



Cuoricino and CUORE detectors: developing big arrays of large mass bolometers for rare events physics

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On behalf of CUORE Collaboration



Outline

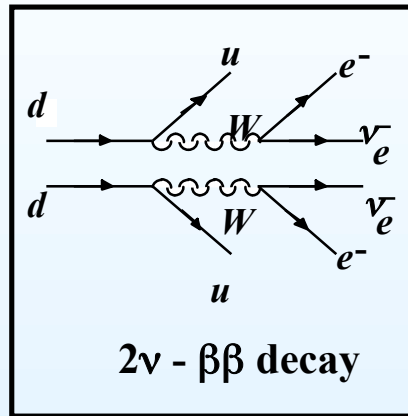
- Physical goal: search of neutrinoless DBD
 - Bolometric approach in Cuoricino/CUORE
 - Cuoricino
 - Present performance
 - Developments for Cuore
 - Background reduction
 - Resolution improvements
 - Detector reproducibility
 - Conclusion
- The Cuoricino results on neutrino physics



Neutrino properties and DBD

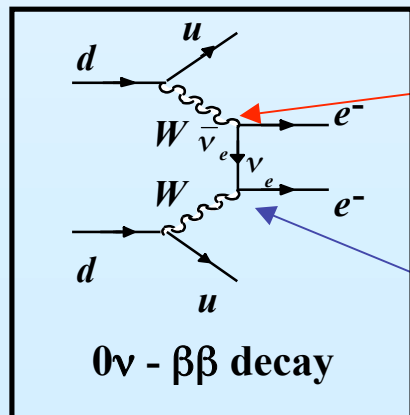
Nuclear $\beta\beta$ decay is a rare nuclear process; **Two** interesting decay modes are usually discussed:

① **DBD 2ν** : $(A,Z) \rightarrow (A,Z+2) + 2e^- + 2\bar{\nu}_e$



2nd order electroweak process allowed by SM, has already been observed in various nuclei $\tau^{2\nu} \sim 10^{19} - 10^{21}$ y

② **DBD 0ν** : $(A,Z) \rightarrow (A,Z+2) + 2e^-$



a **RH (L=1) antineutrino** is emitted

a **LH (L=-1) neutrino** is absorbed

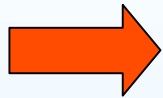
In pre-oscillation standard particle physics (massless neutrino), the process is forbidden because neutrino has not the correct **helicity/lepton number** to be adsorbed at the second vertex



Neutrino properties and DBD

But from neutrino oscillation measures we know that $m_\nu \neq 0$.

IF neutrinos are massive Majorana particles ($\nu = (\nu)^c$) : helicity can be accommodate thanks to finite mass, AND Lepton number is not relevant.

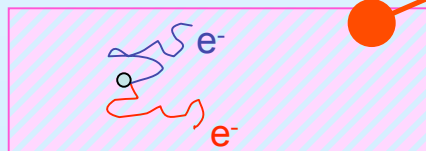


Measuring 0ν -DBD is a very important test of neutrino nature



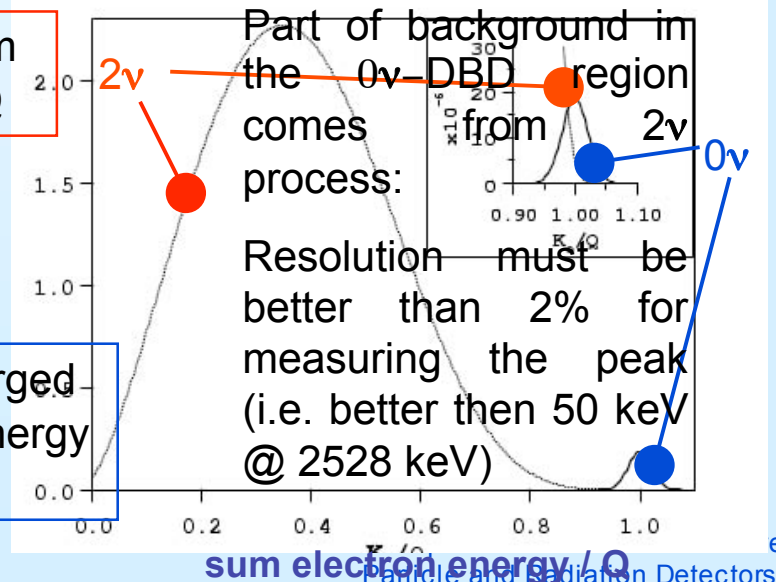
Experimental approach

Cuoricino (and CUORE) are experiments for measuring 0ν -DBD of ^{130}Te , using bolometric detectors with the, so called, calorimetric technique



Source \equiv Detector
(calorimetric technique)

An energy line is expected at the the Q value of the reaction. 0ν -DBD peak enlarged only by detector's energy resolution
 $Q(^{130}\text{Te}) = 2528 \text{ keV}$

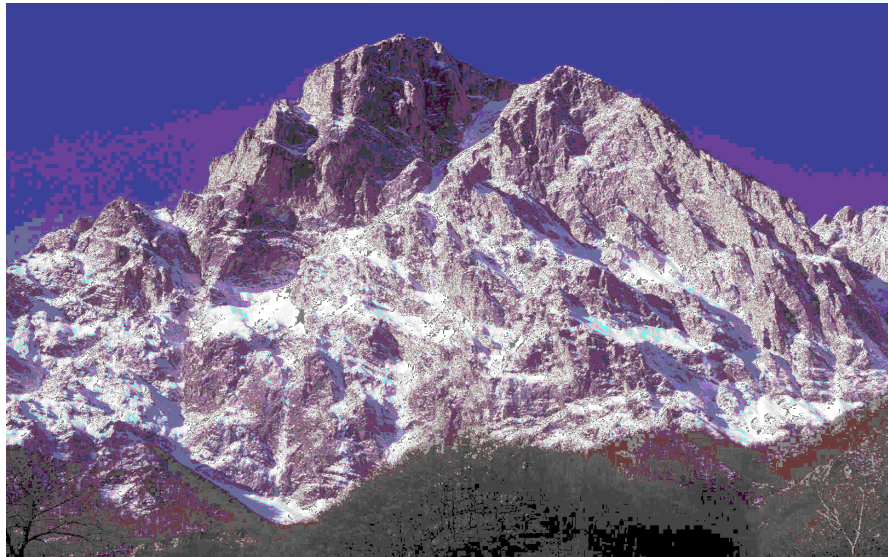


Part of background in the 0ν -DBD region comes from 2ν process:

Resolution must be better than 2% for measuring the peak (i.e. better than 50 keV @ 2528 keV)



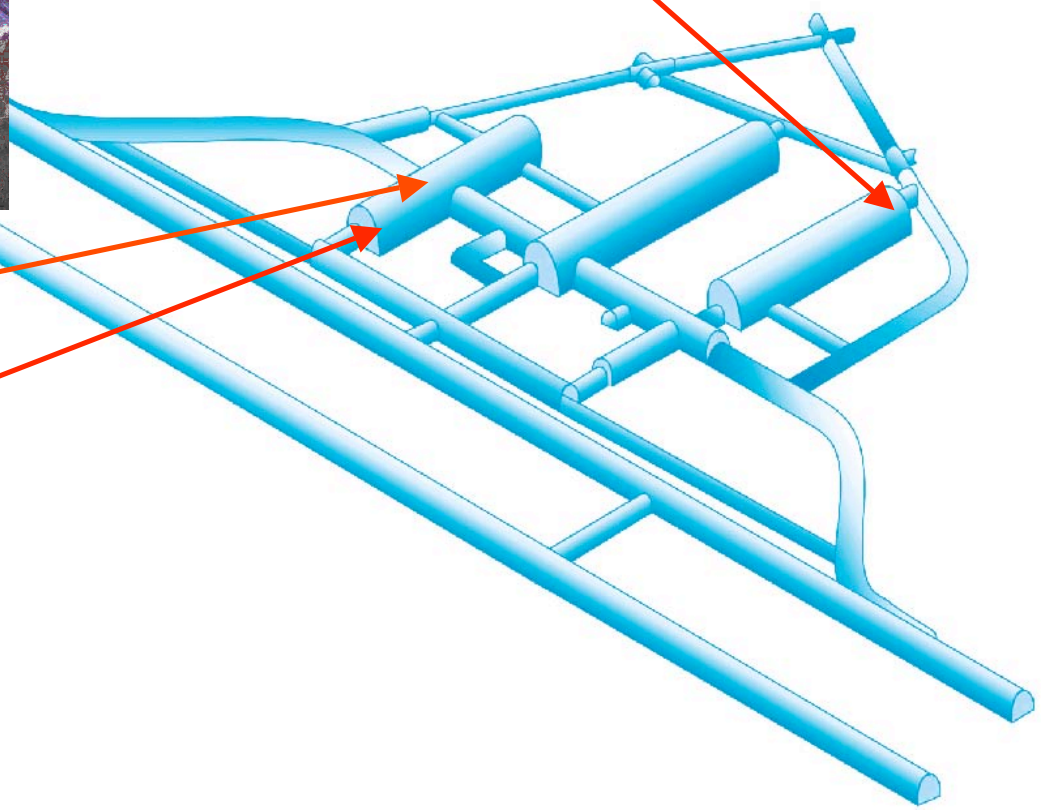
CUORE @ LNGS



CUORE R&D (Hall C)

CUORE location (Hall A)

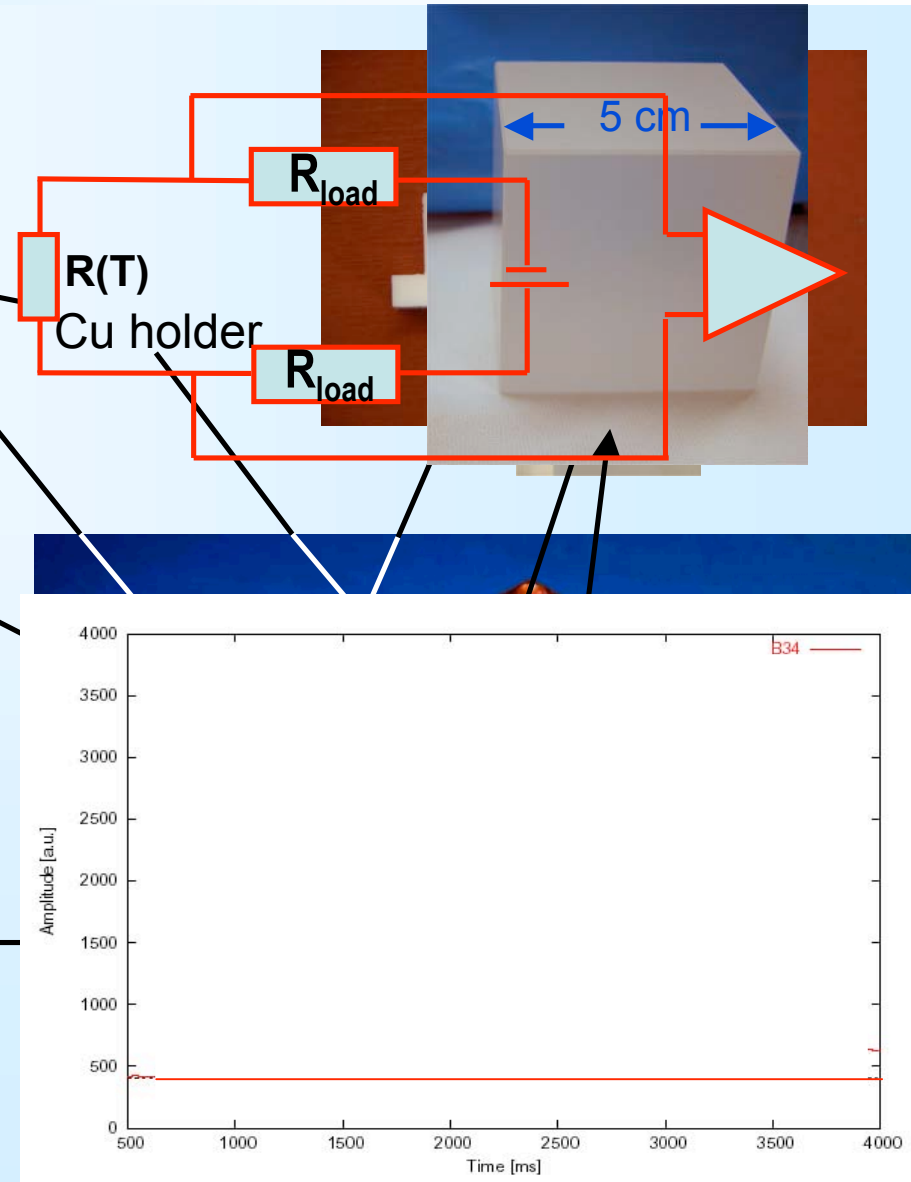
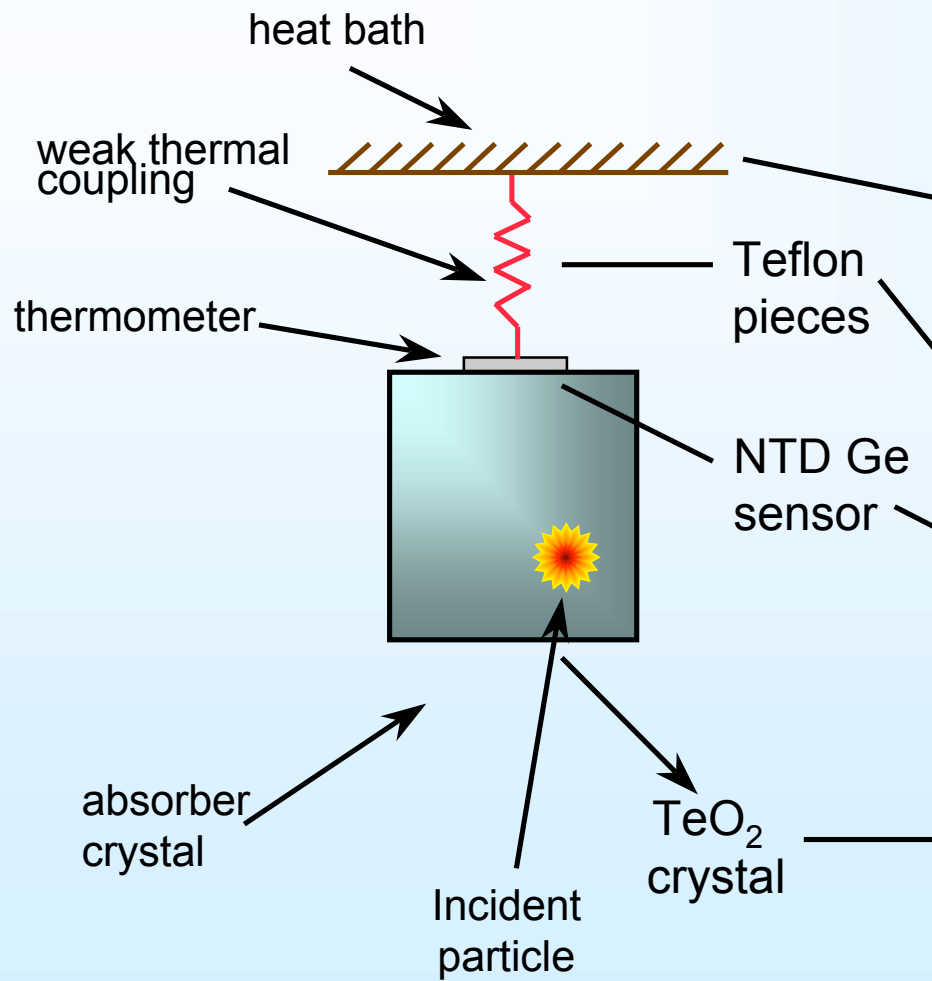
Cuoricino (Hall A)





Cuoricino and CUORE

Cuoricino is an array of 62 bolometers:





Detector sensitivity for DDB-0 ν

Sensitivity Lifetime corresponding to the minimum number of detectable events above background at a given C.L.

$$S^{0\nu} = \ln 2 \times N_A \times \frac{a}{A} \left[\frac{M T}{b \Gamma} \right]^{1/2} \times \epsilon$$

The diagram illustrates the components of the sensitivity equation for DDB-0 ν . The equation is $S^{0\nu} = \ln 2 \times N_A \times \frac{a}{A} \left[\frac{M T}{b \Gamma} \right]^{1/2} \times \epsilon$. The variables are defined as follows:

- a : Isotopic abundance
- A : Atomic mass
- M : Detector mass (kg)
- T : Measurement Time (y)
- b : Background (counts/keV/kg/y)
- Γ : Energy resolution (keV)
- ϵ : Efficiency

Detector improvements must be on:

- Increase detector mass
- Background reduction (and if possible rejection)
- Increase energy resolution



Increasing the mass

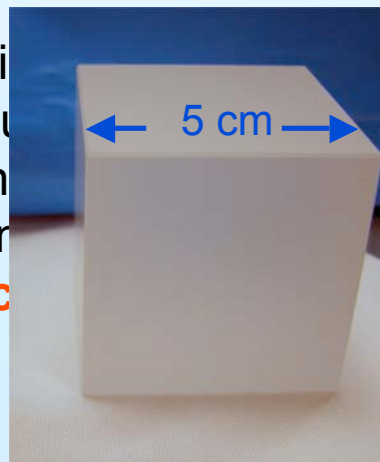
Two ways:

1. Increasing the absorber crystal of the single detector
2. Increasing number of detectors

- ① **1991** : first TeO₂ bolometer, 73 g of mass
1993 : first 340 g TeO₂ bolometer, 3x3x6 cm³
2000 : first 790 g TeO₂ bolometer, 5x5x5 cm³
2004 : first 1300 g TeO₂ bolometer, 6x6x6 cm³



Thermal noise oscillation due to the crystal (mainly due to this type of material) is proportional to the square root of C. **S/N is constant**



i.e. baseline fluctuation on the detector. The amplitude of the increasing

Temperature variation measure by NTD

BUT

Energy release

$$\Delta T = E / C$$

Absorber heat capacity

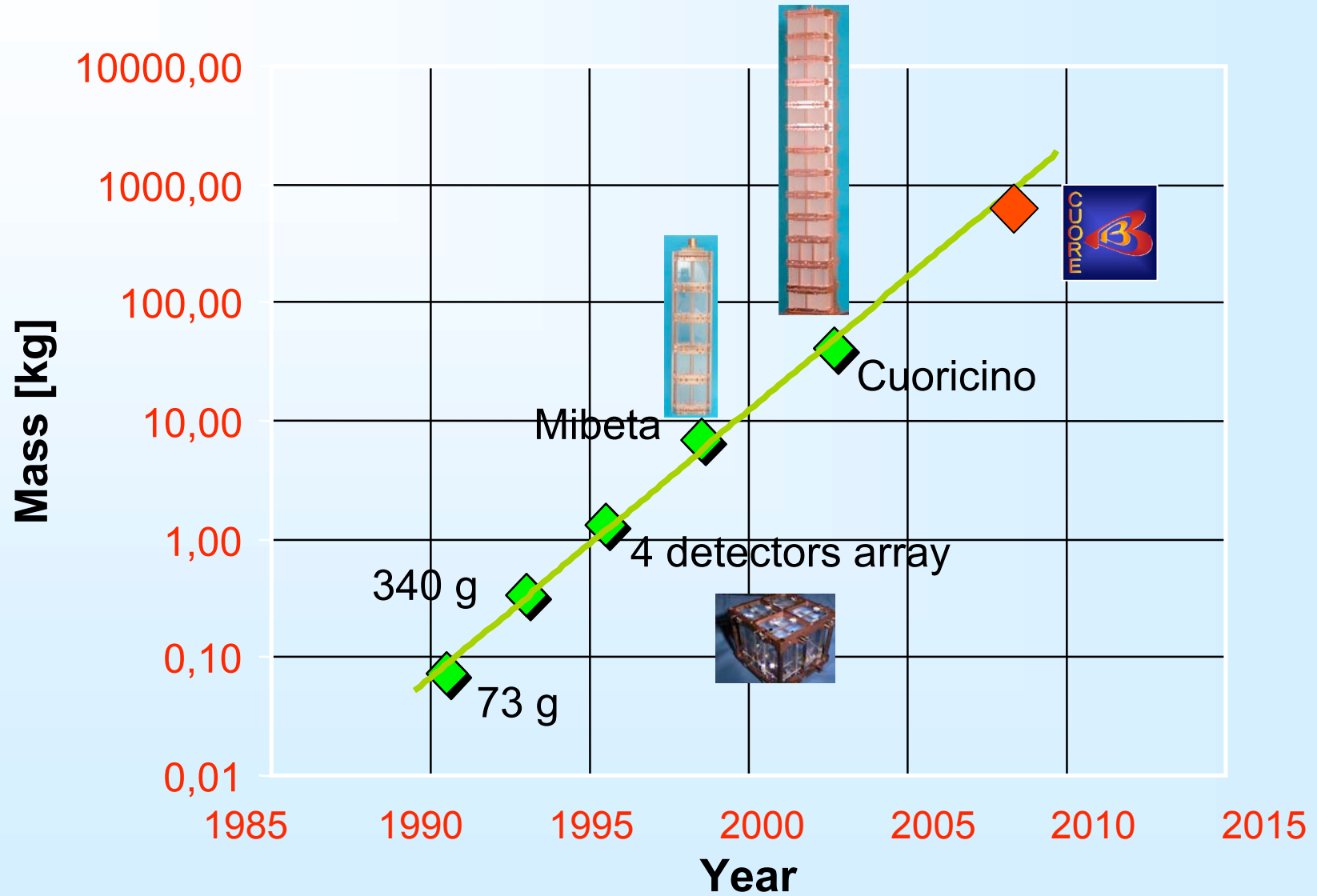
the only limitation is given by the heat capacity of the crystal that is proportional to volume.

The real limitation is given by the growing crystals of big dimensions. **Best choice: 5x5x5 cm³**



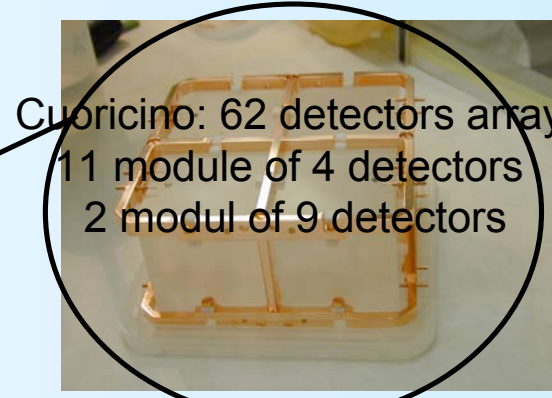
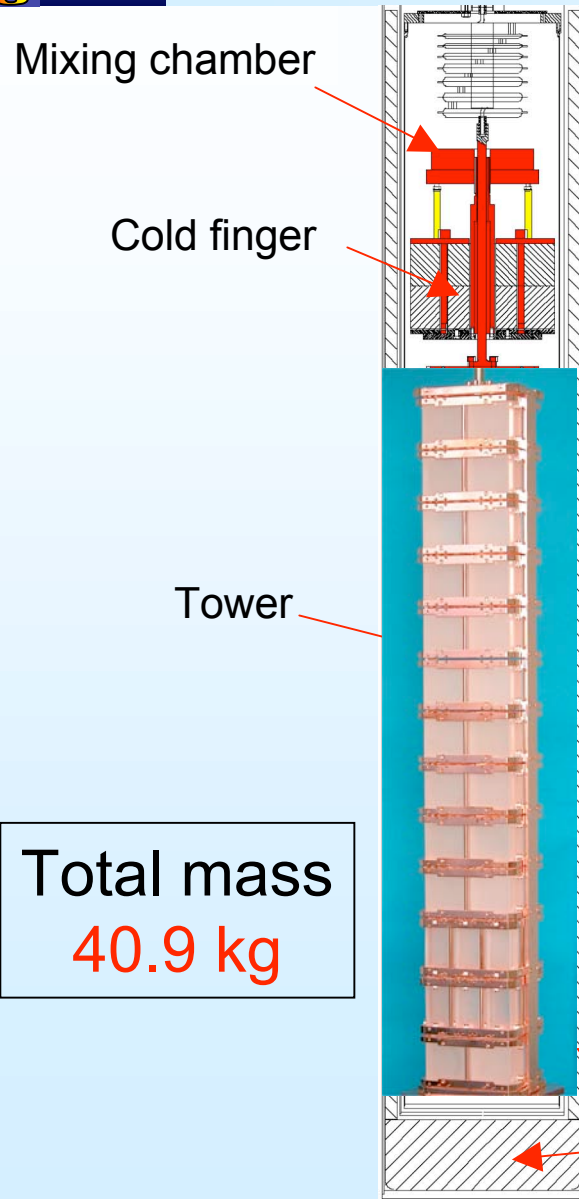
Increasing the mass

②



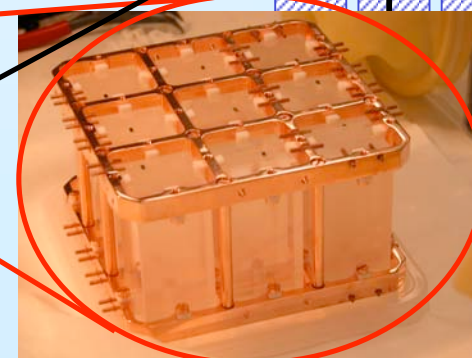
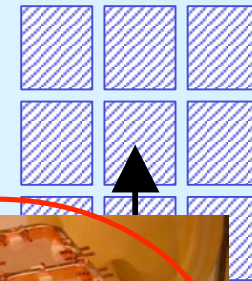


Cuoricino



4 detectors 5x5x5 cm³, 790 g each

Planar Section



9 detectors 3x3x6 cm³, 340 g each

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Background

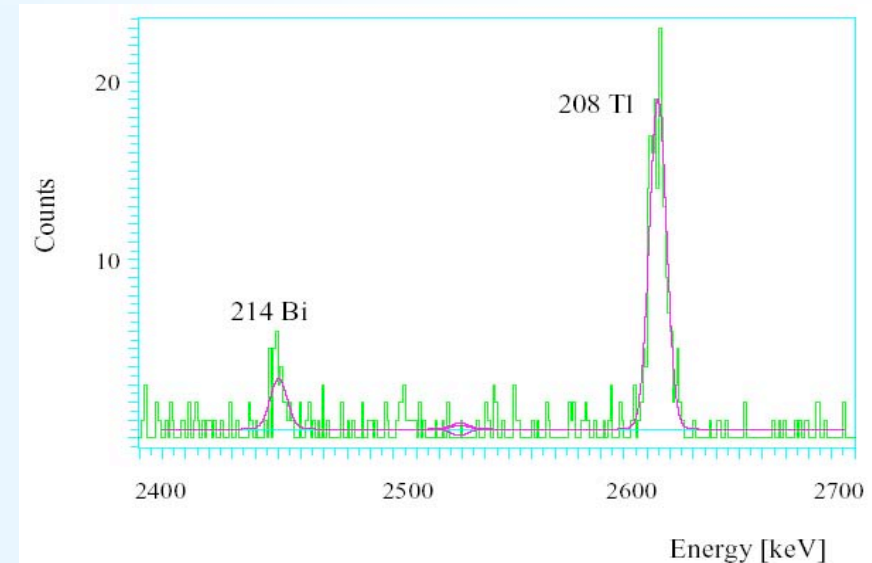
Bkg sources in the DBD region:

- Degraded α particles from material surface
- Multicompton effects and other gamma process

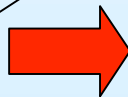


Coincidence detection on faced crystals allows:

- Reconstruction of the sum energy of the event
- Rejection of the events in anticoincidence



Surface contamination from non active material (Cu, Teflon,...) cannot be rejected with anticoincidence.



- Reduce material possible source of contamination: no matter how clean is the material you use, treatment always introduce surface contamination.
- Clean surfaces with special techniques



Detector preparation



Surface cleaning:
Lapping of the crystals surface
with radio pure powders
(18/62 in nitrogen atmosphere)

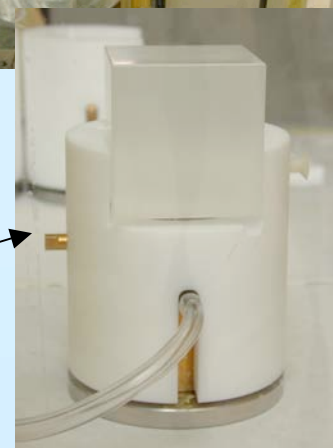


Nitrogen



Detector assembly:
Gluing thermistors and heaters:
all done in nitrogen atmosphere

Gluing tool





Clean procedures

Completely done in nitrogen atmosphere





Resolution: 790 g detector performance

In the past years a big effort was made in R&D on the $5 \times 5 \times 5 \text{ cm}^3$ TeO_2 detectors (790 g)

The best 790 g detector:

1.4 keV FWHM @ 0.351 MeV

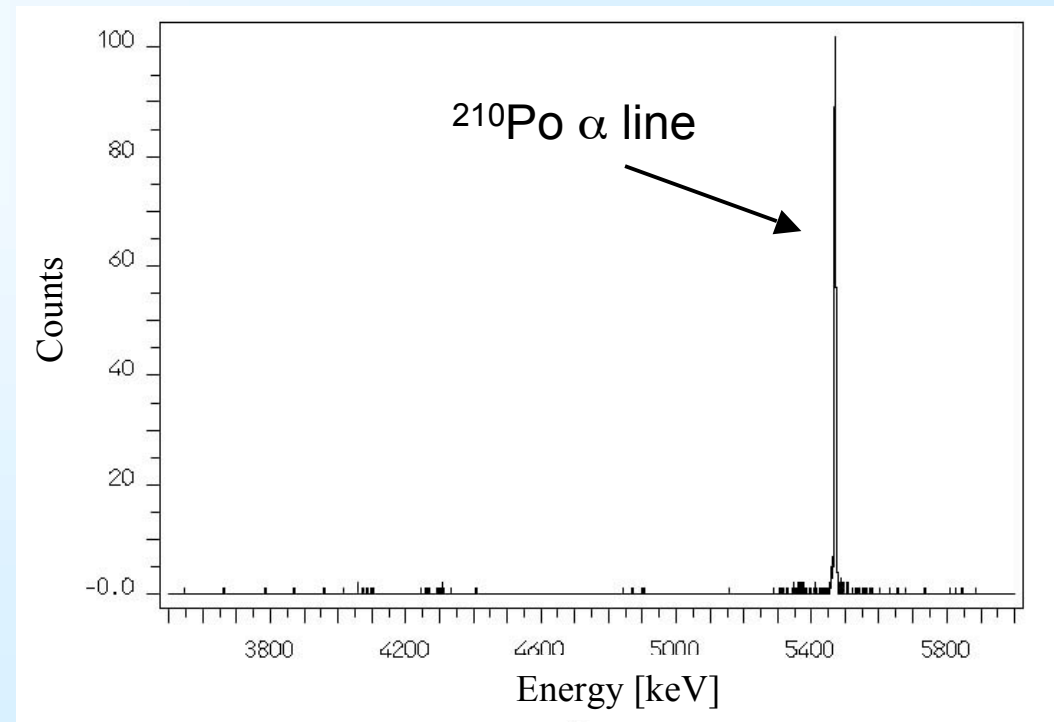
2.1 keV FWHM @ 0.911 MeV

2.6 keV FWHM @ 2.615 MeV

3.2 keV FWHM @ 5.407 MeV

(the best α spectrometer ever realized)

A. Alessandrello et al, NIMA 440 (2000) 397-402

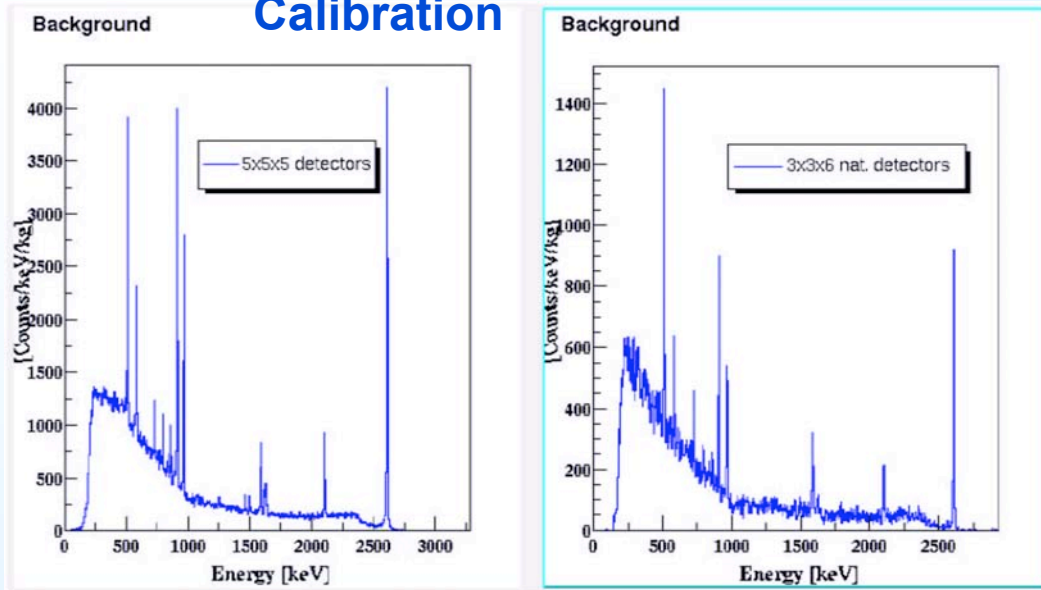


One of the best techniques for measuring very low radioactive contaminations



Resolution: Cuoricino

Calibration

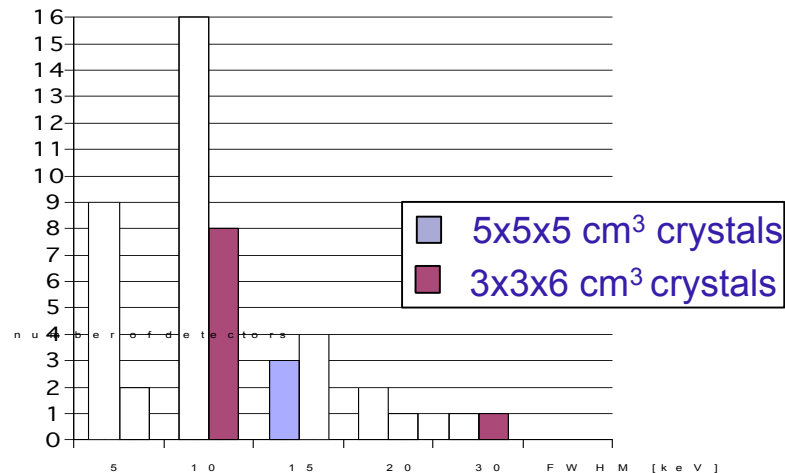


FWHM [keV] of the 2615 keV gamma line of ^{208}Tl (calibration with a ^{232}Th source ~ 3 days)

average 5x5x5 cm³ crystals ~ 7 keV

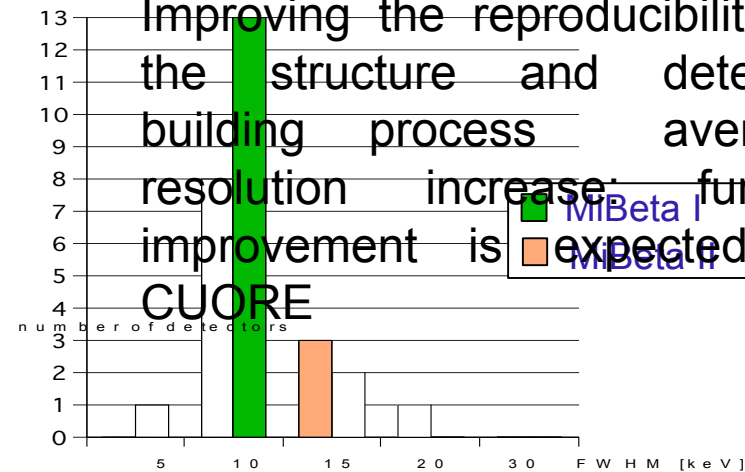
average 3x3x6 cm³ crystals ~ 9 keV

Energy resolution FWHM @ 2615 keV



Energy resolution

Improving the reproducibility of the structure and detector building process average resolution increase further improvement is expected in CUORE





Cuoricino results

After ~4 month of measure, 5.4 kg·y:

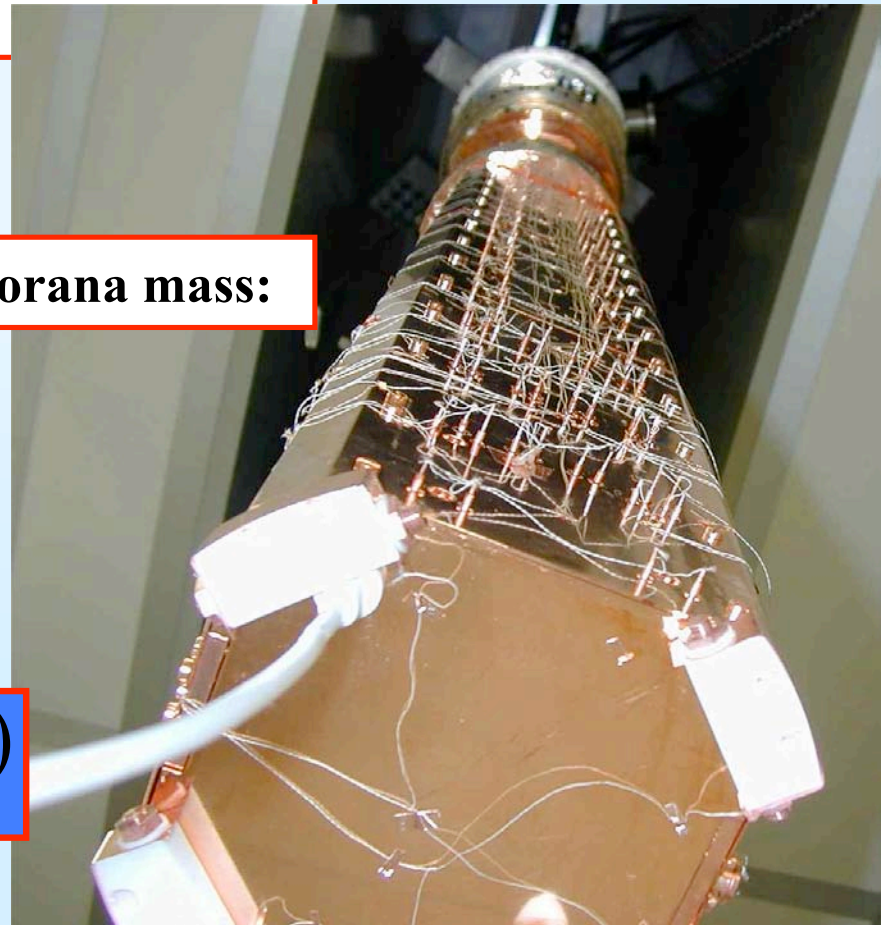
Background $\beta\beta$ region
(anticoincidence spectrum)
= 0.19 c/keV/kg/y

Total results for life time and Majorana mass:

$$\tau_{1/2}^{0\nu} \geq 7.5 \cdot 10^{23} \text{ y (90\% CL)}$$



$$\langle m_{\nu} \rangle \leq 0.31 - 1.63 \text{ eV (90\% CL)}$$





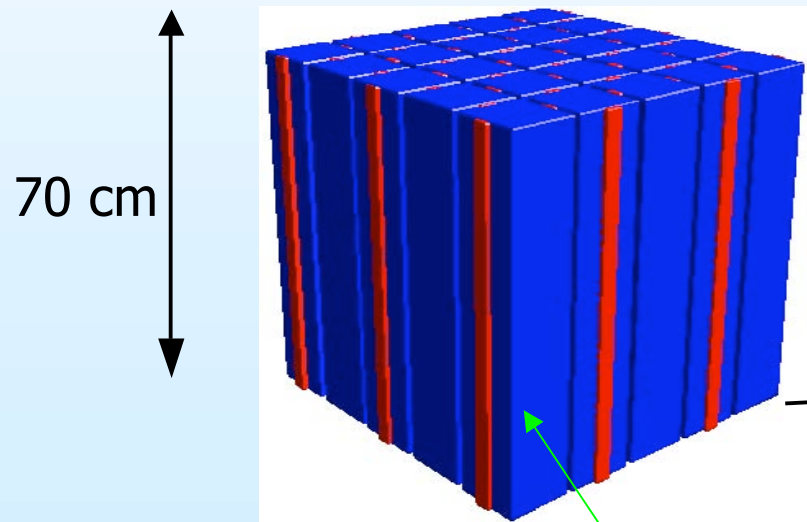
Cryogenic Underground Observatory for Rare Events

Array of 1000 detectors:
25 towers - 10 module/tower - 4 detectors/module

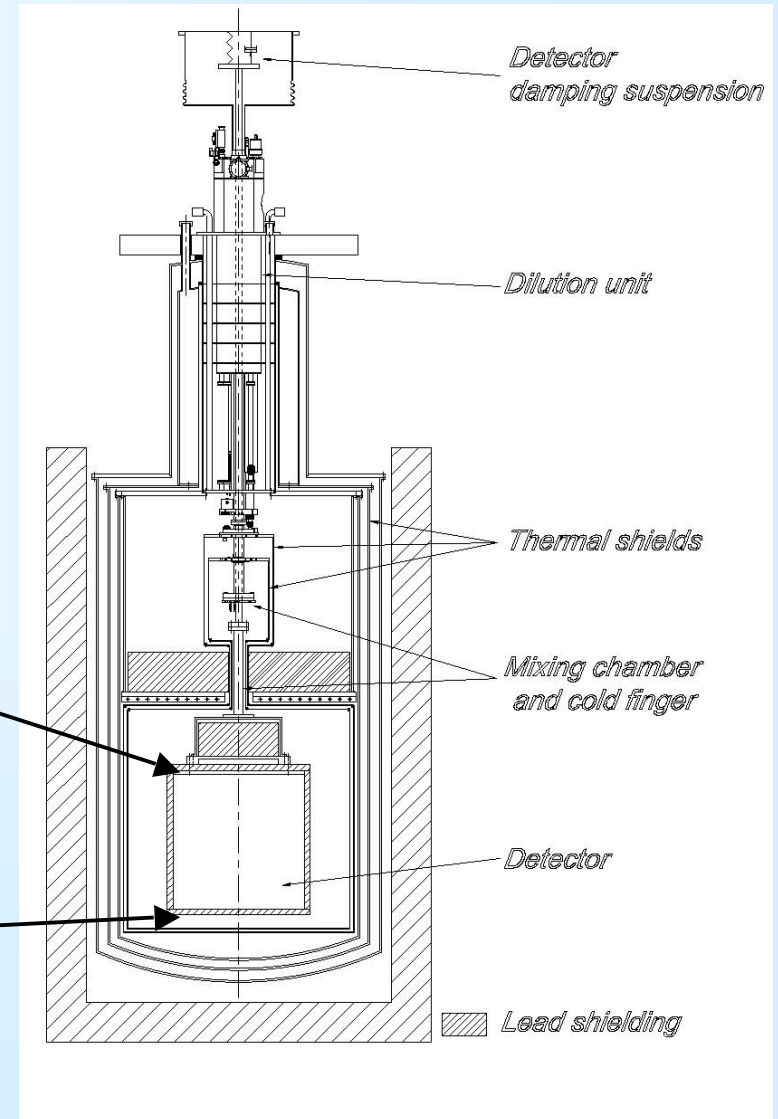
→ $M = 0.79 \text{ ton}$ di TeO_2

Sensitivity on DBD-0 ν (estimated by 1 c/keV ton y)

$T^{5y} \sim 6.6 \cdot 10^{26} \text{ y}$ $\langle m_\nu \rangle \sim 15 - 40 \text{ meV}$



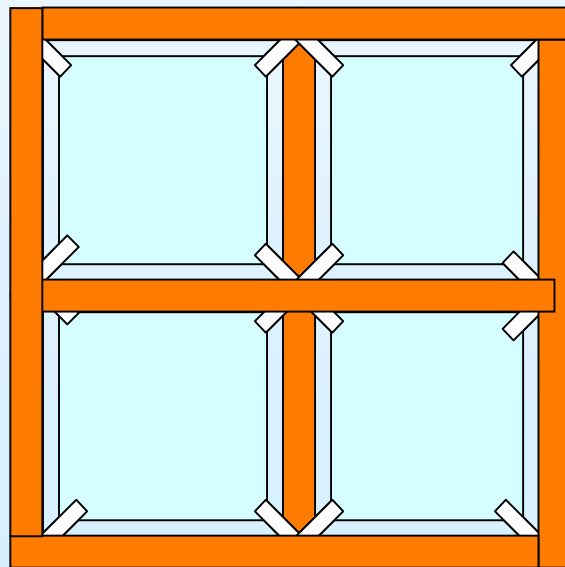
25 CUORICINO-like towers



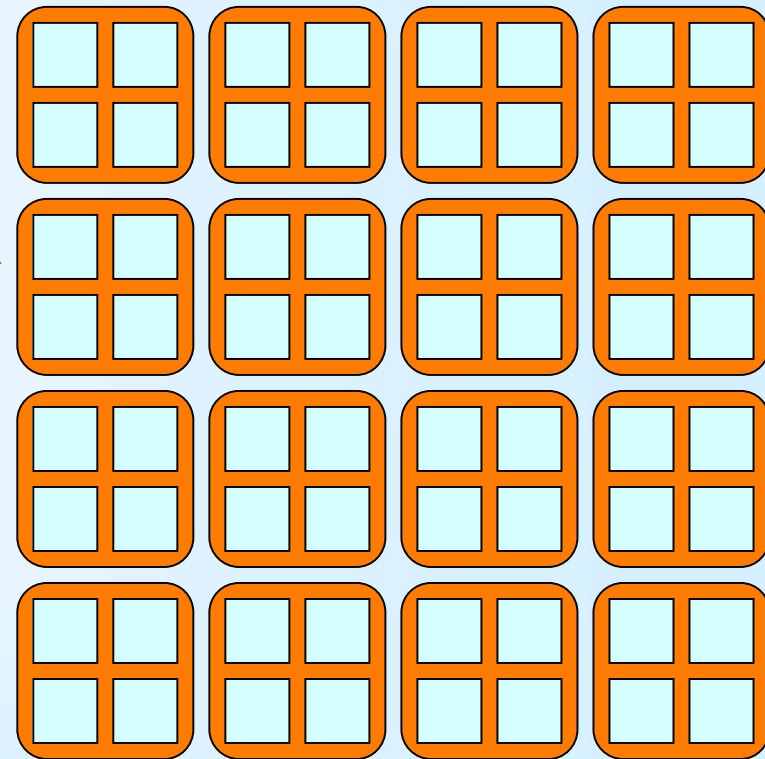


Cryogenic Underground Observatory for Rare Events

Cuore is a huge improvement not only in mass but in segmentation (anticoincidence) and modularity (resolution)



More stable older



A CUORE slice



Conclusion

- Big arrays of macrobolometers are reliable detectors for rare events physics
- Cuoricino demonstrate the possibility of having an average good energy resolution on many detector and tested the improvements of anticoincidence technique
- CUORE, a second generation detector developed on this new approaches will be build in the next 5 years.