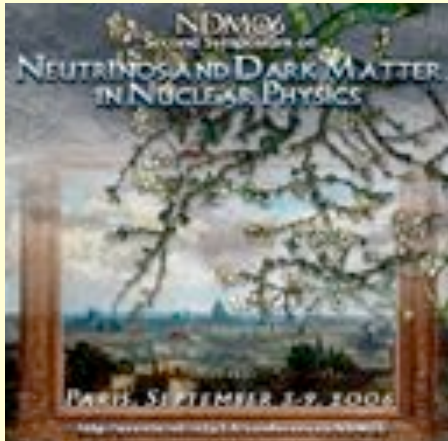




Double Beta Decay: future experiments

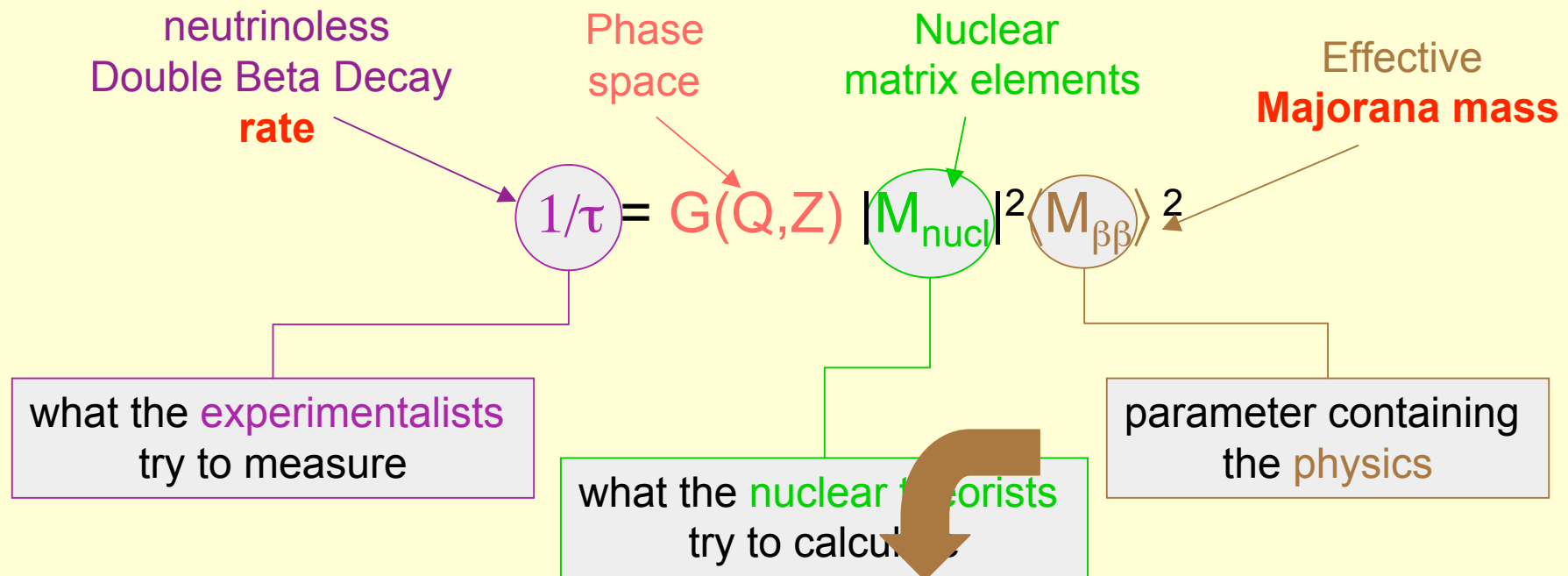
Andrea Giuliani

University of Insubria, Como, and INFN Milano

- crucial role of DBD for the future of neutrino physics
- from where we start...and where we want to go to
- the DBD experimental program: challenges and techniques
- overview of the experiments
- discovery potential and branching points

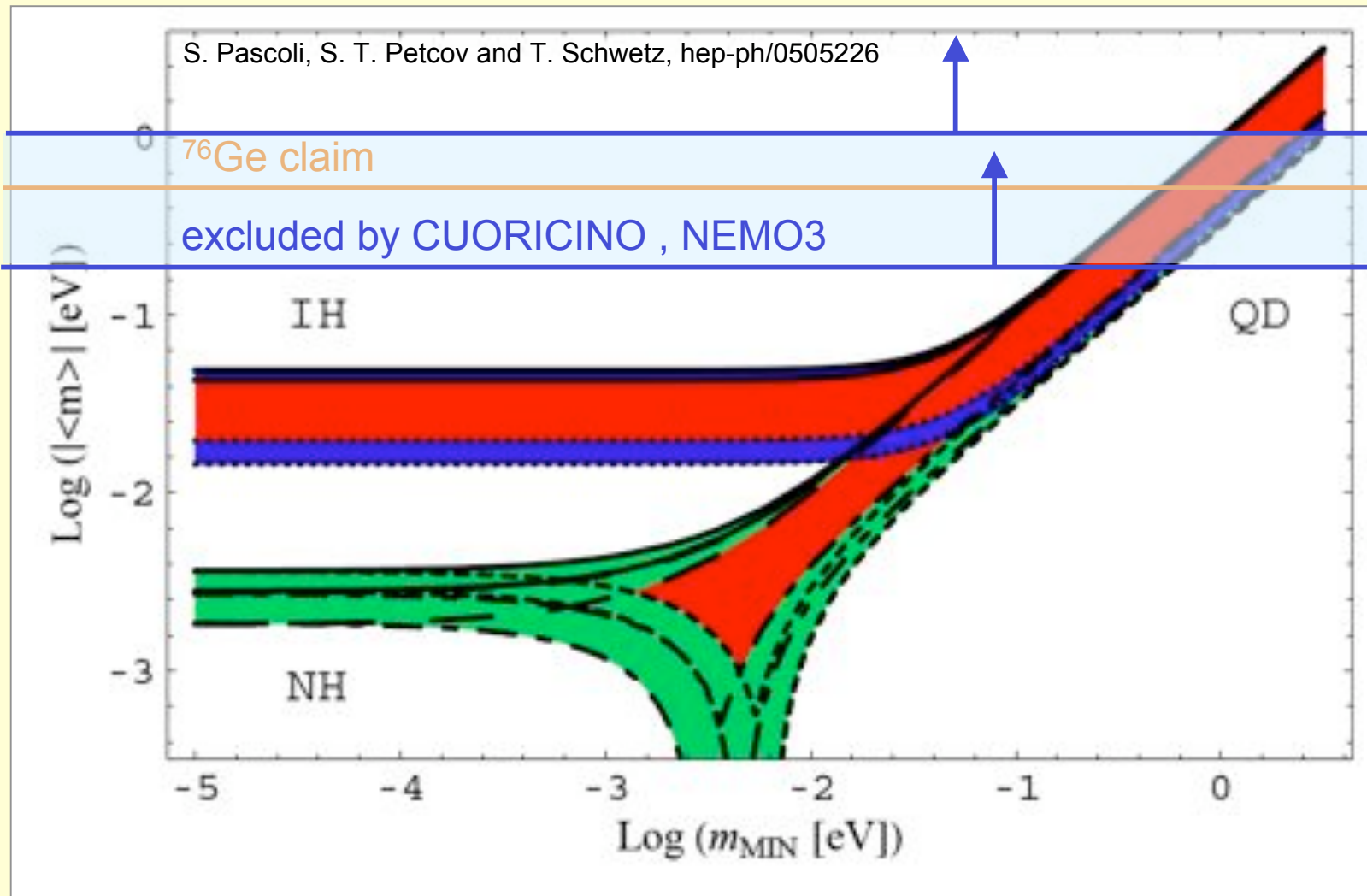
0ν -DBD and neutrino physics

how 0ν -DBD is connected to neutrino mixing matrix and masses
in case of process induced by mass mechanism

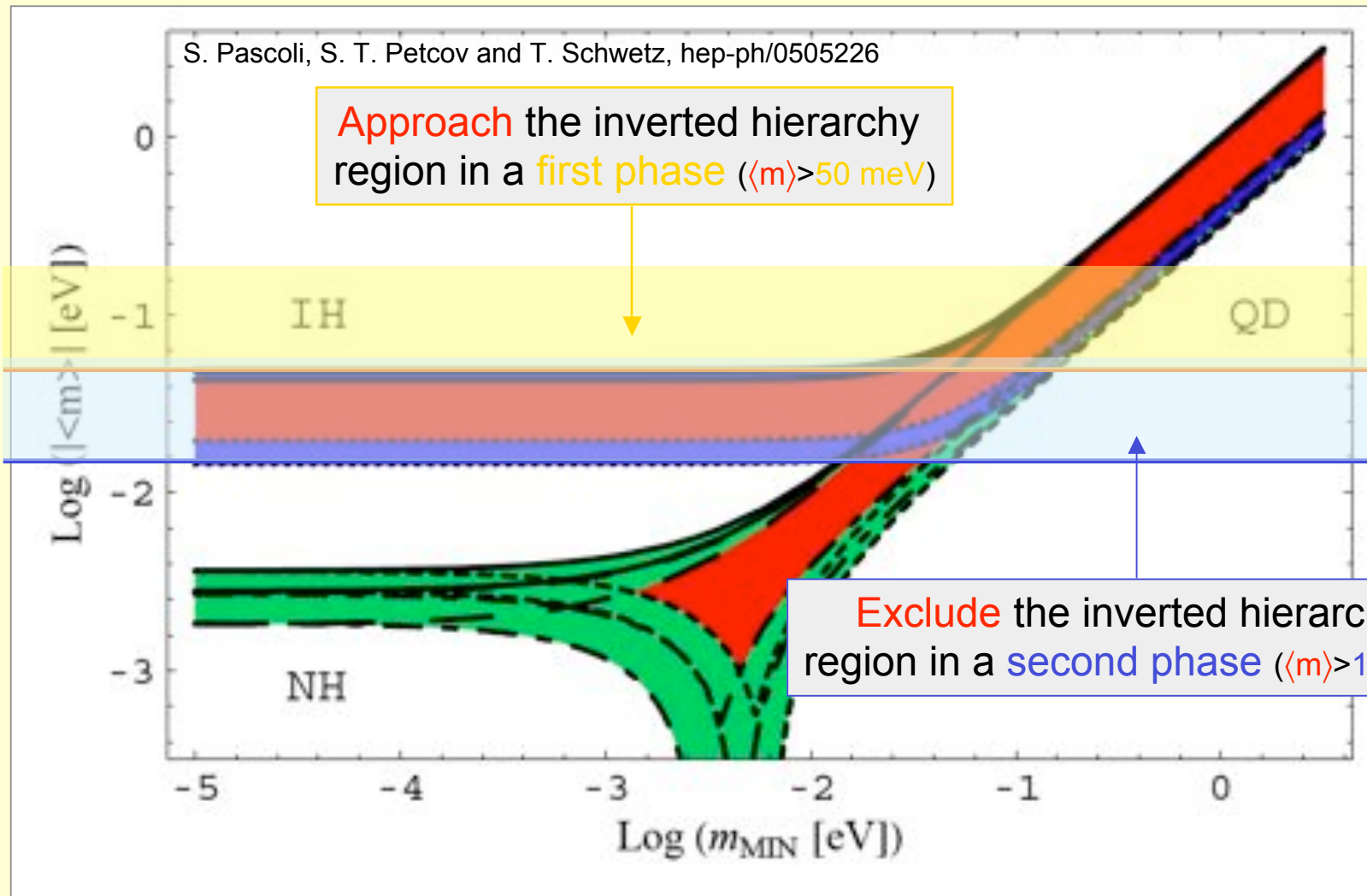


$$\langle M_{\beta\beta} \rangle = \left| |U_{e1}|^2 M_1 + e^{i\alpha_1} |U_{e2}|^2 M_2 + e^{i\alpha_2} |U_{e3}|^2 M_3 \right|$$

From where we start...

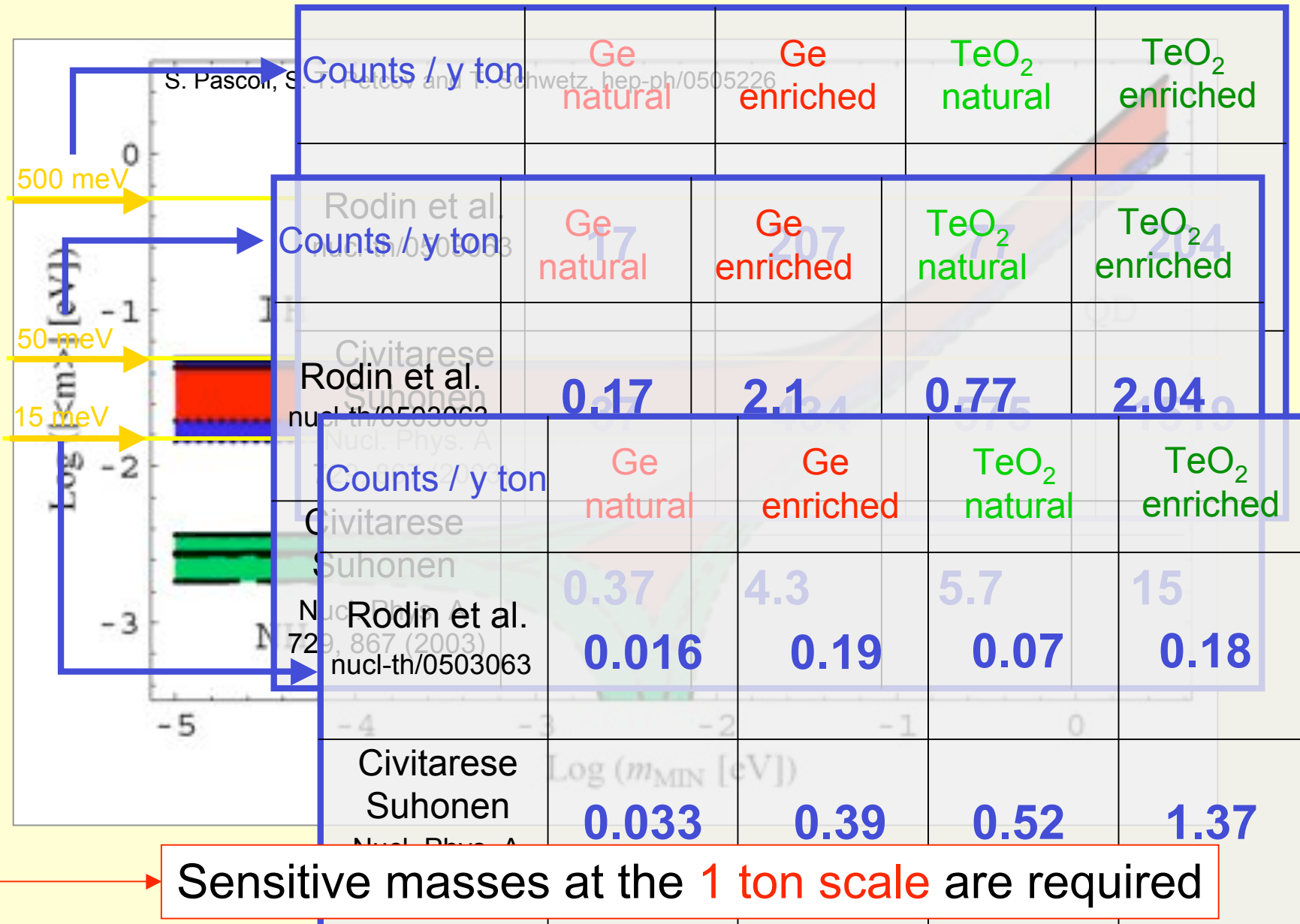


...and where we want to go



There are techniques and experiments in preparation which have the potential to reach these sensitivities

The size of the challenge



Background requirements

To start to explore the inverted hierarchy region

Sensitivity at the level of 1-10 counts / y ton

To cover the inverted hierarchy region

Sensitivity at the level of 0.1 -1 counts / y ton



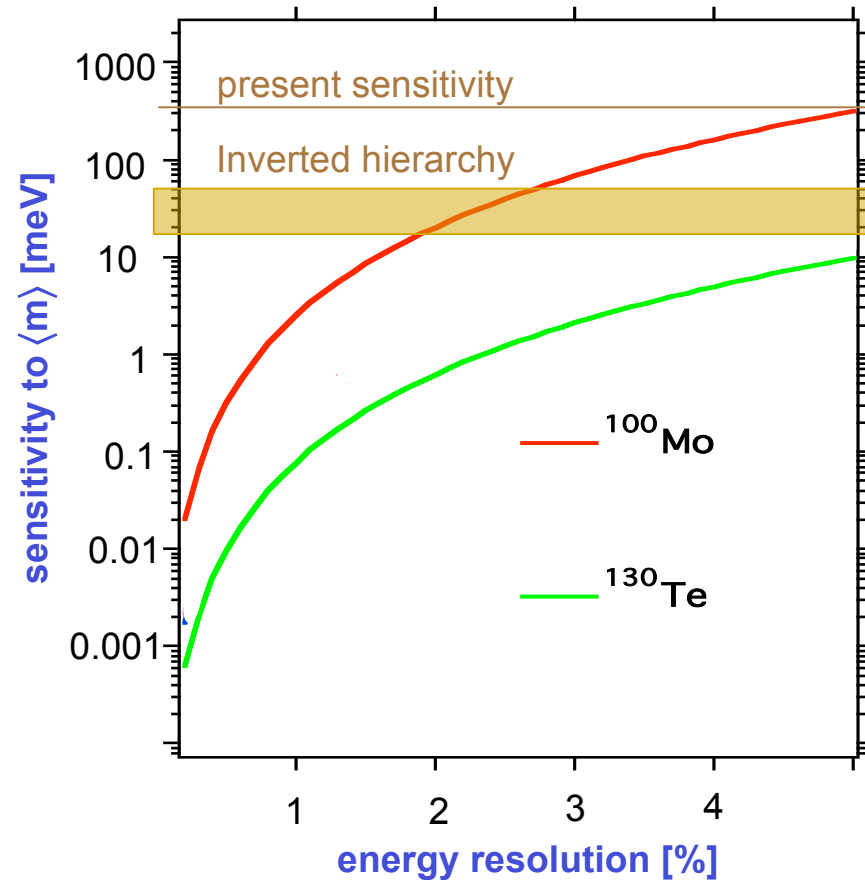
The order of magnitude of the target background is
~ 1 counts / y ton

A challenge for the space-resolving techniques,
which normally have **low energy resolution** ($\sim 10\%$)

- **Natural radioactivity**
(source itself)

- **Neutrons**

- **Cosmogenic**



- **2ν Double Beta Decay**

Experimental sensitivity to 0ν -DBD

sensitivity F: lifetime corresponding to the minimum detectable number of events over background at a given confidence level

$b \neq 0$

b: specific background coefficient
[counts/(keV kg y)]

$b = 0$

source mass live time energy resolution

$$F \propto (MT / b\Delta E)^{1/2}$$

$$F \propto MT$$

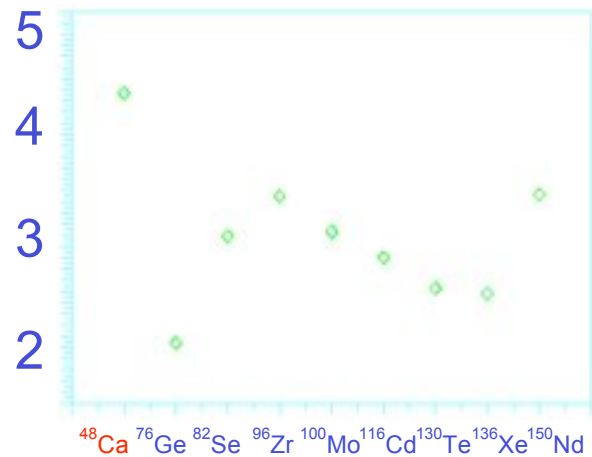
background level

importance of the **nuclide choice**
(but **large uncertainty** due to nuclear physics)

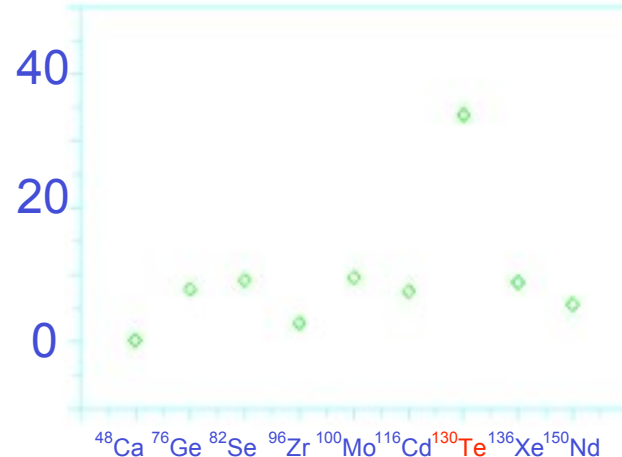
$$\text{sensitivity to } \langle M \rangle \propto (F/Q |M_{\text{nucl}}|^2)^{1/2} \propto \frac{1}{Q^{1/2} |M_{\text{nucl}}|} \left[\frac{b\Delta E}{MT} \right]^{1/4}$$

Choice of the nuclide

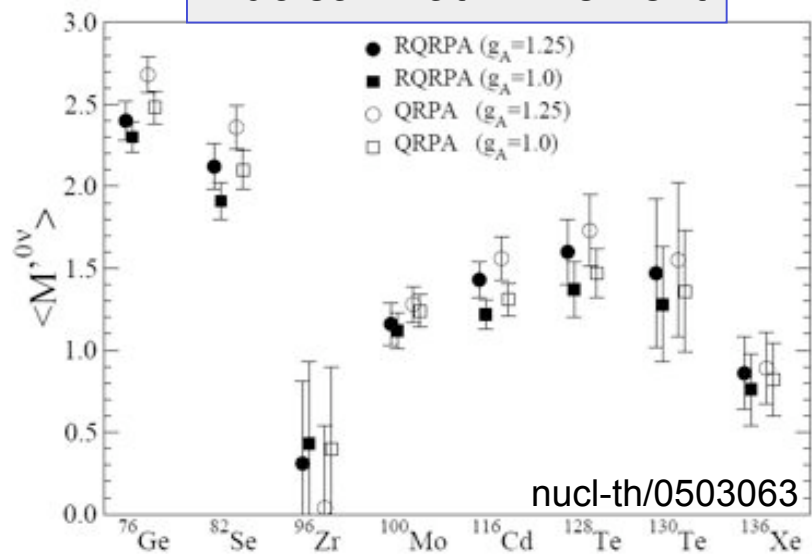
Transition energy (MeV)

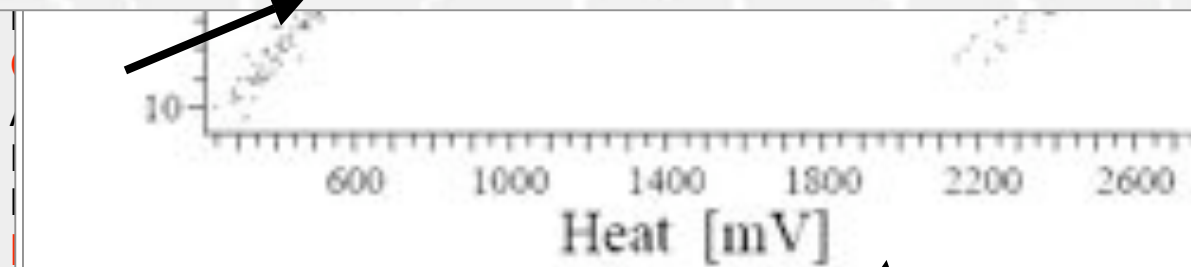
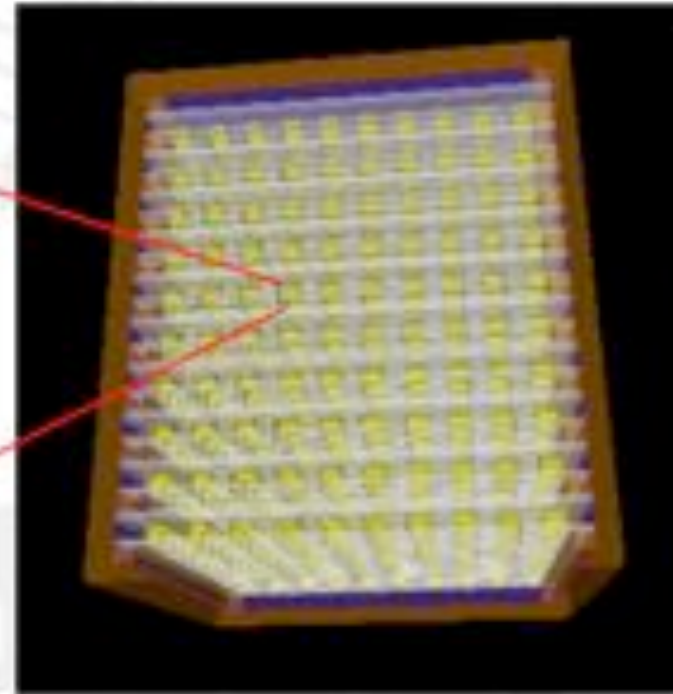
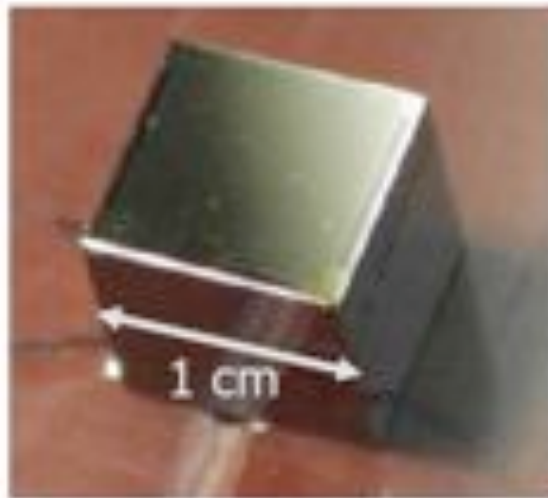


Isotopic abundance (%)



Nuclear Matrix Element





LOW
events

350 400

Granularity can be achieved through electrodes segmentation

⇒ R&D in progress for GERDA, MAJORANA, COBRA

SURFACE SENSITIVITY in bolometers

- R&D in progress in CUORE against energy-degraded α and β background

Simultaneous LIGHT and PHONON detection in bolometers

- R&D in progress in CUORE-like detectors for α / γ rejection

Small scale prototype at LNGS

Proved energy resolution: **1.9% FWHM**

Easy

2%)
ability
nd

Experiments and techniques



Source $\equiv$ Detector

Easy to approach the ton scale

CANDLES – ^{48}Ca

Array of natural pure (not Eu doped)

Prove of principle completed (CANDLES I)

Next step (CANDLES II: under construction)

Further step (CANDLES IV: requires 100 ton)

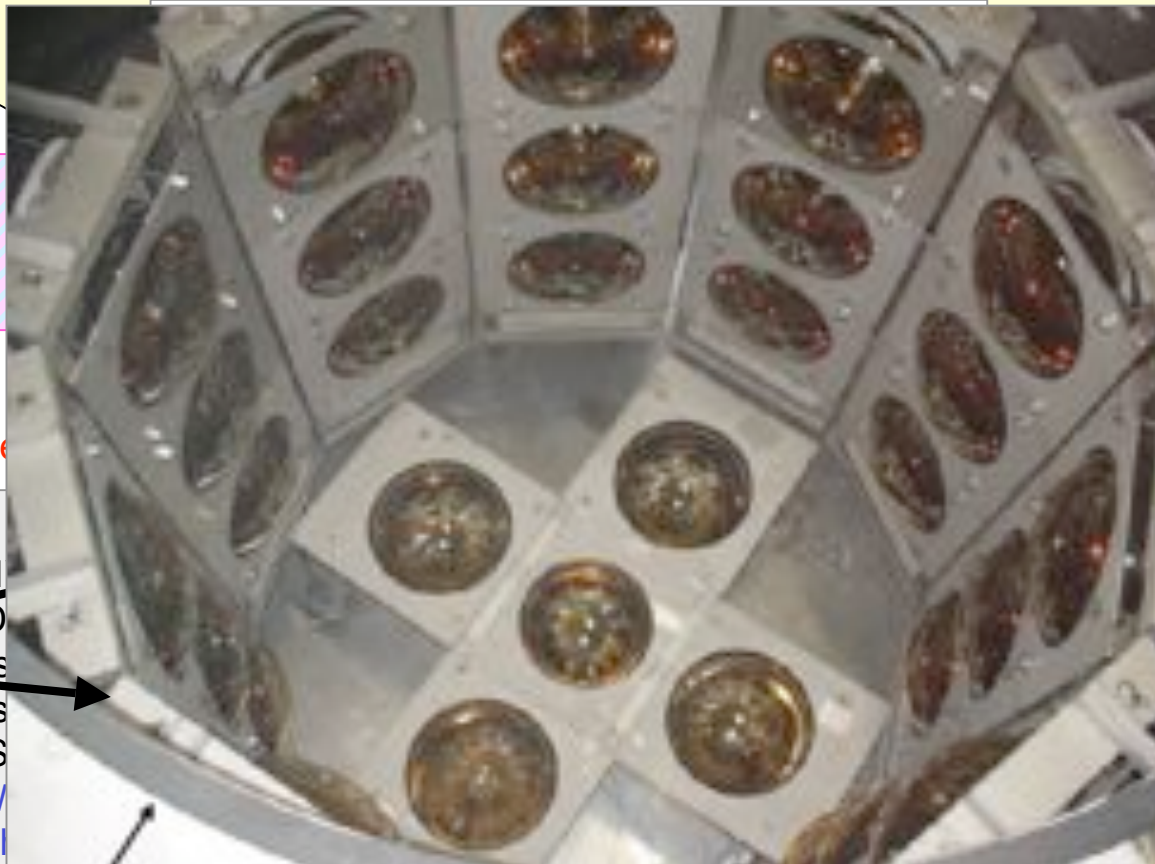
Final goal (CANDLES V): 100 ton (S)

Proved energy resolution: 3.4 % FWHM

The good point of this search is the low background

$\Rightarrow$ out of γ (2.6 MeV end point), β (3.3 MeV end point) and α (max 2.5 MeV with quench) natural radioactivity

Other background cuts come from PSD (α/β different timing) and space-time correlation for Bi-Po and Bi-Tl



Source $\neq$ Detector

Source $\neq$ Detector

Easy to get tracking capability

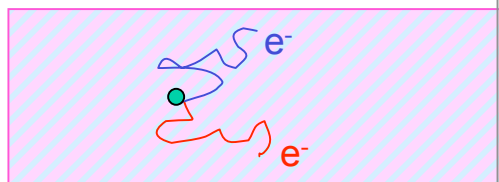
Low energy resolution (>2%)

Tracking / topology capability

Easy to approach zero background

(with the exception of 2ν DBD component)

Experiments and techniques



Source $\equiv$ Detector

Easy to approach the ton scale

EXO – ^{136}Xe

TPC of enriched liquid Xenon

Event position and topology: in prospe

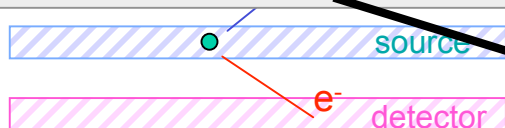
Next step (EXO-200: funded, under co

Further steps: 1-10 ton

Proved energy resolution: 3.3 % FWHM

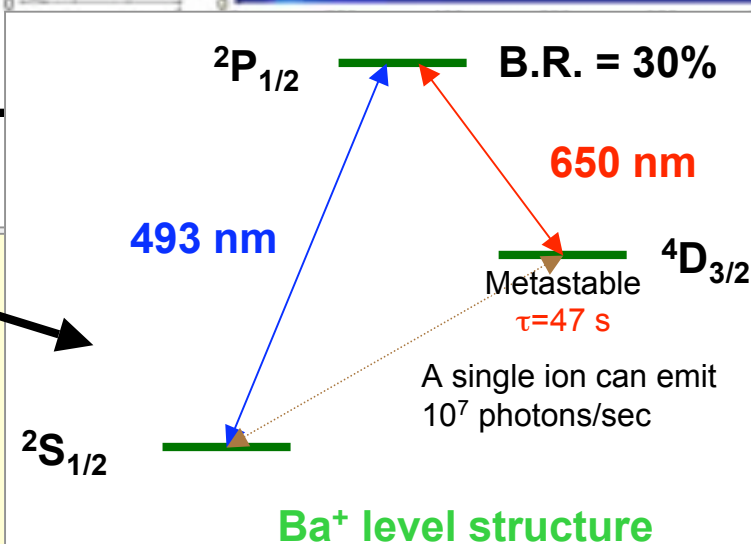
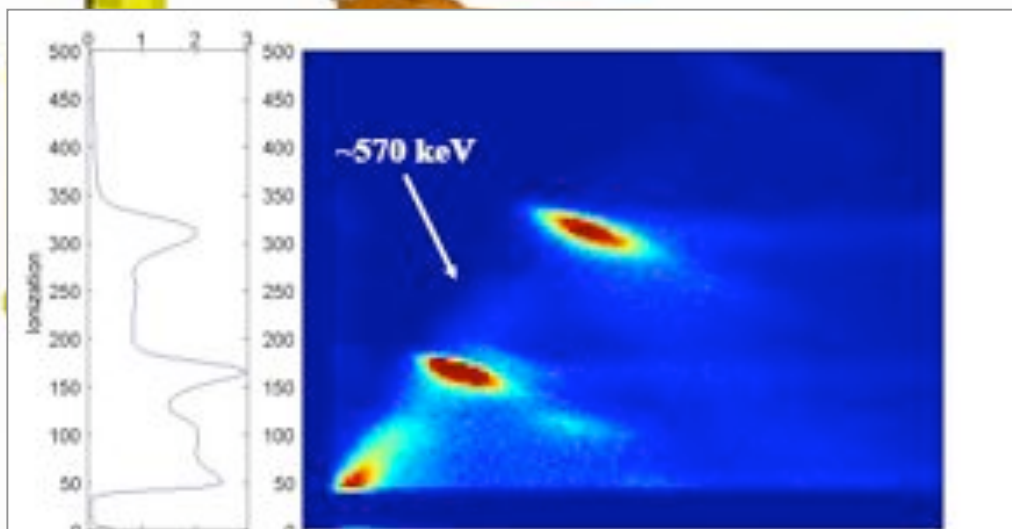
In parallel with the EXO-200 developm

Ba^{++} $e^- e^-$ final state is identified through opt



Source $\neq$ Detector

Easy to get tracking capability



on (>2%)
capability
ackground
of
nt)

Experiments and techniques

SUPERNEMO - ^{82}Se or ^{150}Nd

Modules with source foils, tracking (drift)
Magnetic field for charge sign

Possible configuration: 20 modules with

Energy resolution: 4 % FWHM

li can take advantage of NEMO3 experi

MOON - ^{100}Mo or ^{82}Se or ^{150}Nd

Multilayer plastic scintillators interleaved
MOON-1 prototype without tracking sec

MOON-2 prototype foreseen for 2006-2

Proved energy resolution: 6.8 % FWHM

Final target: collect 5 y x ton

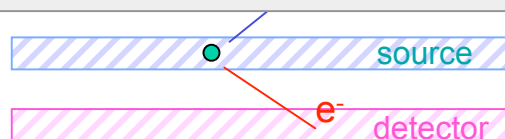
DCBA - ^{82}Se or ^{150}Nd

Momentum analyzer for beta particles

Prototype under construction: Nd_2O_3 fo

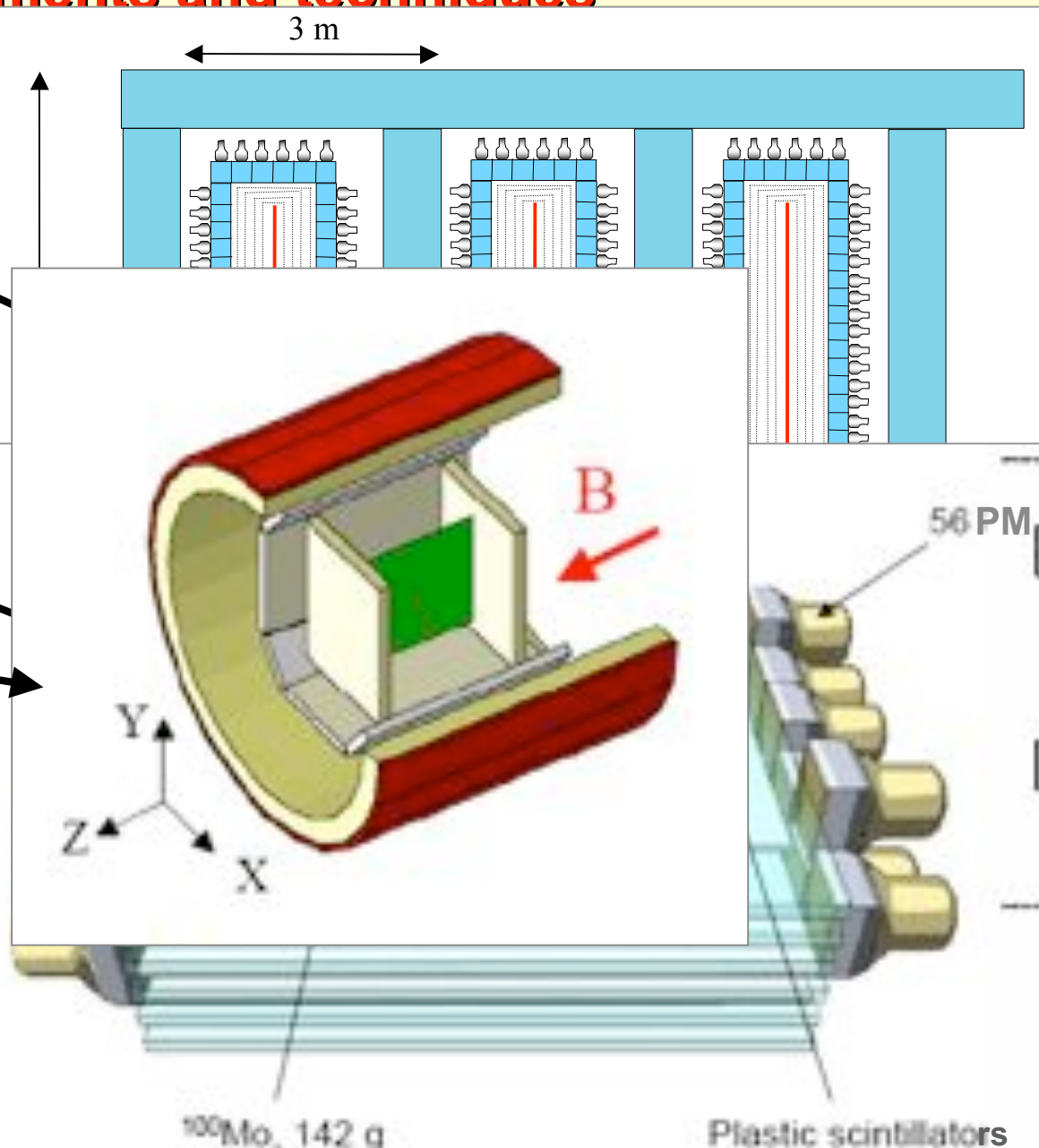
Space resolution ~ 0.5 mm; energy res

Final target: 10 modules with 84 m² so



Source $\neq$ Detector

Easy to get tracking capability



GERDA sensitivity

- Phase I: operate refurbished HM & IGEX enriched detectors (~20 kg)

Background: 0.01 counts/ keV kg y

Scrutinize ^{76}Ge claim with the same nuclide (exclude 99% c.l. or confirm 5σ)

Half life sensitivity: 3×10^{25} y

- Phase II: additional ~20 kg ^{76}Ge diodes (segmented detectors)

Background: 0.001 counts / keV kg y

Sensitivity after 100 kg y (~3 years): 2×10^{26} y

$\langle M \rangle < 90 - 290$ meV

- Phase III (depending on physics results of Phase I/II)

⇒ ~ 1 ton experiment in world wide collaboration with MAJORANA

$\langle M \rangle < 20 - 50$ meV

EXO 200 sensitivity

Assumptions on detector performance and background:

- 200 kg of Xe enriched to 80% in 136
- $\sigma(E)/E = 1.4\%$ obtained in EXO R&D
- Low but finite radioactive background: 20 events/year in the $\pm 2\sigma$ interval centered around the 2.481 MeV endpoint
- Negligible background from 2ν DBD ($T_{1/2} > 1 \cdot 10^{22}$ yr R. Bernabei et al. measurement)

Background: 40 counts in 2 y

Sensitivity (2 years, 90% c.l.): 6.4×10^{25} y

$$\langle M \rangle < 270 - 380 \text{ meV}$$

QRPA

NSM

If ^{76}Ge claim is correct :

- Worst case (QRPA, upper limit) 15 events (40 events bkg) $\Rightarrow 2\sigma$
- Best case (NSM, lower limit) 162 events (40 events bkg) $\Rightarrow 11\sigma$

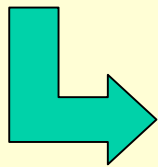
CUORE sensitivity

Montecarlo simulations of the background show that

$b = 0.001$ counts / (keV kg y)

is possible with the present bulk contamination of detector materials

The problem is the **surface background** (alpha, beta energy-degraded)



it must be reduced by a factor 10 – 100 with respect to Cuoricino work in progress! (only a factor 2 from the **conservative** assumption)

5 y sensitivity (1σ) with **conservative** assumption: $b = 0.01$ counts/(keV kg y)

$$F^{0\nu} = 9.2 \times 10^{25} \times (T [y])^{1/2}$$

$$\langle M \rangle < 20 - 100 \text{ meV}$$

5 y sensitivity (1σ) with **aggressive** assumption: $b = 0.001$ counts/(keV kg y)

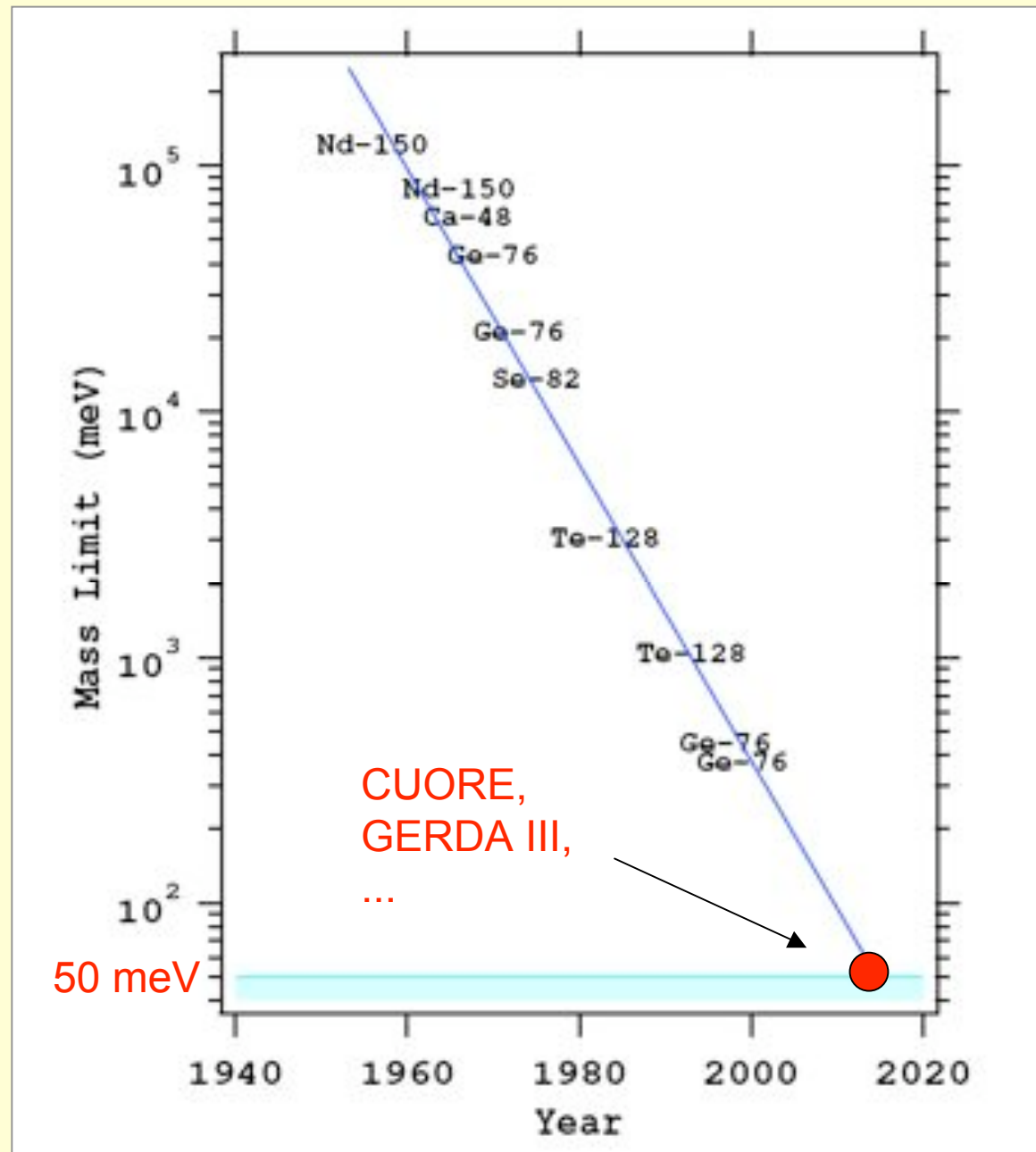
$$F^{0\nu} = 2.9 \times 10^{26} \times (T [y])^{1/2}$$

$$\langle M \rangle < 11 - 60 \text{ meV}$$

$$\langle M \rangle < 7 - 38 \text{ meV}$$

enriched CUORE

Prediction of the Moore's law for the sensitivity



Future scenarios and branching points in terms of discovery

sensitivity to

100 - 500

15 - 50

2 - 5

if this range holds:

- SUPERNEMO could marginally see it in ^{82}Se or ^{150}Nd
- GERDA phase III could see it in ^{76}Ge
- CUORE could see it in a **couple of isotopes** in sequence (after ^{130}Te , ^{116}Cd ?)

discovery in 3 or 4 isotopes necessary (and possible...)

if this range holds:

- **new strategies** have to be developed
- it is worthwhile to start to elaborate them now
- next generation experiments are precious for the selection of the future approaches
- a **large investment in enrichment** is mandatory

multi-isotope searches simultaneously
(^{130}Te - ^{116}Cd - ^{100}Mo)

large scale enrichment required
reduction of uncertainties in NME

precision measurement era for $0\nu\text{-DBD}$!

unpredictable time scale

EMO

estimate

^{82}Se or ^{150}Nd)