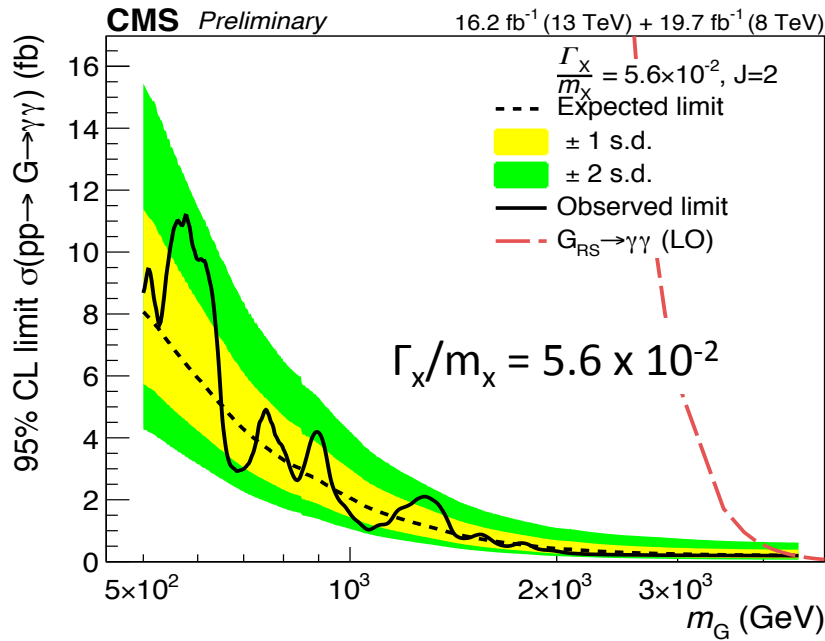
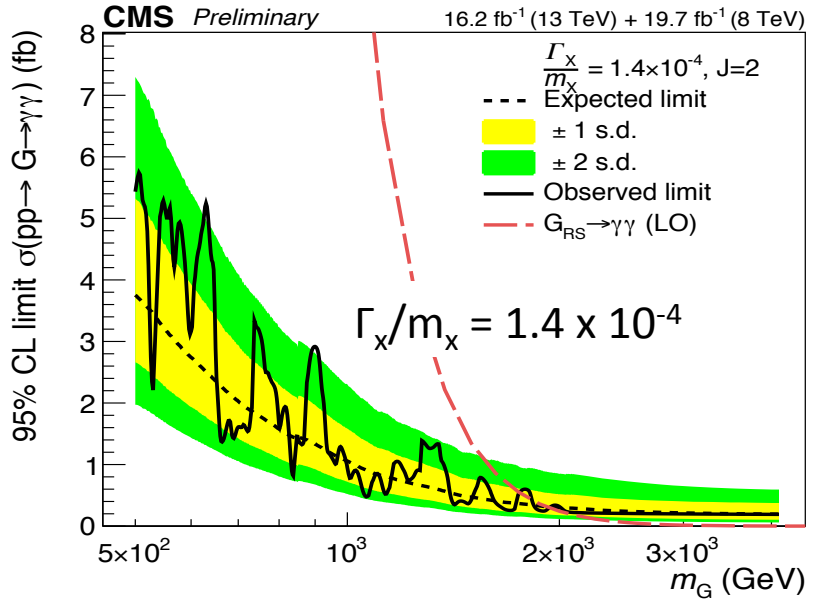


Compatibility of data with a common signal strength at $m_{\chi}=750 \text{ GeV}$ tested with a likelihood ratio test.

- Found to be at the level of
- 13TeV data only: 2.7σ
 - 8TeV+13TeV data: 2.4σ

A signal with cross-section as the largest excess in 2015+8TeV would be like this

Run1+Run2 limits



Sensitivity driven by 2016 dataset
13TeV

Coupling	Exclusion
0.01	$m_G < 1.95 \text{ TeV}$ except for [1.75, 1.85]
0.1	$m_G < 3.85 \text{ TeV}$
0.2	$m_G < 4.45 \text{ TeV}$

8TeV contribution:

~10% at low mass, negligible at high mass

Conclusions

Search for new resonances decaying to di-photon pairs presented, based on 12.9/fb of 13TeV CMS 2016 data

- Mass region between 0.5 and 4.5 TeV
- Tested hypothesis: spin-0 and spin-2 resonances with different widths

Data consistent with Standard Model expectations

Modest excess presented based on 2015 (+ 8TeV) data in the region around 750 GeV not confirmed by the new data

- Compatibility of data with a common signal strength at level of 2.4σ

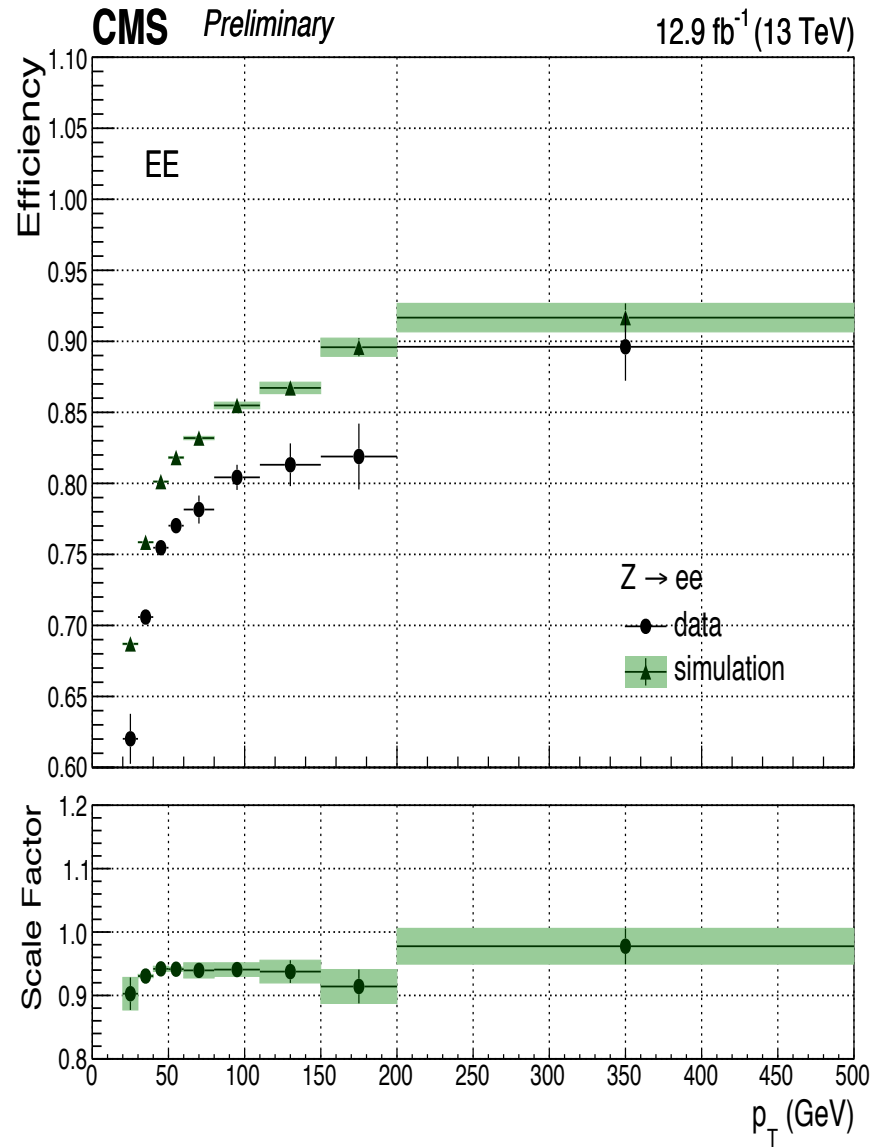
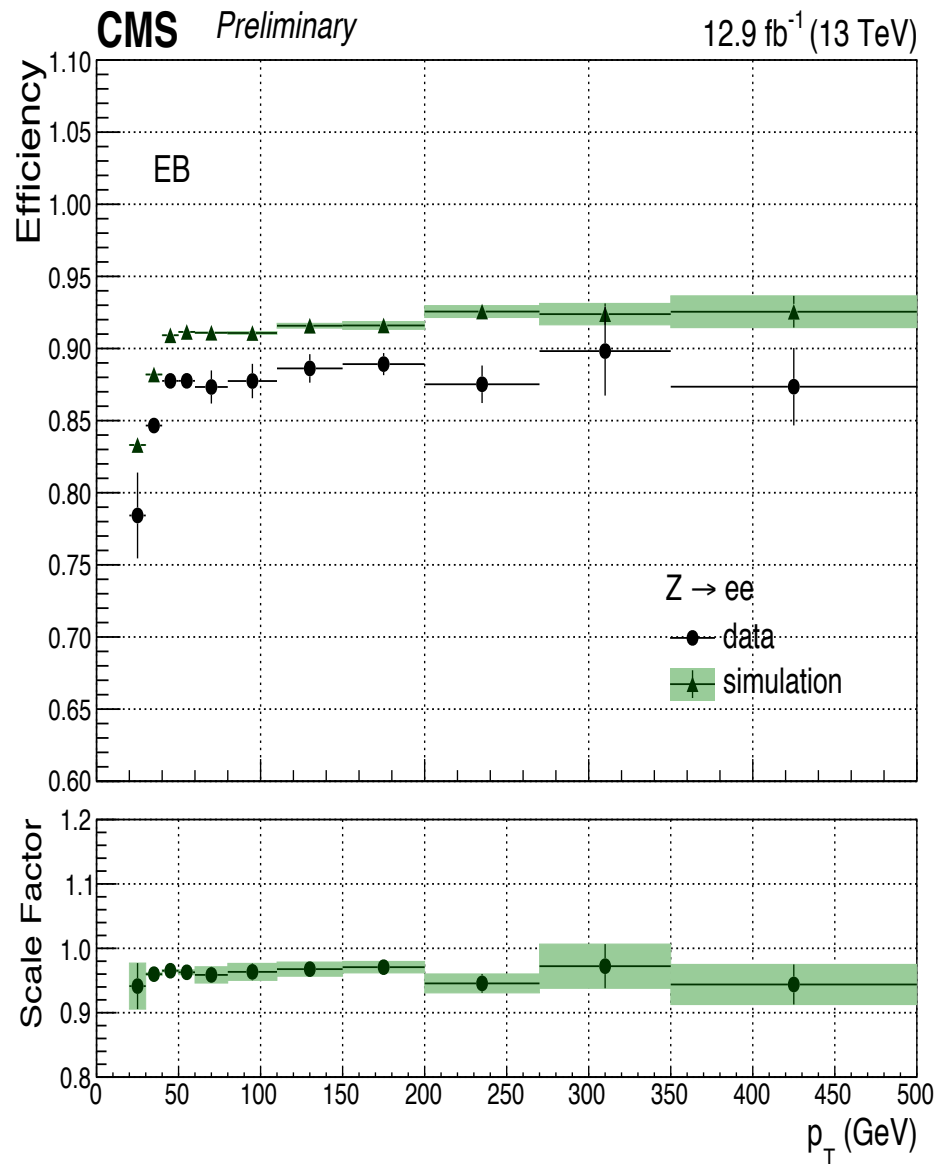
2016 data combined with 2015 and 8TeV data

Limits set on the production cross section times di-photon branching ratio

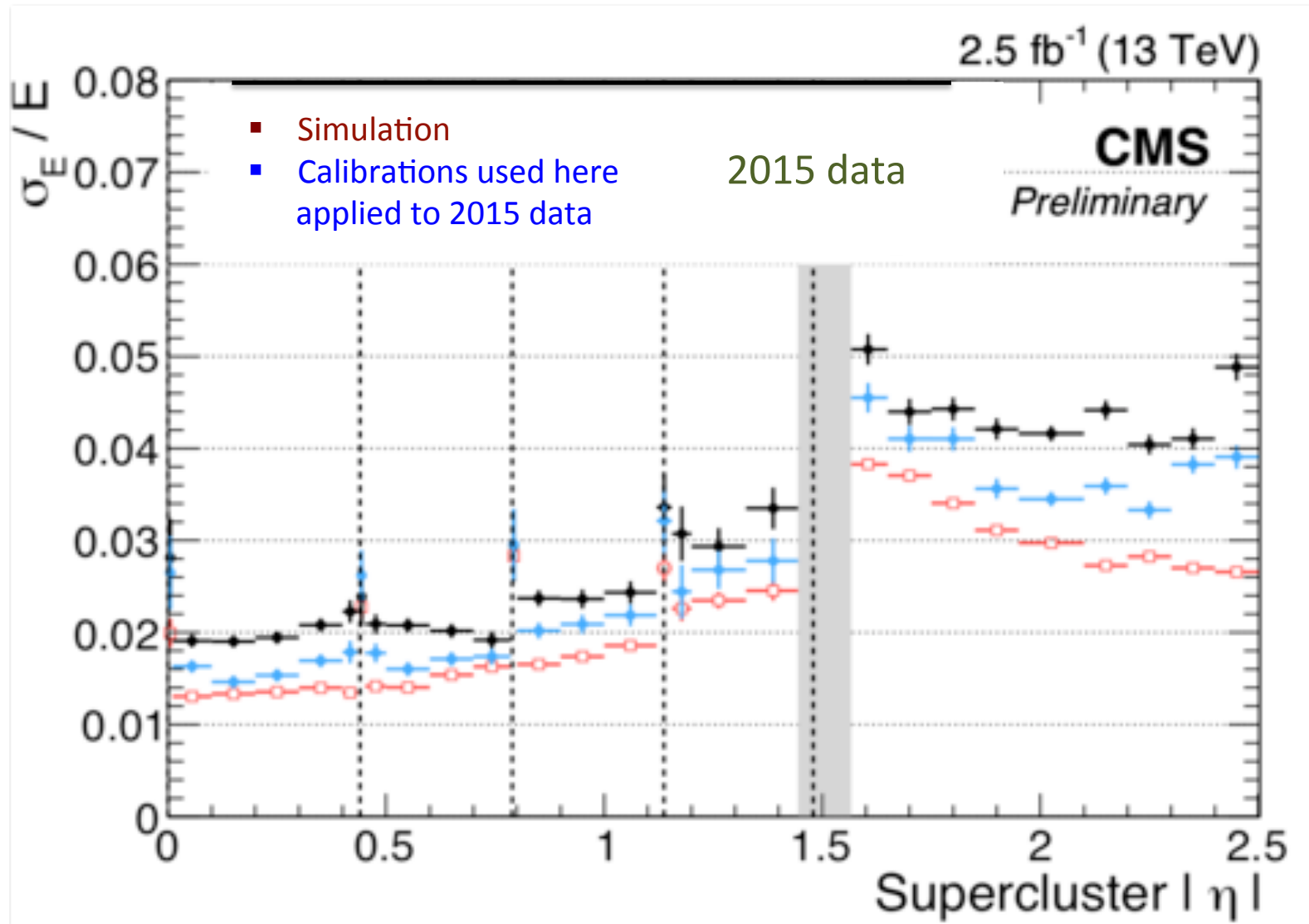
- Negligible contribution of the 8TeV dataset
- 2016 data dominating limits and significances

Backup

Photon selection efficiencies

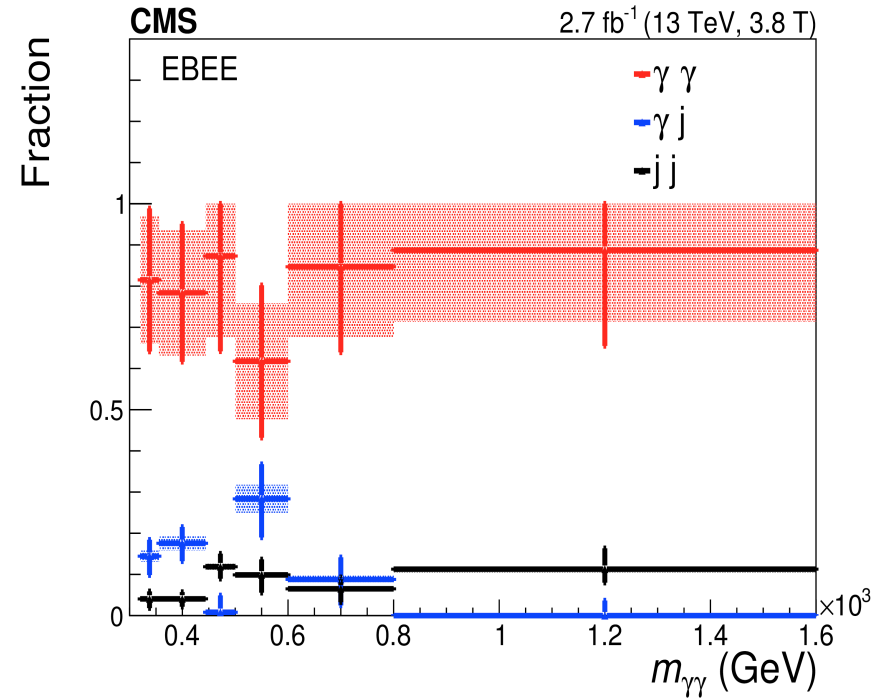
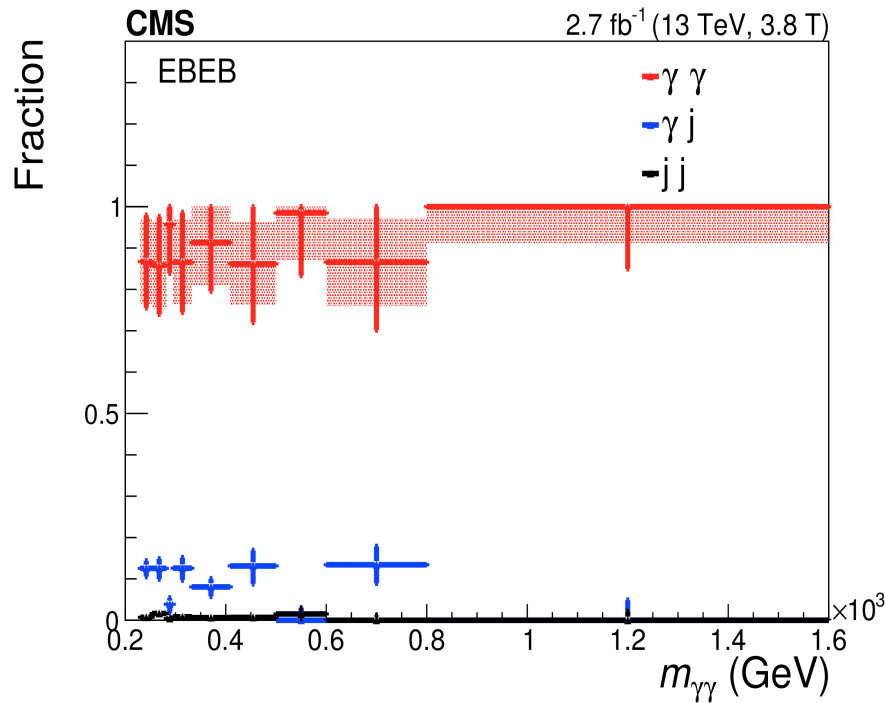


Energy calibration



Background purities

2015 plots. Consistent purities measured this year



Dominant contribution: **2 prompt photons**

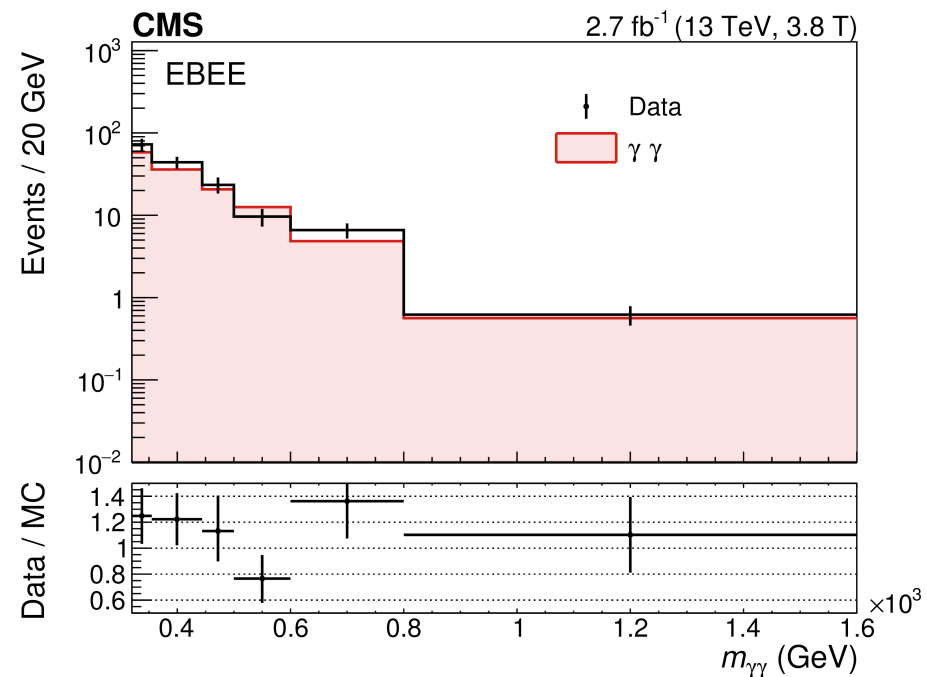
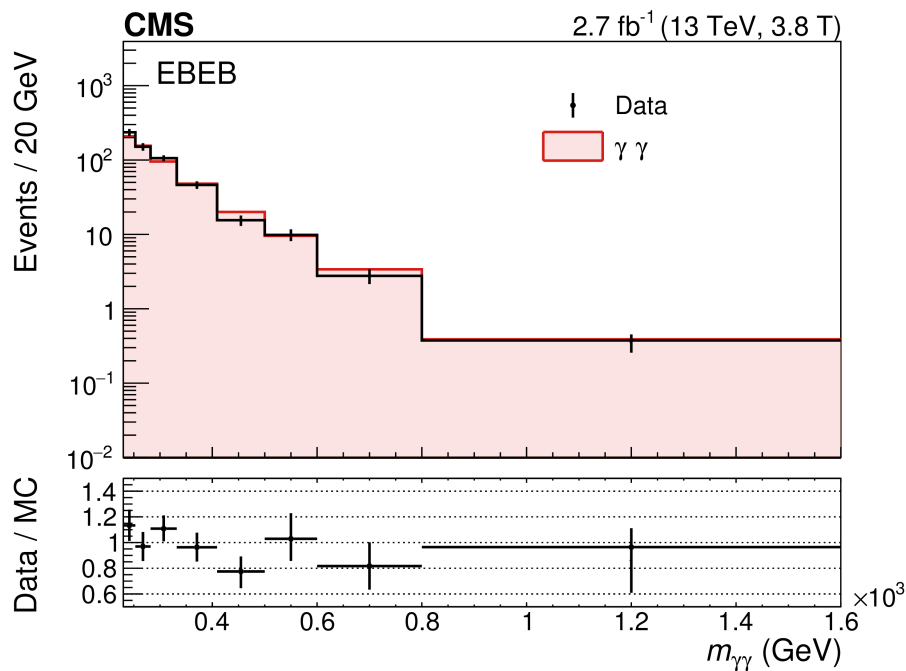
QCD and **photon+jets**: <10% (20%) in EBEB (EBEE)

Background composition, closure test

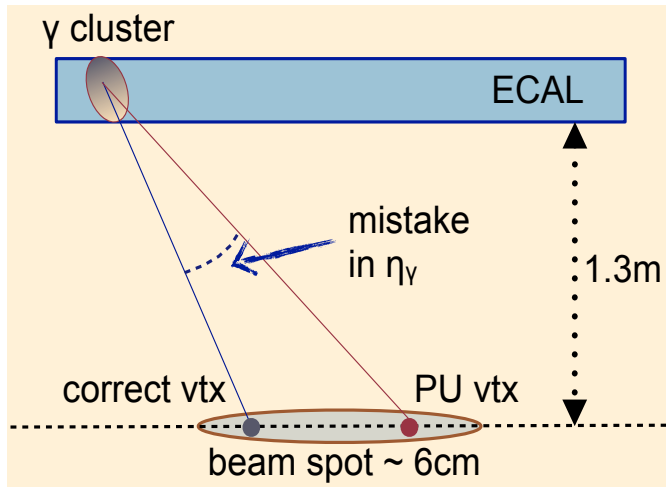
Data driven prediction for the prompt-prompt component compared with theory

✓ Sherpa generator rescaled to $2\gamma\text{NNLO}$

2015 plots. Consistent purities measured this year



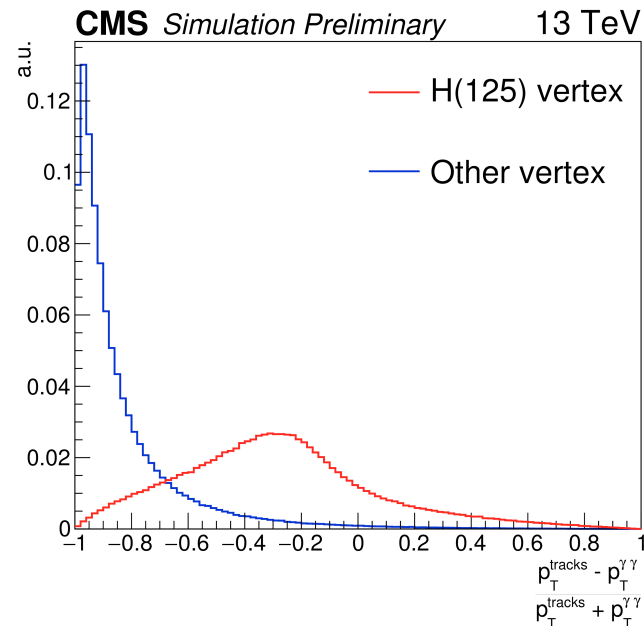
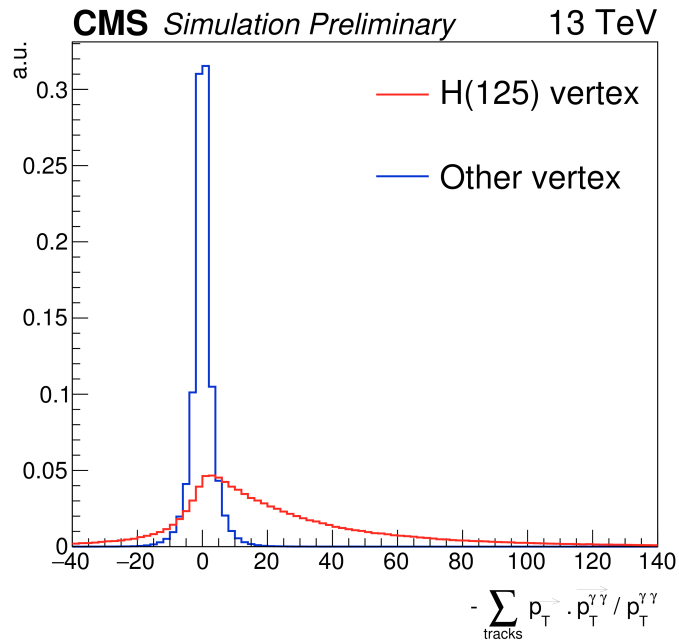
Vertex determination



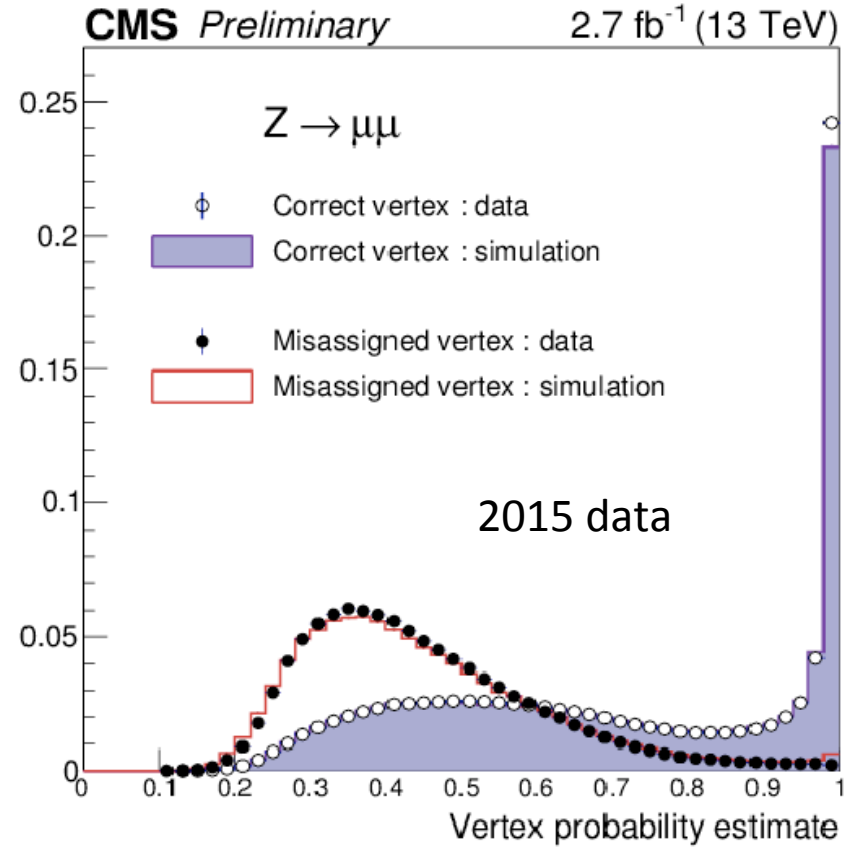
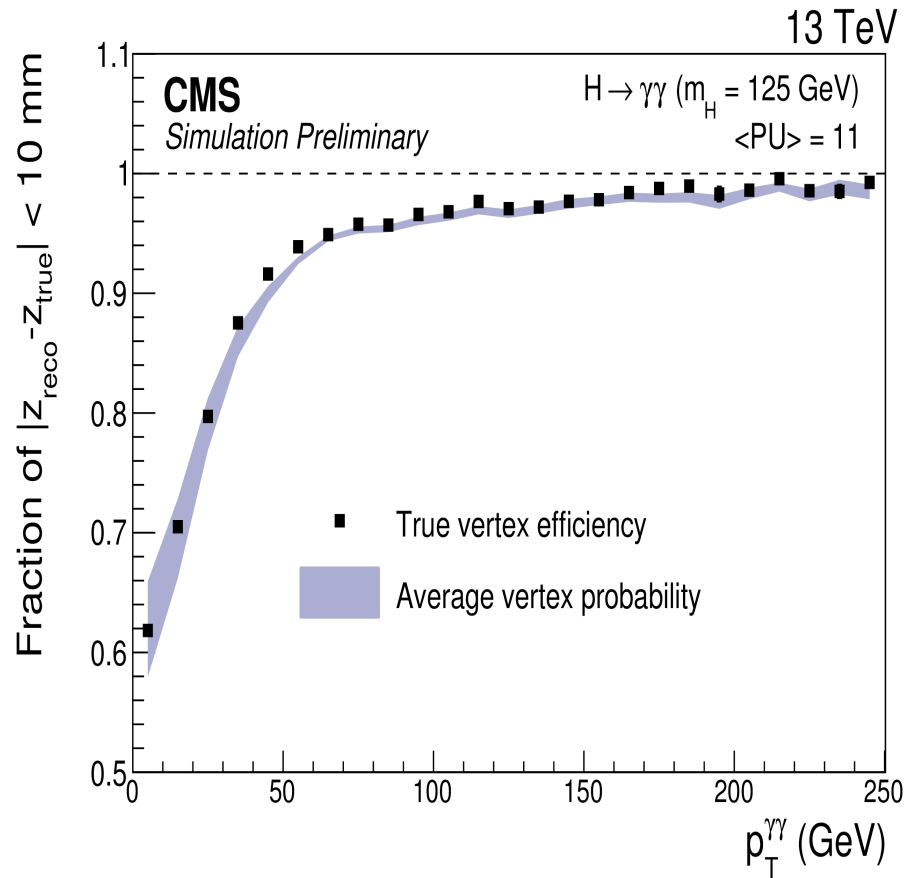
Interaction vertex identified using
recoiling tracks and conversions when present:
 $\text{Sum}(p_T^2)$, $p_T(\gamma\gamma)$ vs $p_T(\text{tracks})$, $Z(\text{conv})$

Combined in a BDT

Same as in $H \rightarrow \gamma\gamma$ analysis



Vertex ID performance



Probability to assign the correct vertex in $H \rightarrow \gamma\gamma$ analysis

Background model

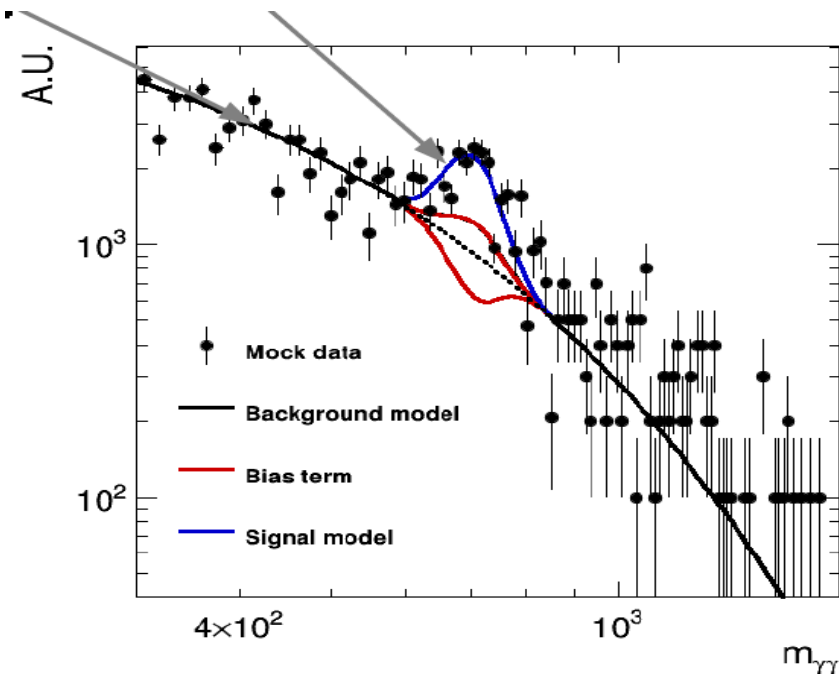
Parametric 1dim fit to data in the 2 categories, $f(m) = m^{a+b \log(m)}$

- Model coefficients treated as unconstrained nuisance parameters in the hypothesis test

Goodness of background fit assessed locally in $m_{\gamma\gamma}$ using MC

- Study pull of mean number of background events
- Model ok if $b = |\text{median}(p)| < 0.5$ for all windows
 - Uncertainty on mean number of B events underestimated by $< 10\%$
- If not => error increased with a bias term

$$p_i^j = \frac{N_{\hat{g}_i}^{w_j} - N_h^{w_j}}{\sigma(N_{\hat{g}_i}^{w_j})}$$



$$\tilde{p}_j^i = \frac{N_{\hat{g}_i}^{w_j} - N_h^{w_j}}{\sqrt{\sigma^2(N_{\hat{g}_i}^{w_j}) + \beta_I^2(w_j)}}$$

Bias term = signal like component added to the model

- negligible impact on sensitivity

Systematics

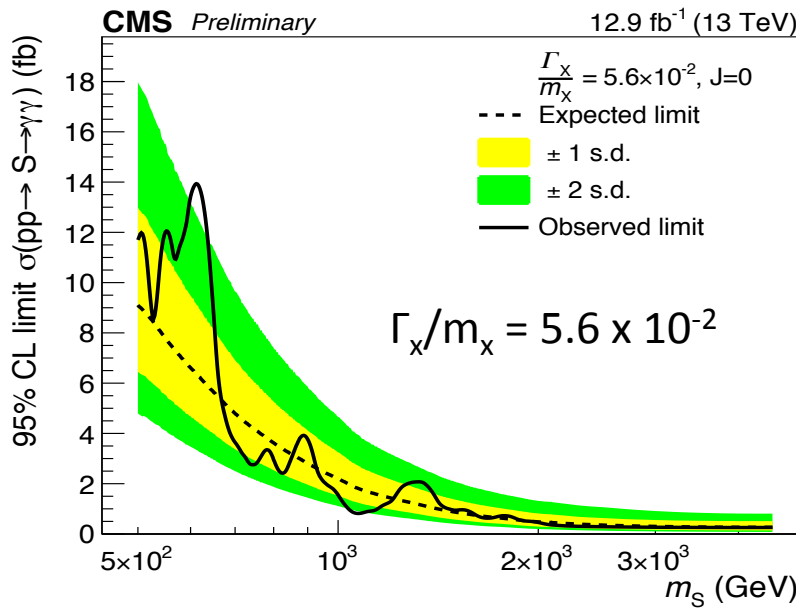
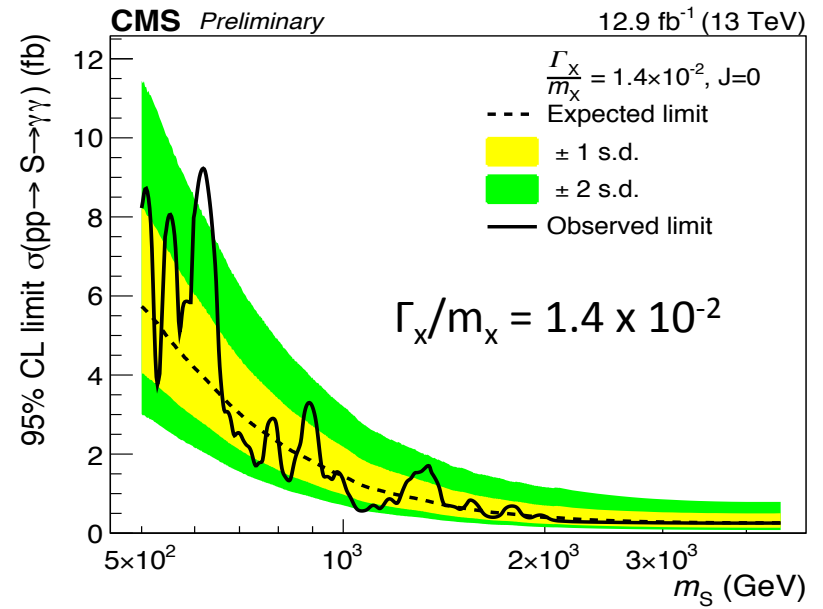
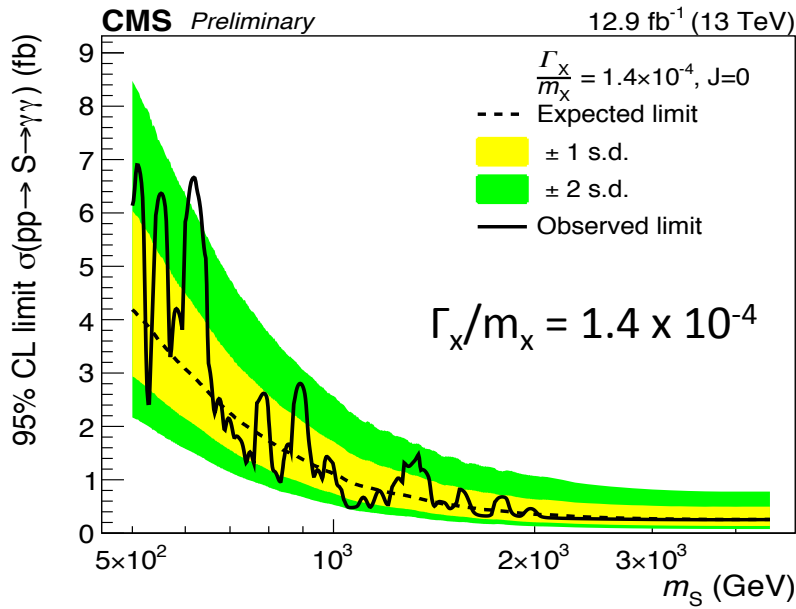
Signal model:

- Luminosity: 6.2%
- Trigger and photon selection: 6%
- Photon energy scale: 1%
- Photon energy resolution: 0.5%
- PDF: 6%

Background model:

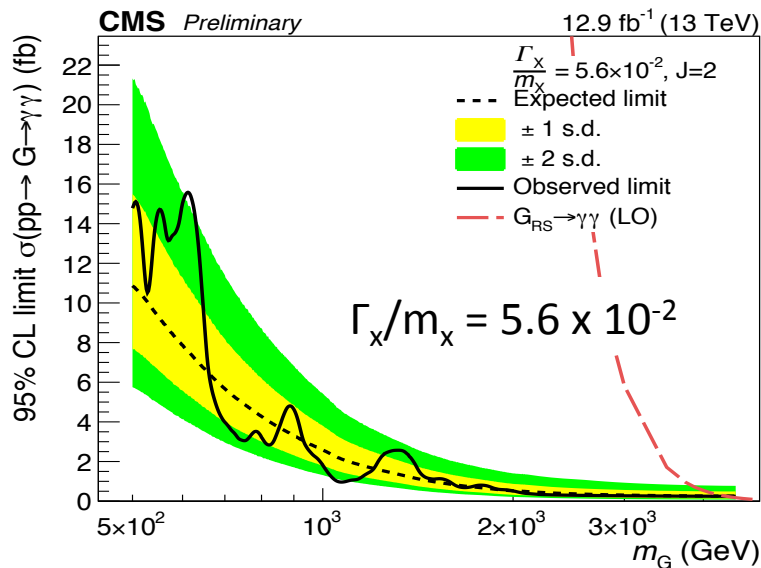
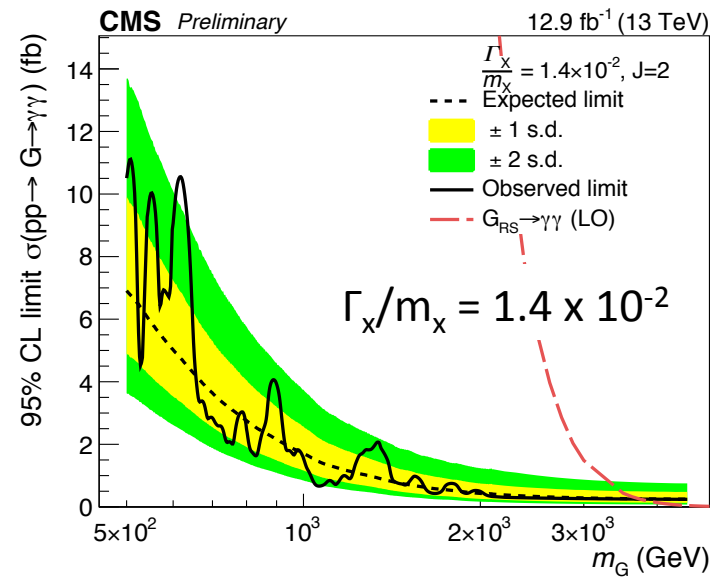
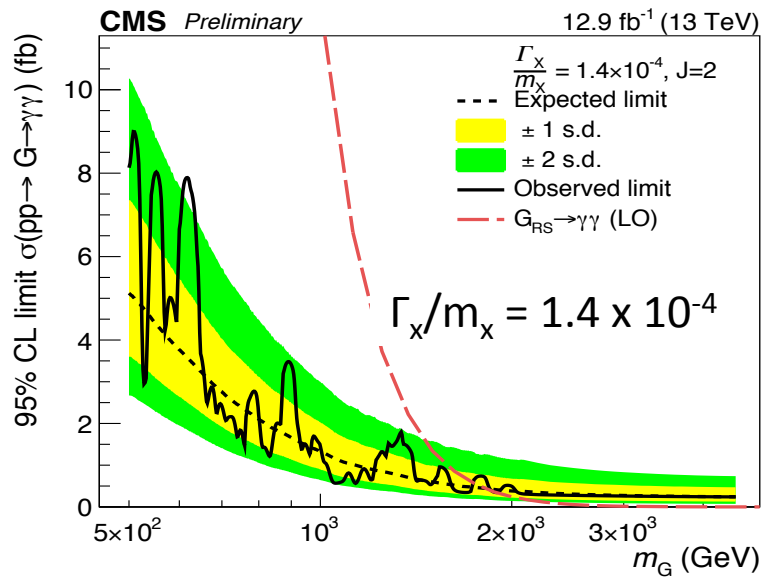
- Bias term only
- [Parameter coefficients: unconstrained nuisance parameters
 - contribute to statistical error]

2016 exclusion limits



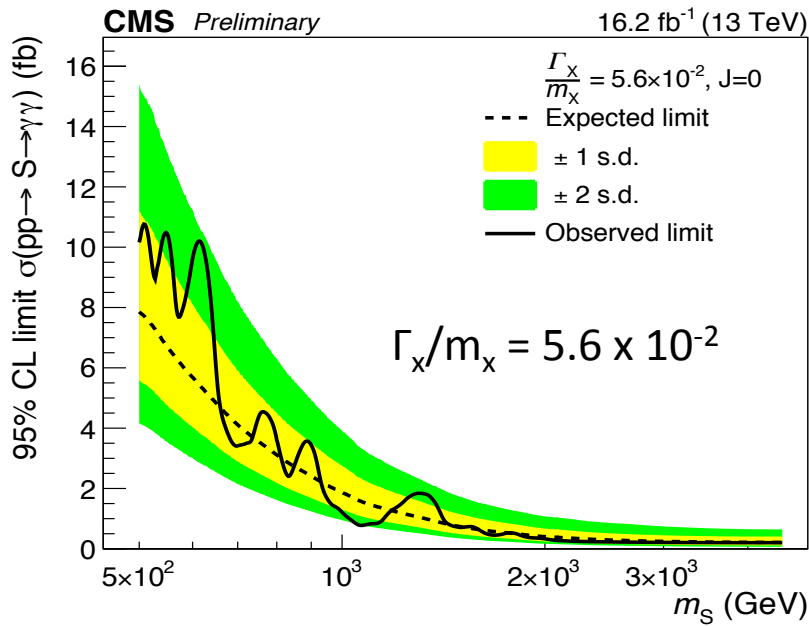
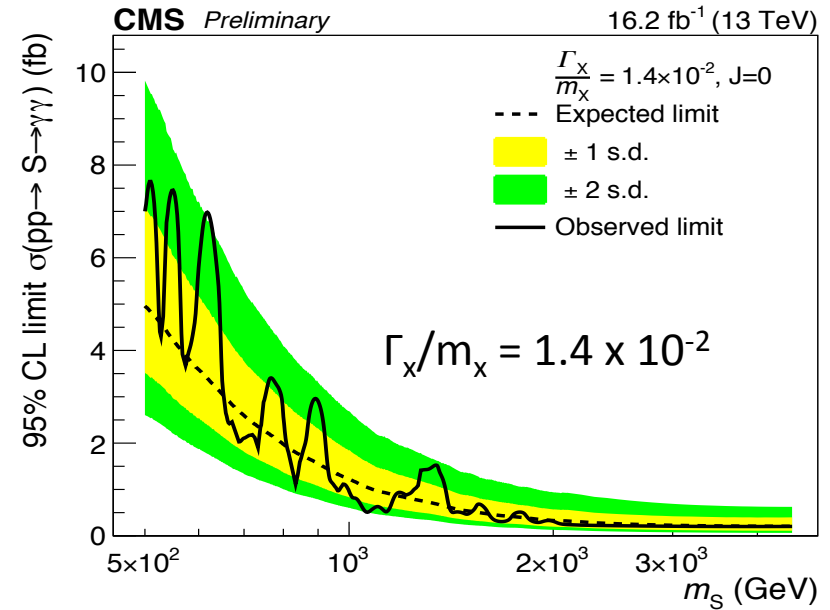
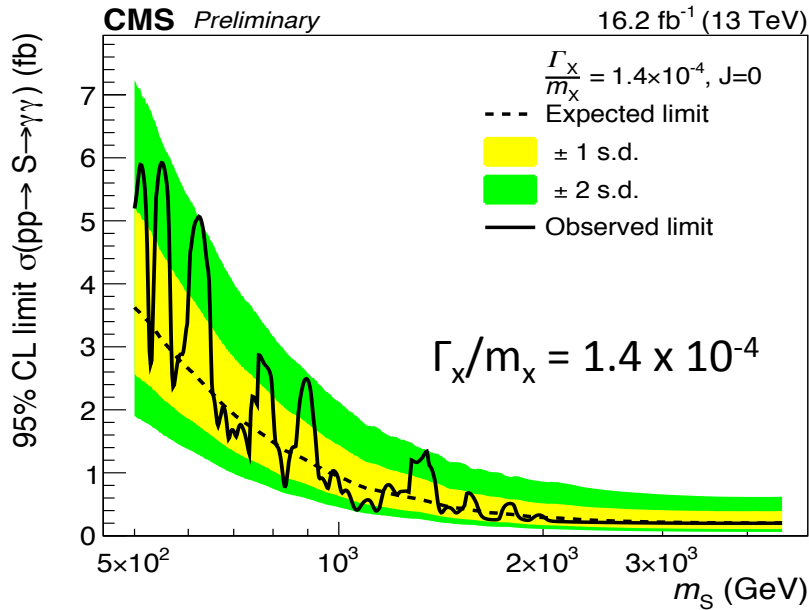
Spin-0

2016 exclusion limits



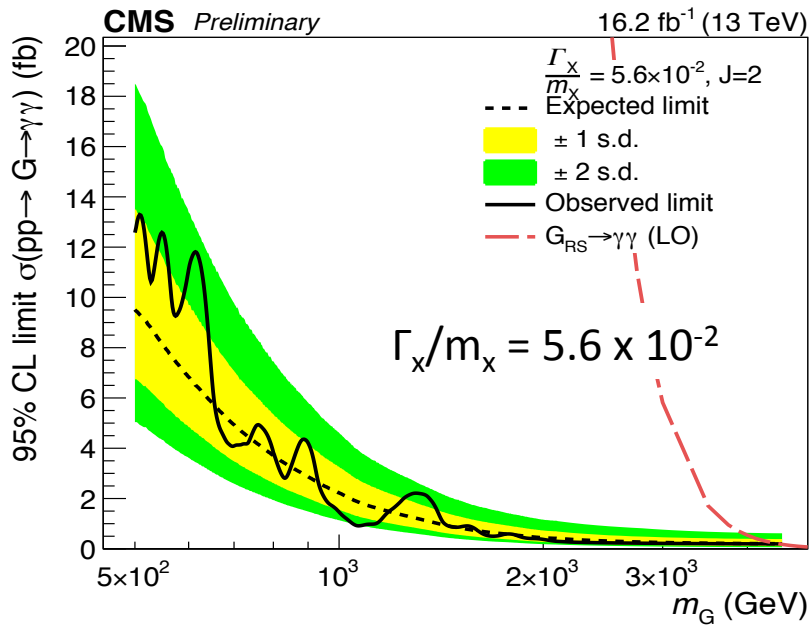
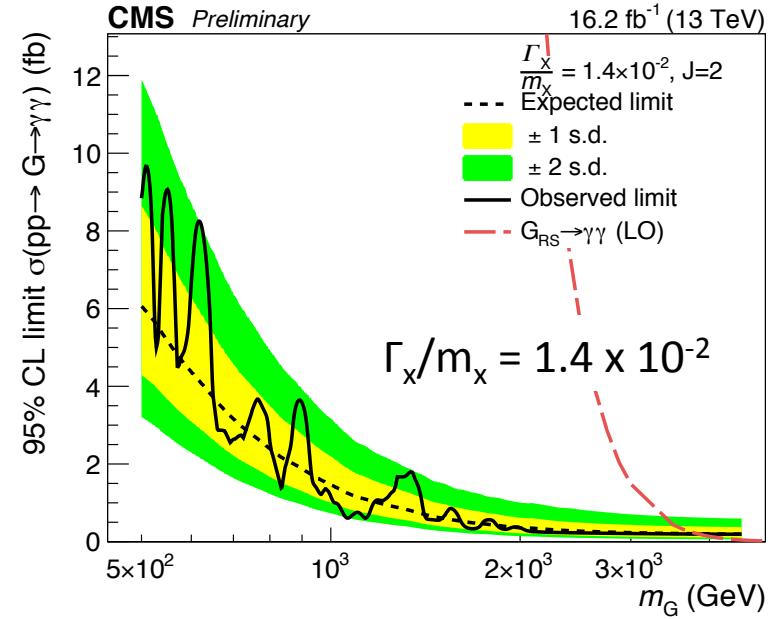
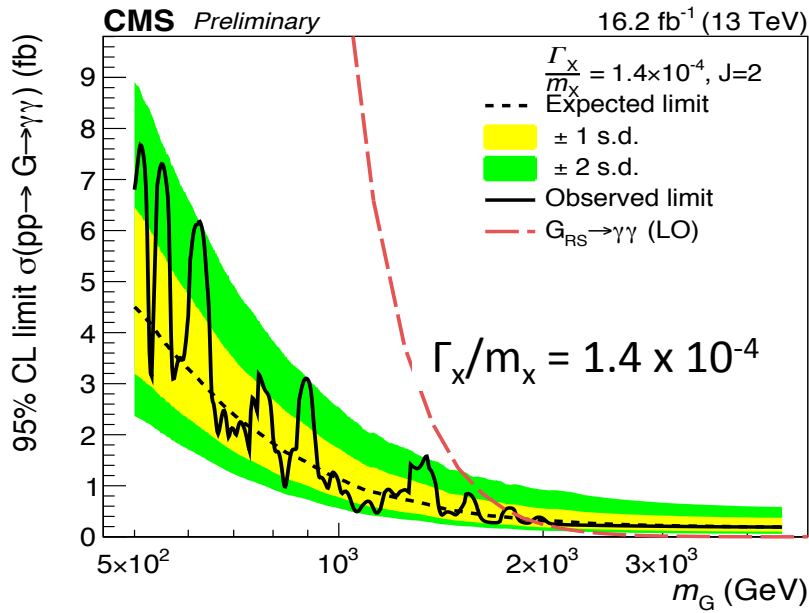
Spin-2

2015 + 2016 exclusion limits



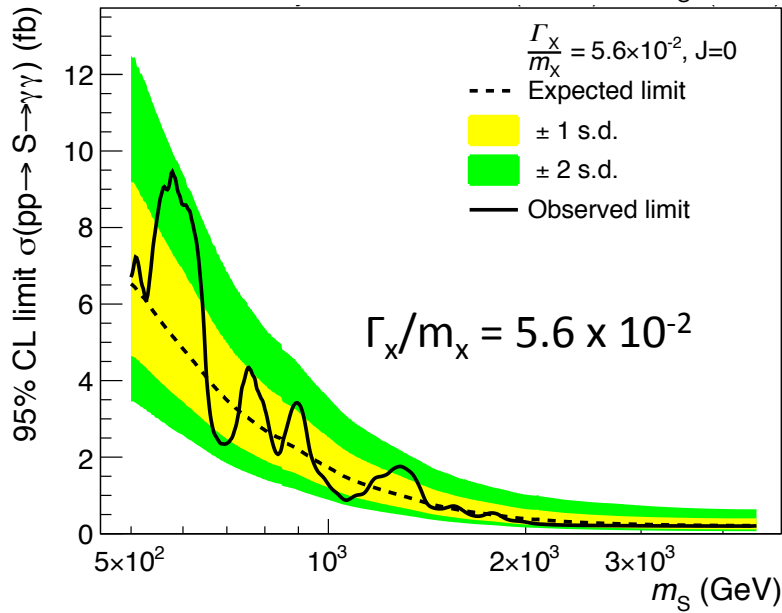
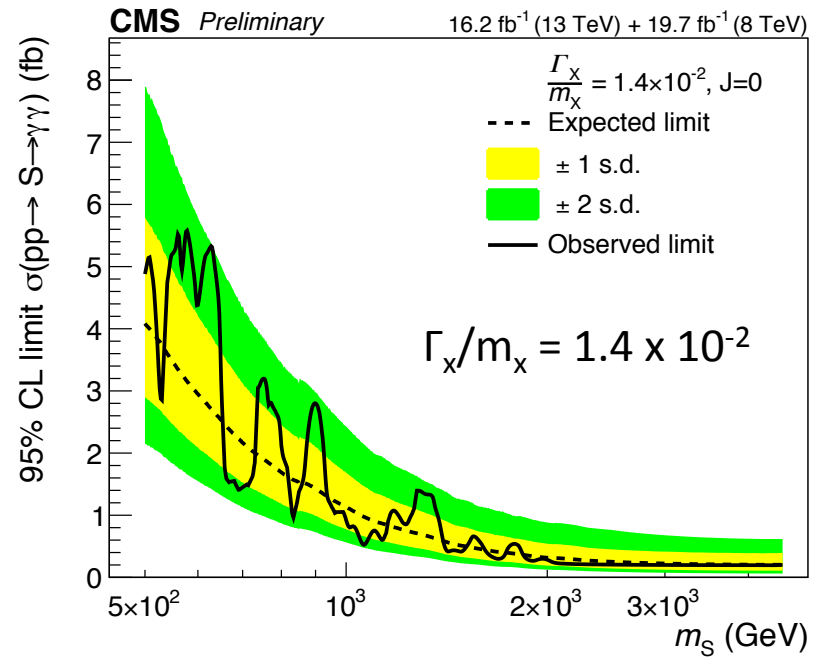
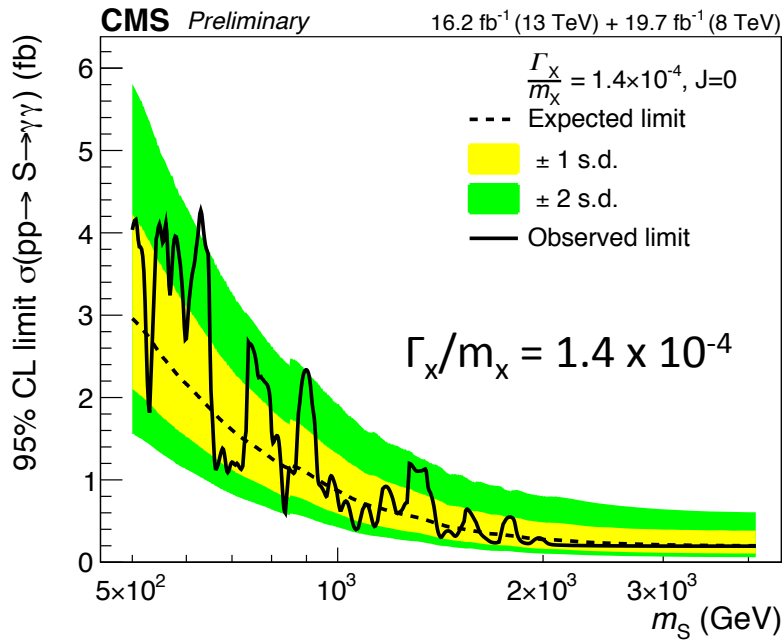
Spin-0

2015 + 2016 exclusion limits



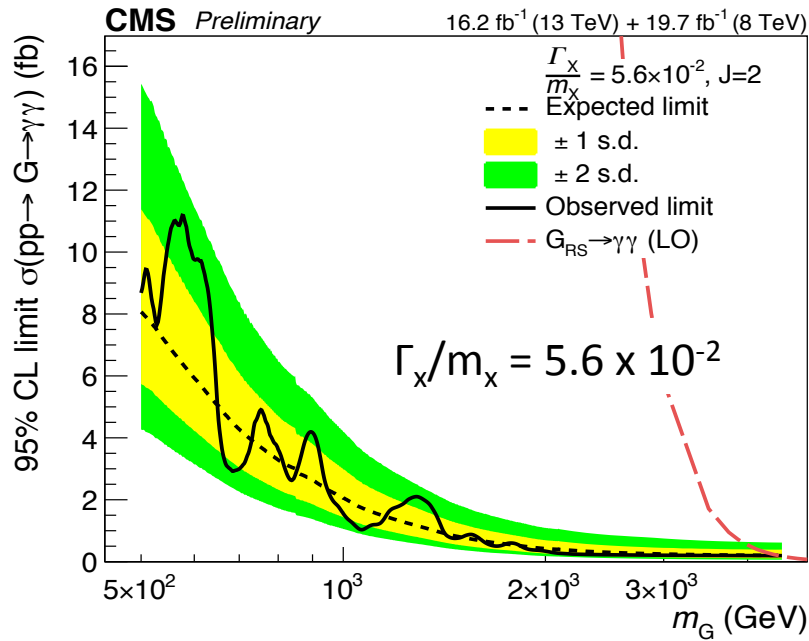
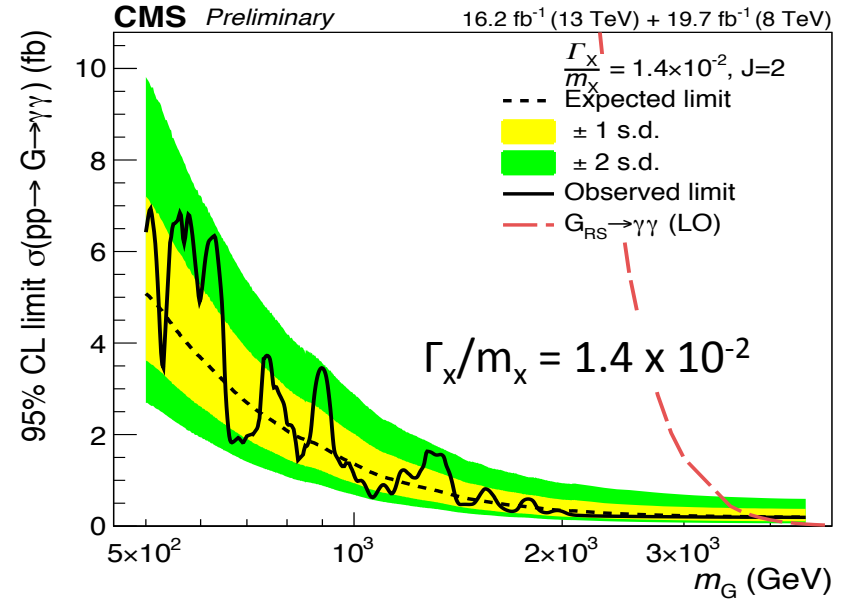
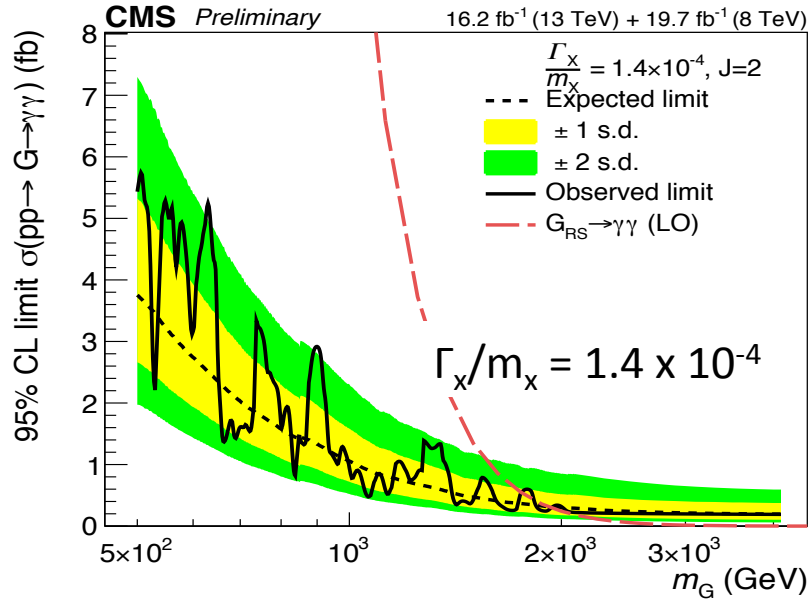
Spin-2

8TeV + 13TeV exclusion limits



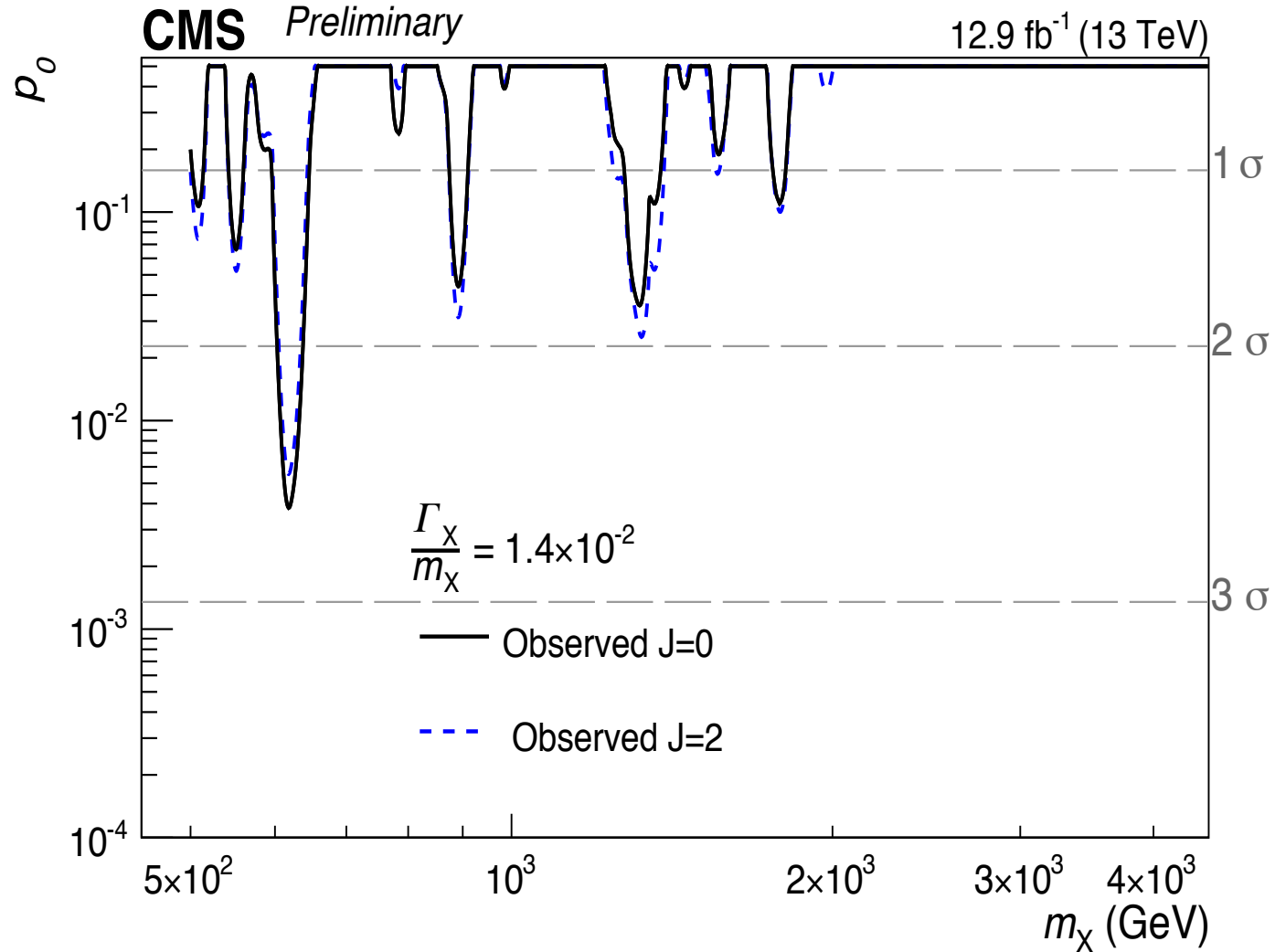
Spin-0

8TeV + 13TeV exclusion limits

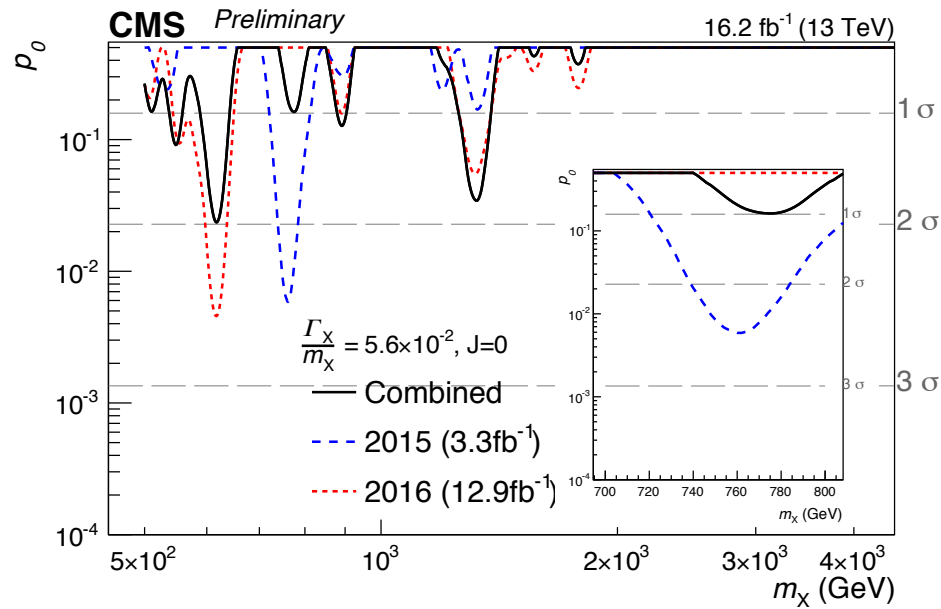
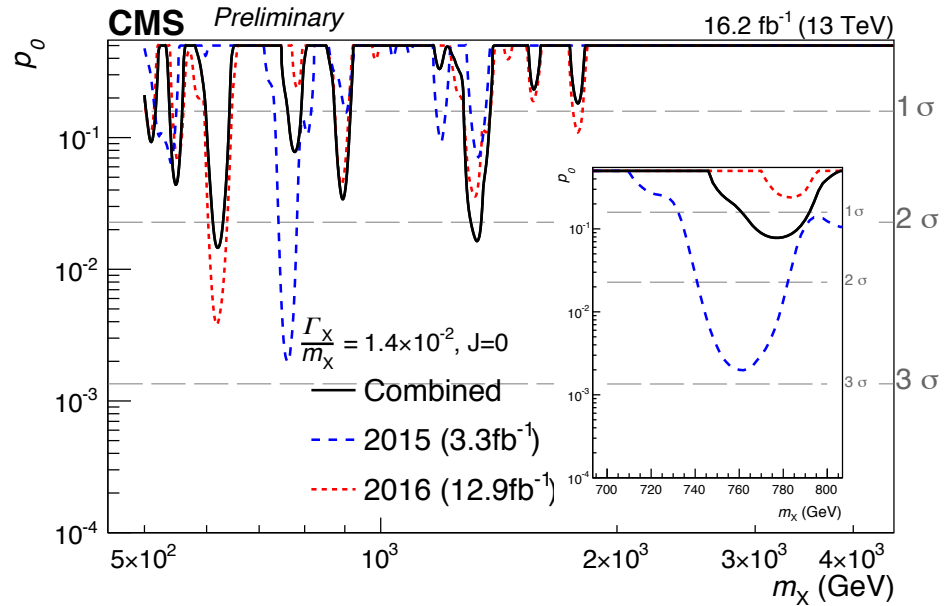
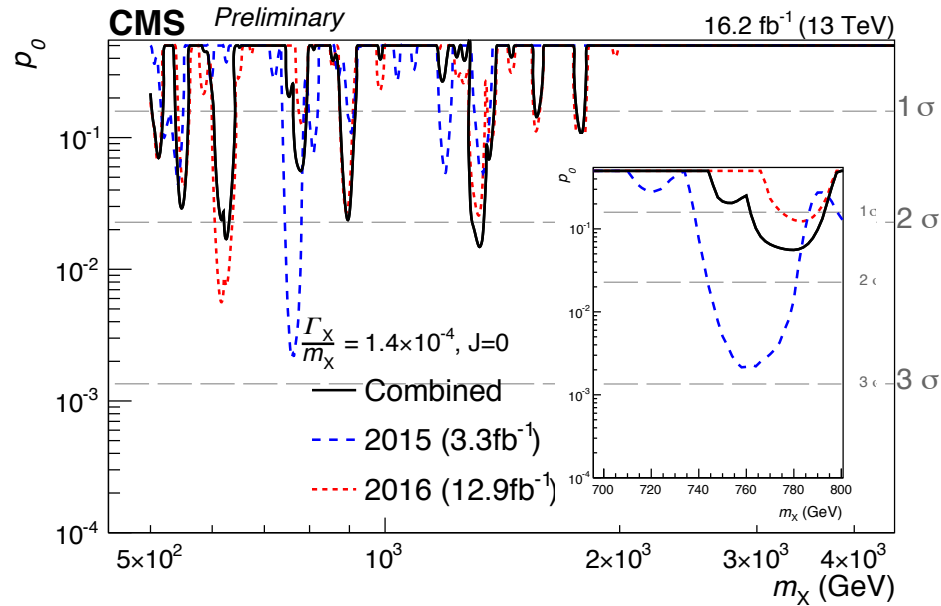


Spin-2

2016 p-value (intermediate width)

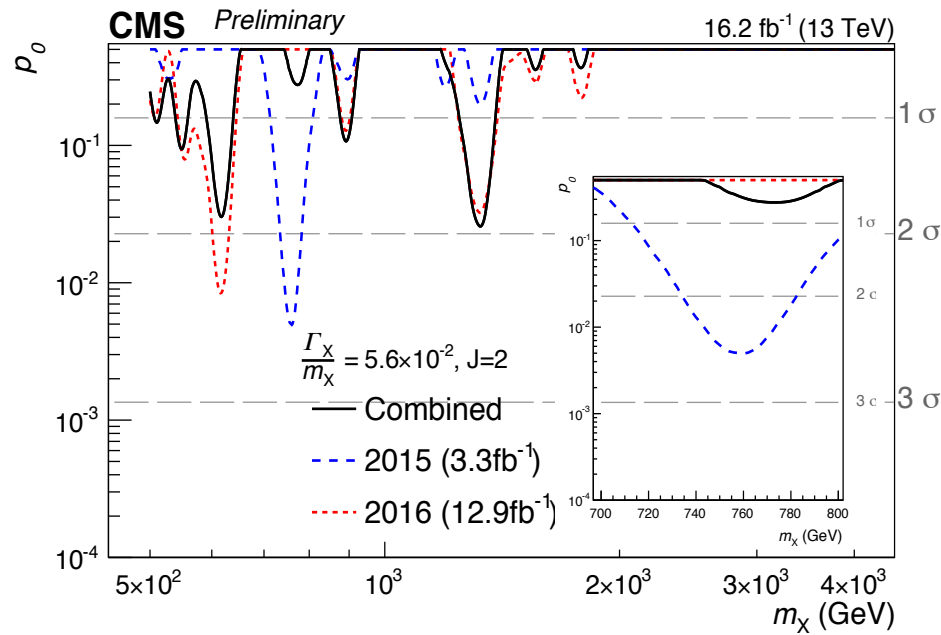
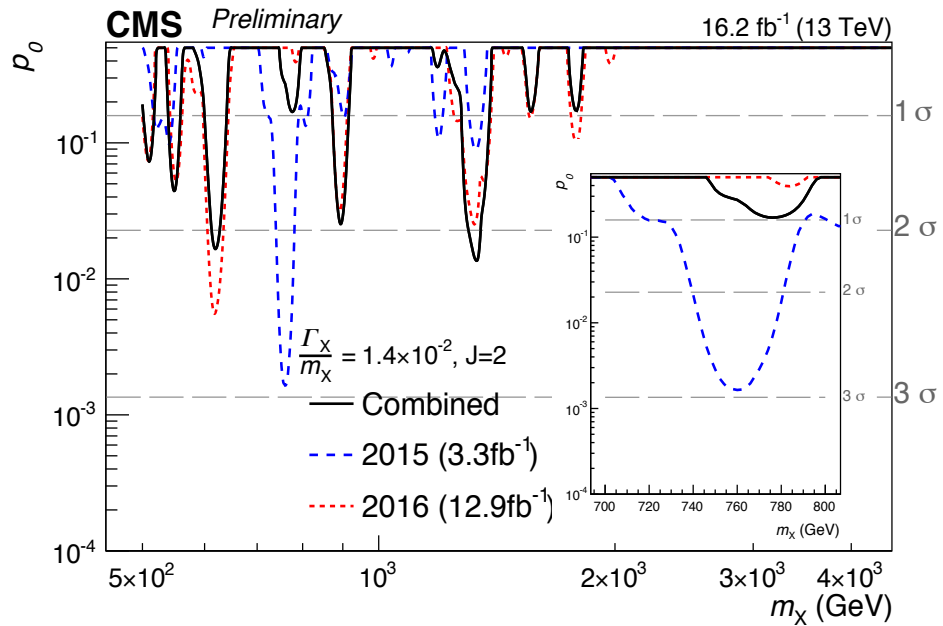
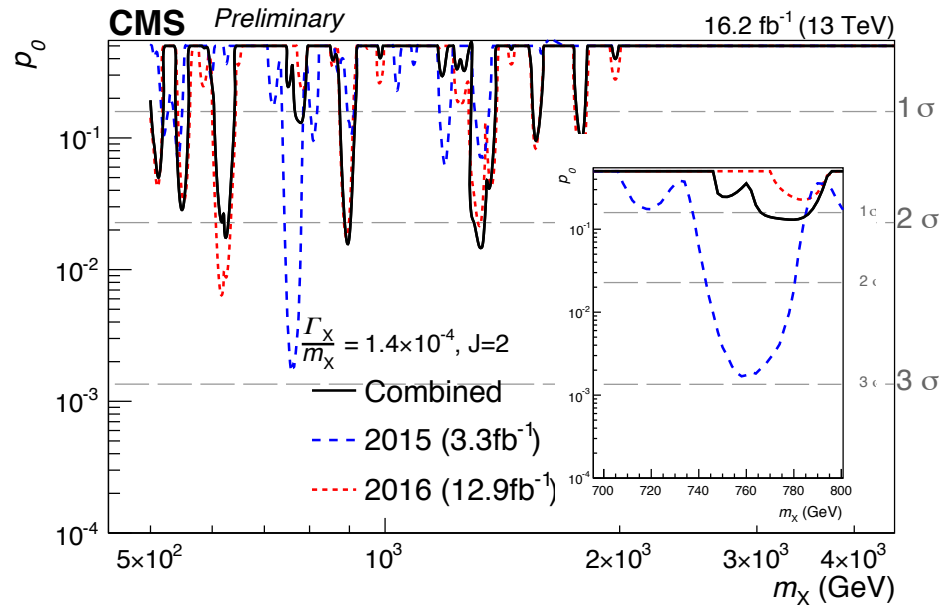


2015 + 2016 p-value



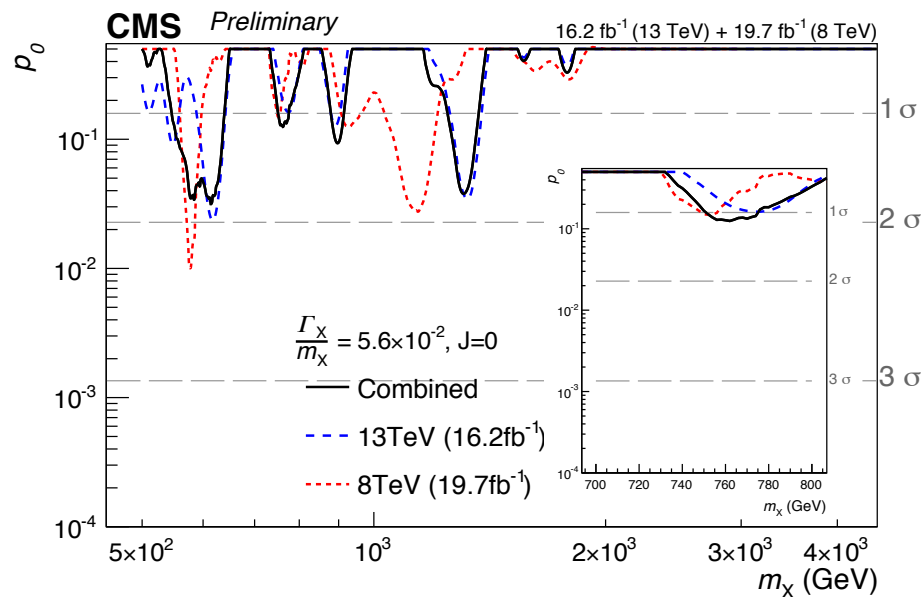
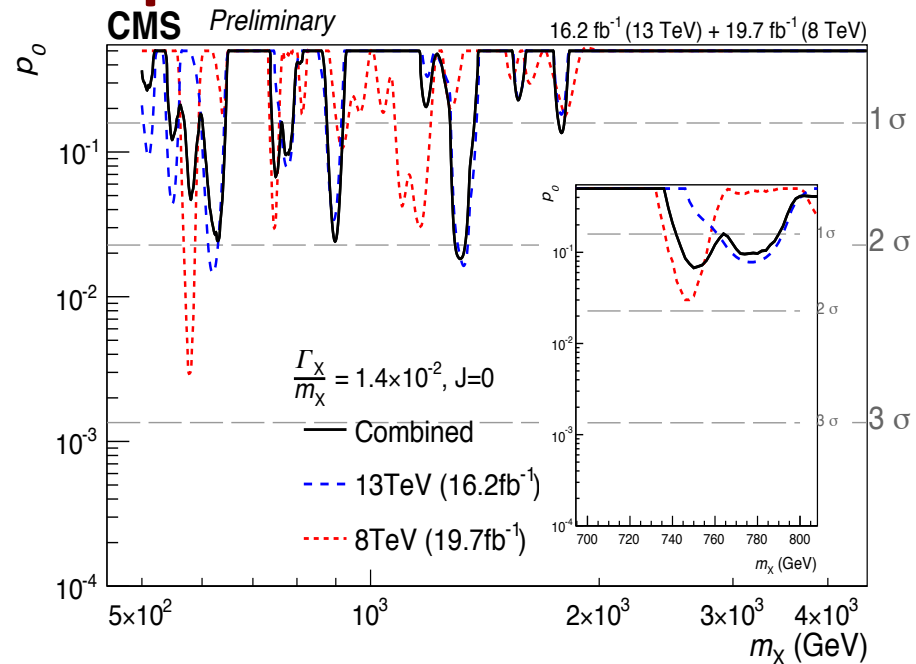
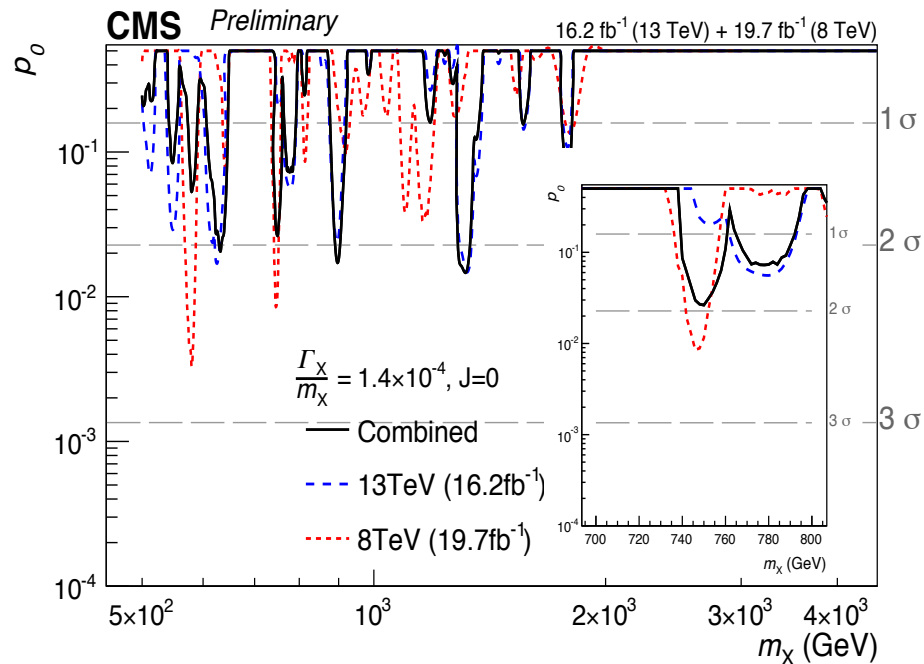
Spin-0

2015 + 2016 p-value



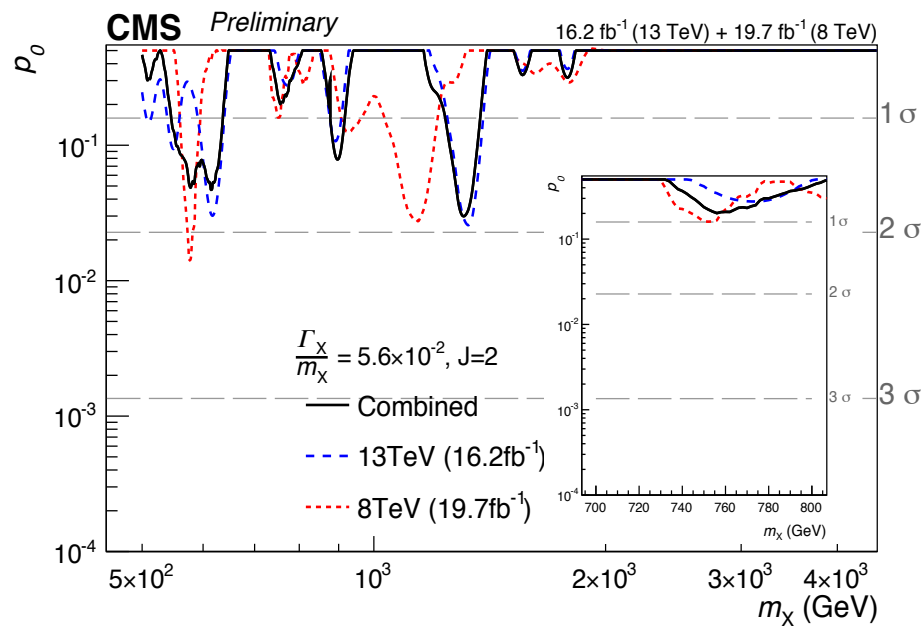
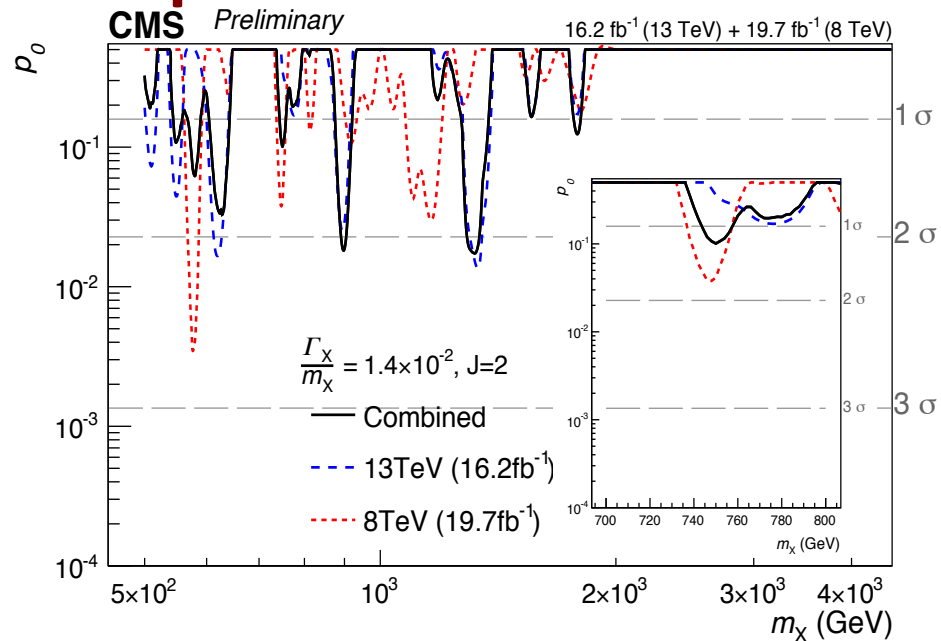
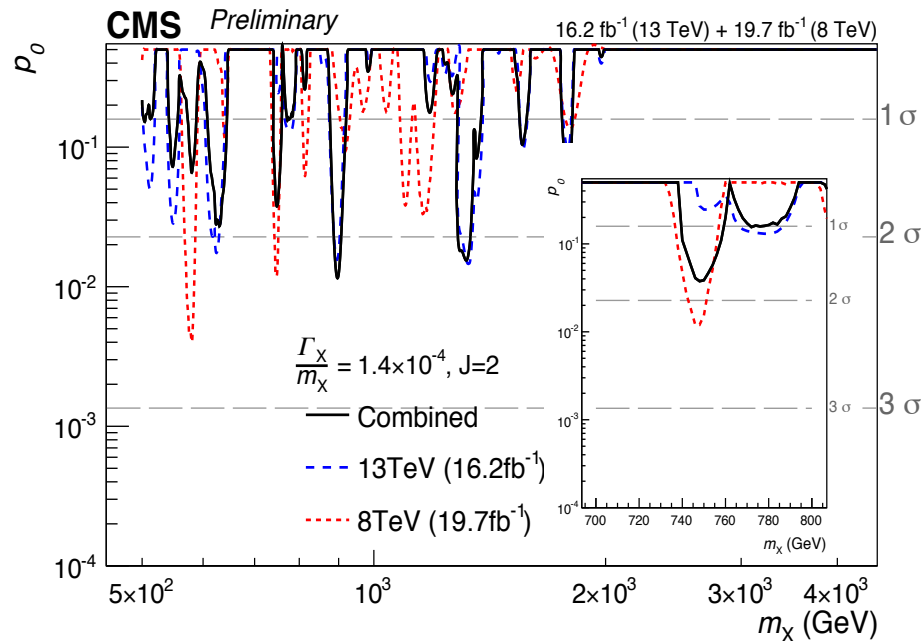
Spin-2

8TeV + 13TeV p-value



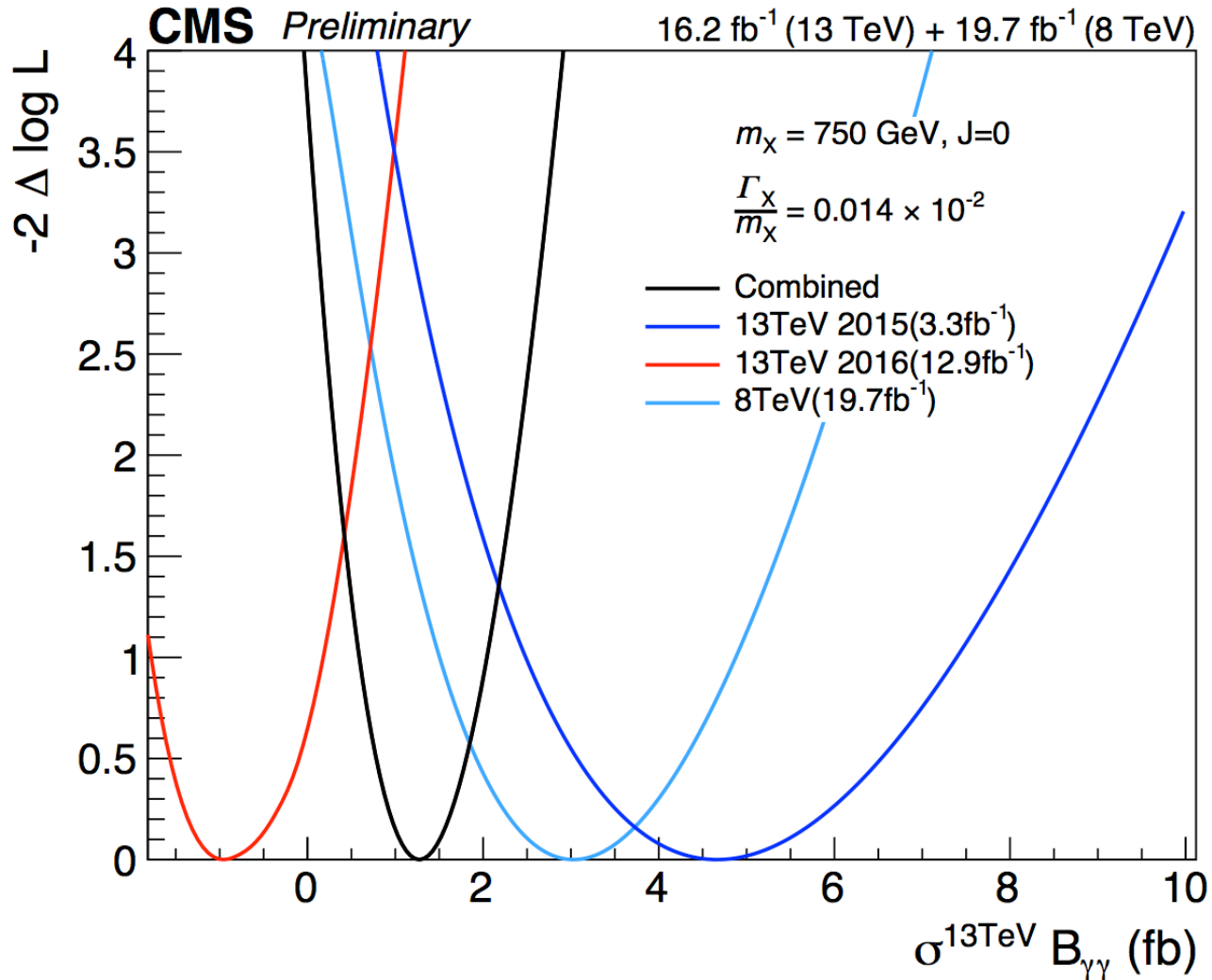
Spin-0

8TeV + 13TeV p-value



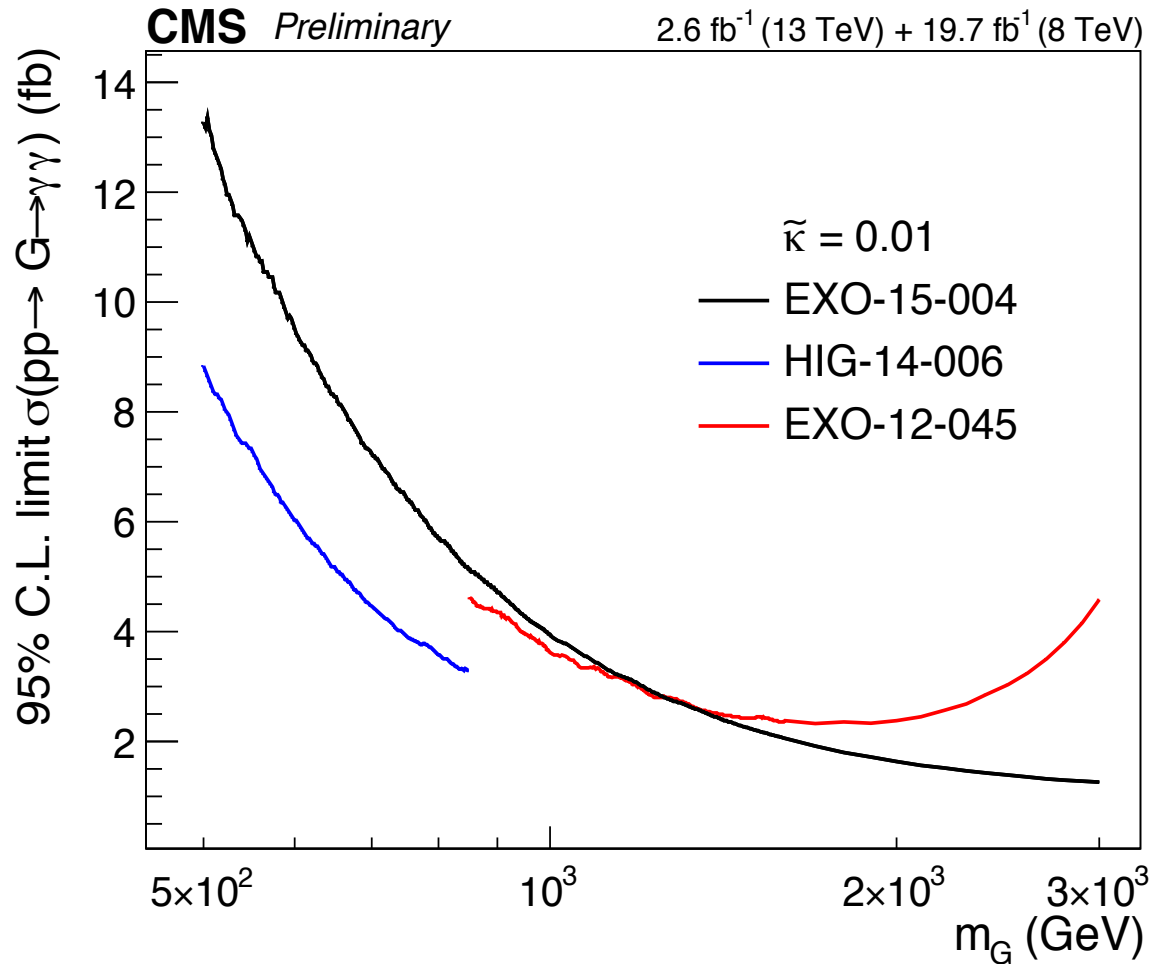
Spin-2

2016+2015+8TeV: σ likelihood scan

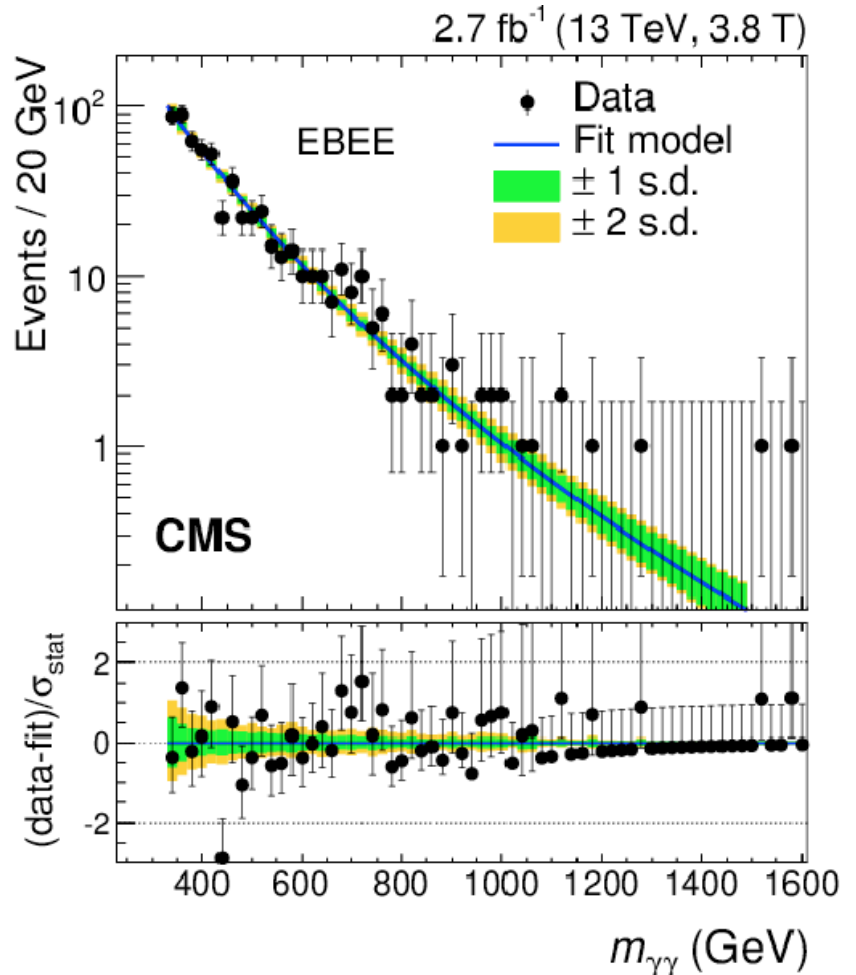
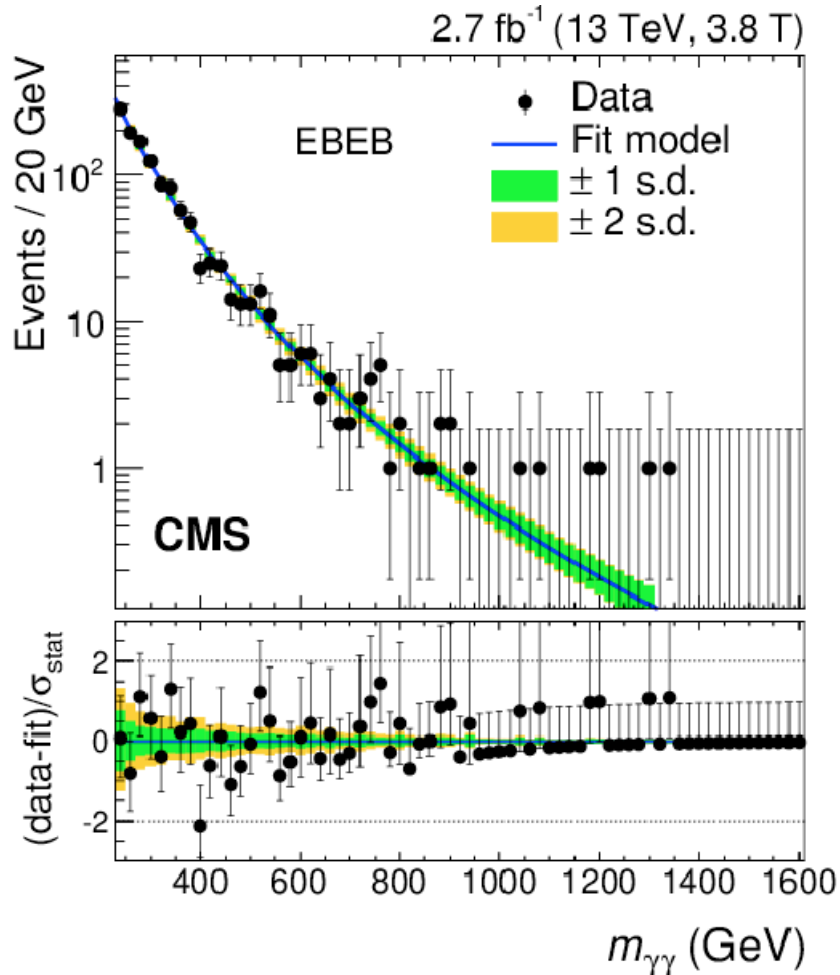


$\Gamma_x/m_x = 1.4 \times 10^{-4}$, spin-0, $m_x=750\text{GeV}$

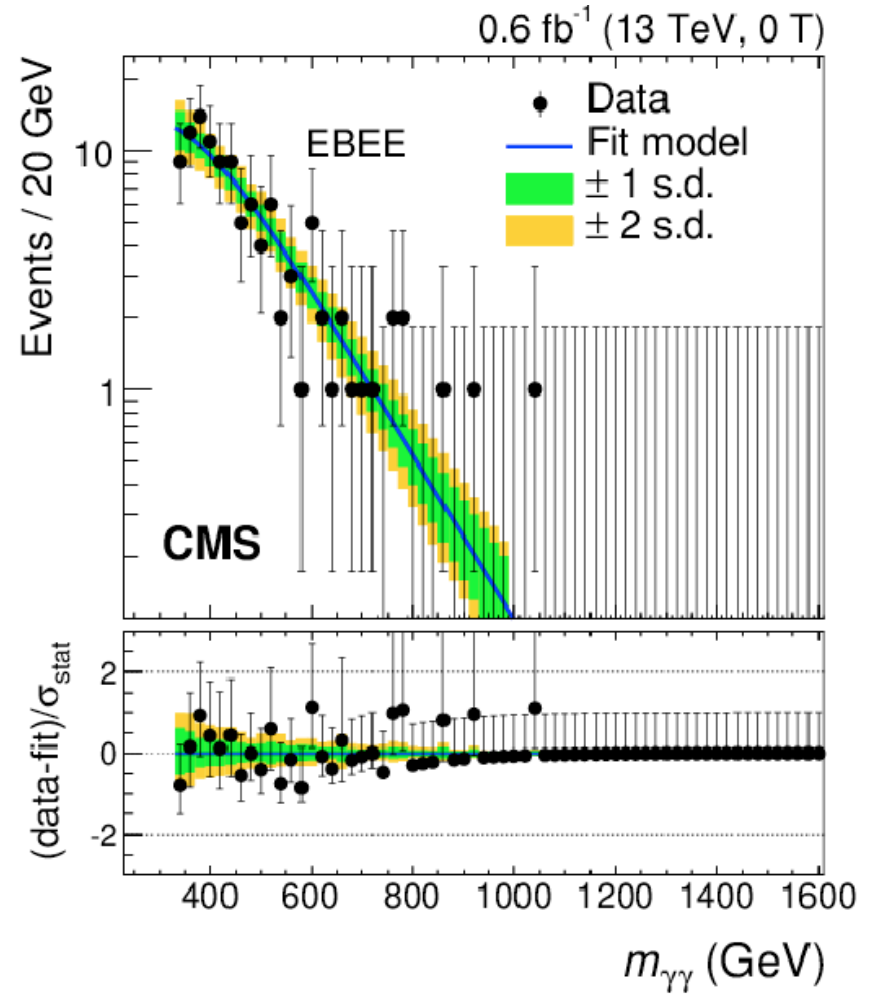
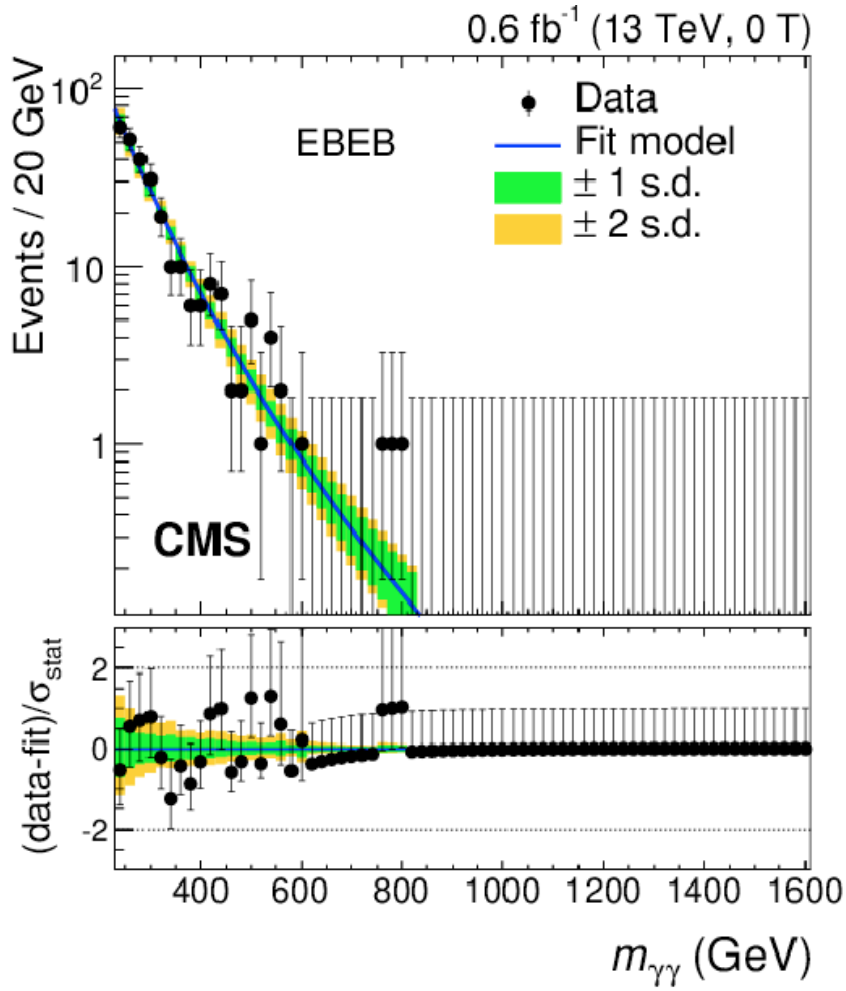
8TeV analysis



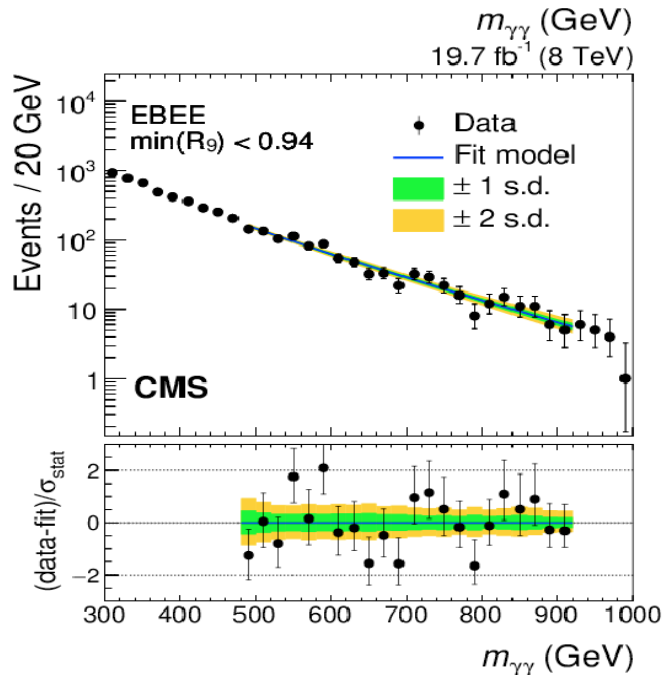
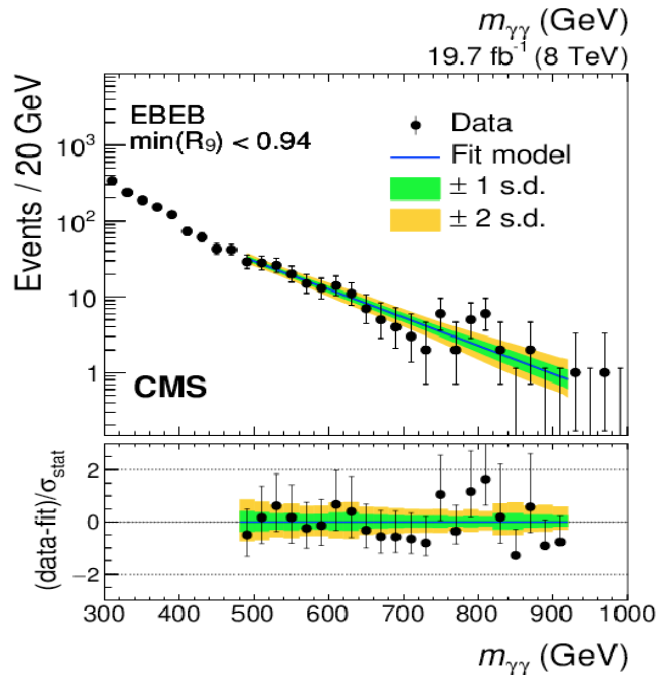
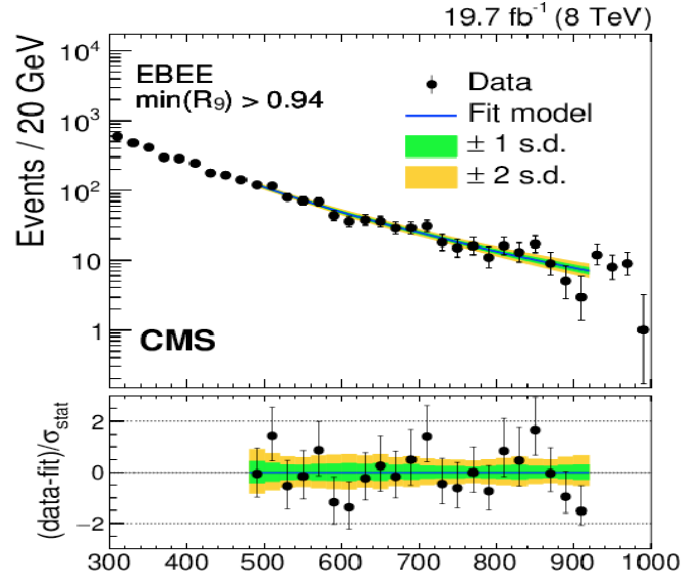
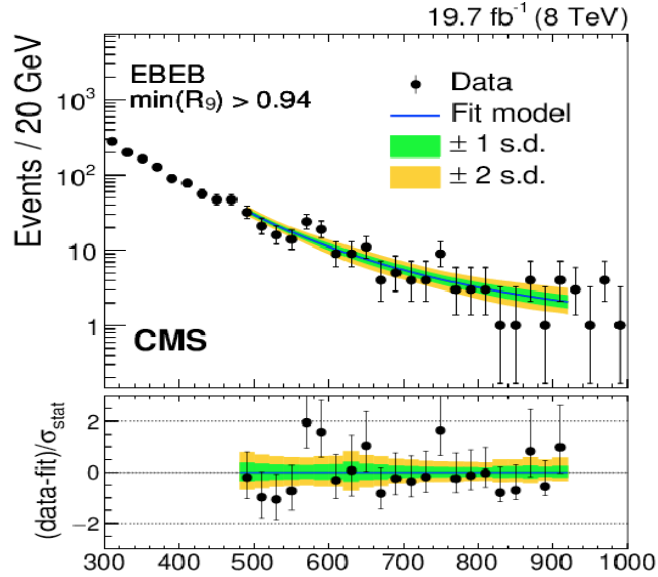
2015 data spectra, 3.8T



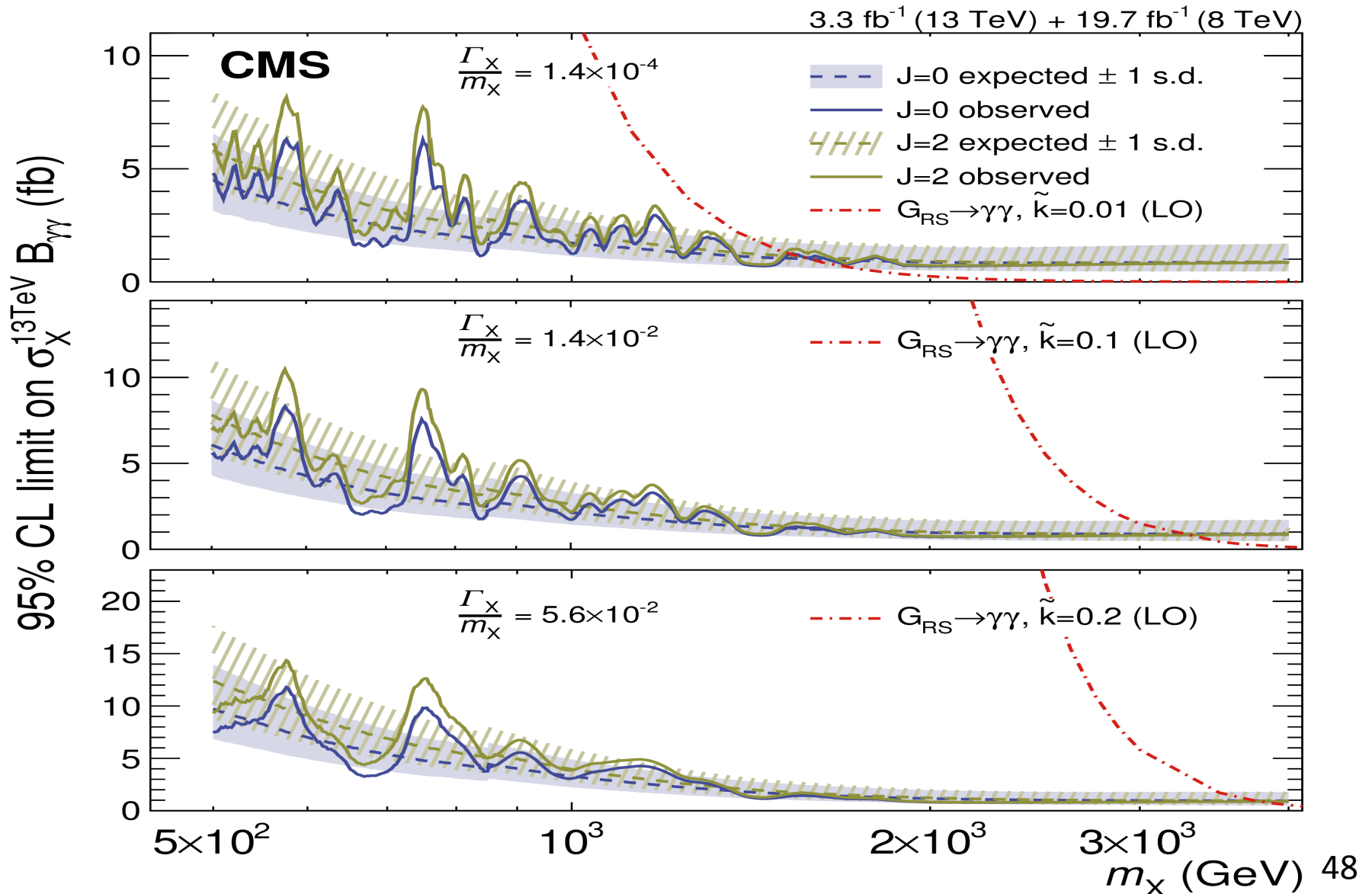
2015 data spectra, 0T



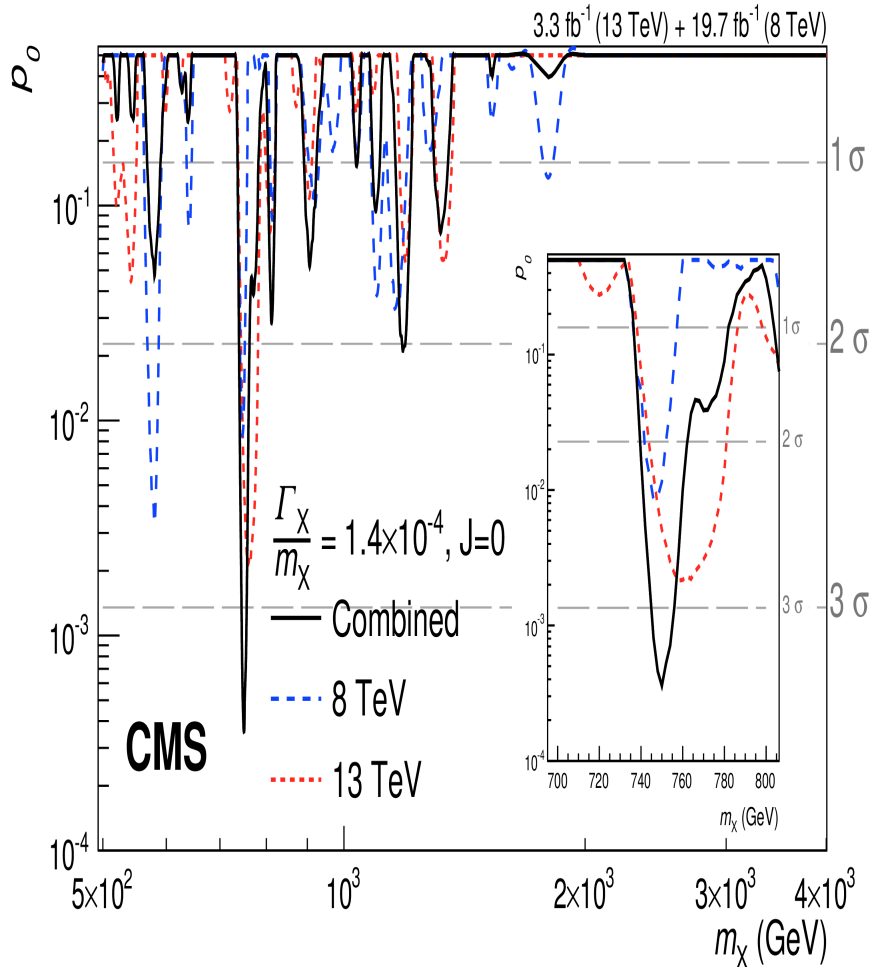
8TeV data spectra



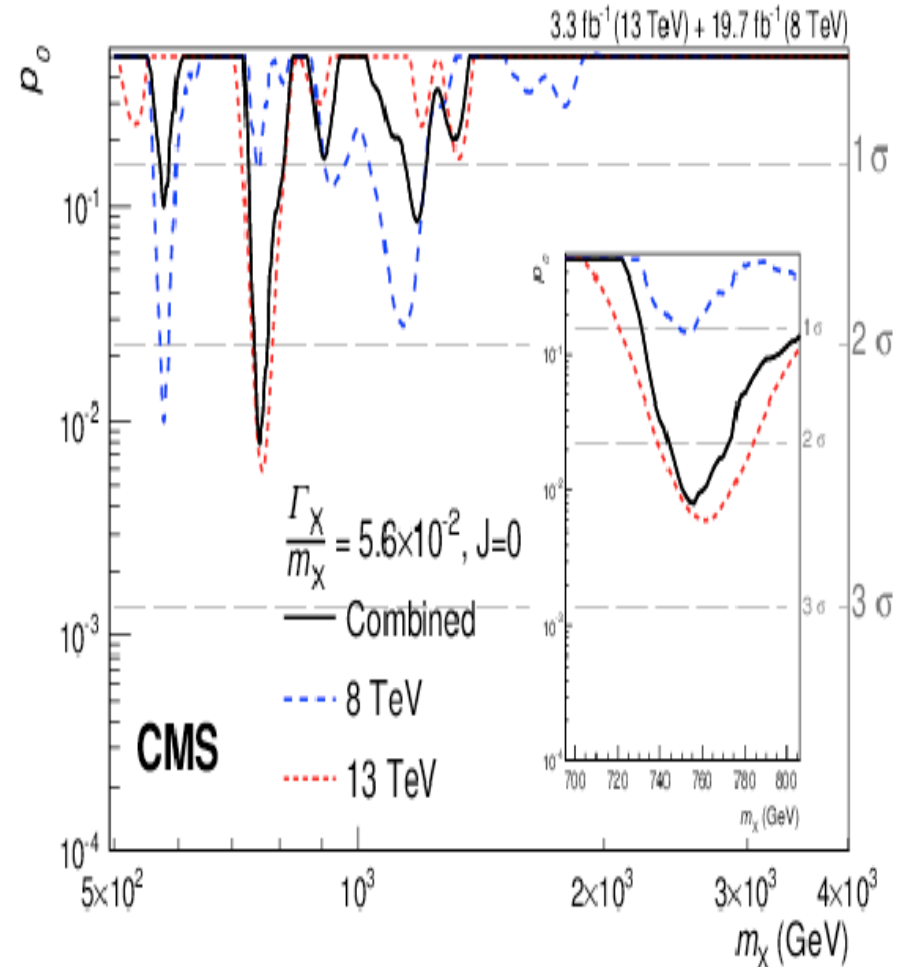
8TeV + 2015 limits



8TeV + 2015 pvalue



$$\Gamma_X/m_X = 1.4 \times 10^{-4}, J=0$$



$$\Gamma_X/m_X = 5.6 \times 10^{-2}, J=0$$

Crystals transparency

