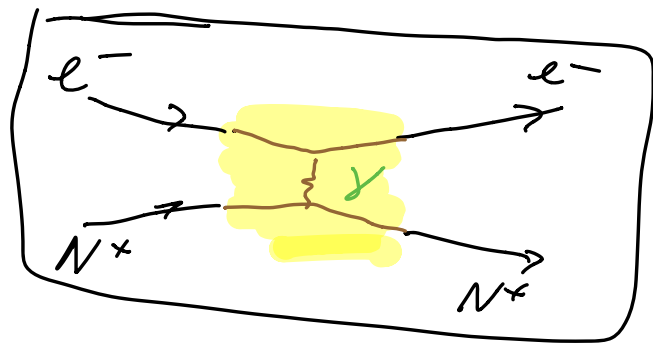




tempo
→



FORZA/INTERAZIONE	MEDIATORE
GRAVITAZIONALE	gravitone
EM.	photon
FORTE	gluoni (8)
DEBOLE	W^+, Z, W^-

ct
x
y
z

$$X^M = \begin{pmatrix} X^0 \\ X^1 \\ X^2 \\ X^3 \end{pmatrix} \quad [\quad] = [L]$$

CONTRAVARIANTE

$$= \begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix}$$

A · B

$$(X^M) \cdot (X^M) = X^0 X^0 - X^1 X^1 - X^2 X^2 - X^3 X^3$$

vettore covariante

$$X_\mu = \begin{pmatrix} ct \\ -x \\ -y \\ -z \end{pmatrix}$$

$$\sum_{\mu=0}^3 X^M \cdot X_\mu = X^0 X^0 - X^1 X^1 - \dots$$

covariance

$$\mu, \nu, \rho, \sigma \in \{0, 1, 2, 3\}$$

$$i, j, k, l, \dots \in \{1, 2, 3\}$$

$$\sum_{\mu=0}^3 X^M X_\mu \equiv X^M X_\mu$$

$$\vec{F} = \frac{d\vec{p}}{dt}$$

$$S = \int_0^{\Delta t} L dt$$

⊙ particelle libera

Galileo $L = T = \frac{1}{2} m v^2$

lorentz $L = ?$

S invariante di lorentz

$$S = \int L dt \quad d\tau = \frac{ds}{c}$$
$$= \int d\tau \cdot \underbrace{\alpha}_{\text{costante}} = \frac{dt}{\gamma}$$

$$\lim_{v \rightarrow 0} L = \lim_{v \rightarrow 0} \frac{\alpha}{\gamma}$$

$$L = \frac{\alpha}{\gamma} = \alpha \sqrt{1 - \frac{v^2}{c^2}} \approx \alpha \left(1 - \frac{1}{2} \frac{v^2}{c^2}\right)$$

$$= L_{\text{classica}} = \frac{1}{2} m v^2$$

a meno di una costante,

$$\frac{1}{2} m v^2 = - \frac{1}{2} \frac{v^2}{c^2} \alpha$$

$$\alpha = -m c^2$$

$$L = -m c^2 \sqrt{1 - \frac{v^2}{c^2}}$$

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{q}} = \frac{\partial L}{\partial q}$$

$$\begin{array}{l} \searrow \quad \nearrow \\ \{v_x, v_y, v_z\} \quad \{x, y, z\} \end{array}$$

$$L = L(q, \dot{q}, t) = L(x_i, \dot{x}_i, t)$$

$$= L(\dot{x}_i)$$

$$\frac{d}{dt} \left[\frac{\partial L}{\partial \dot{x}_i} \right] = \frac{\partial L}{\partial x_i} = 0 \rightarrow P_i \equiv \frac{\partial L}{\partial \dot{x}_i} = \text{const}$$

$$P_i = \frac{\partial L}{\partial \dot{x}_i} = \frac{\partial \left[-mc^2 \sqrt{1 - \frac{\dot{x}_i \dot{x}_i}{c^2}} \right]}{\partial \dot{x}_i}$$

$$= -mc^2 \frac{\partial \left(\sqrt{1 - \frac{\dot{x}_i \dot{x}_i}{c^2}} \right)}{\partial \dot{x}_i}$$

$$= -mc^2 \frac{1}{2} \left(1 - \frac{v^2}{c^2} \right)^{-1/2} \cdot \left(-\frac{2\dot{x}_i}{c^2} \right)$$

$$= m\dot{x}_i \gamma$$

$$\vec{p} = m\gamma \vec{v}$$

$$H \equiv \sum_{i=1}^3 p_i \dot{x}_i - L$$

$$= \sum \frac{\partial L}{\partial \dot{x}_i} \cdot \dot{x}_i - L$$

$$= \sum m\gamma v_i \cdot v_i - (-mc^2 \sqrt{1 - v^2/c^2})$$

$$\begin{aligned}
&= m \gamma v^2 + m c^2 / \gamma \\
&= m \gamma (v^2 + c^2 / \gamma^2) \\
&= m \gamma c^2 (v^2 / c^2 + 1 / \gamma^2) \\
&= m \gamma c^2 \\
&= H = E
\end{aligned}$$

$$\gamma^2 = \frac{1}{1 - \beta^2}$$

$$\beta = v/c$$

$$E = m \gamma c^2$$

$$P = |\vec{p}| = m \gamma v$$

$$p^\mu = \left(\frac{E}{c}, p^1, p^2, p^3 \right) \rightarrow p^{\mu'} = \dots$$

$$\begin{array}{ccc}
\updownarrow & \updownarrow & \updownarrow \\
x^\mu & ct & x^1 x^2 x^3
\end{array}$$

$$x^{\mu'} = \dots$$

Lorentz

INVARIANT: $p_\mu p^\mu = \frac{E}{c} \cdot \frac{E}{c} - p_x^2 - p_y^2 - p_z^2$

$$\begin{aligned}
&= \frac{E^2}{c^2} - p^2 \\
&= \frac{(m \gamma c^2)^2}{c^2} - (m \gamma v)^2 \\
&= m^2 \gamma^2 c^2 - m^2 \gamma^2 v^2
\end{aligned}$$

$$= m^2 \gamma_{c^2}^2 (1 - v^2/c^2)$$

$$= m^2 c^2$$

$$|P^A| \equiv \sqrt{P_\mu P^\mu} = mc$$

$$E^2 = m^2 c^4 + p^2 c^2$$

$$(E = \sqrt{m^2 c^4 + p^2 c^2})$$

$$\equiv \sqrt{m^2 + p^2}$$

$$c = 1$$

$$\hbar c = 1$$

NATURAL
UNITS

E in eV

$$1 \text{ eV} = 1.6 \cdot 10^{-19} \text{ C} \cdot \text{V}$$

$$= 1.6 \cdot 10^{-19} \text{ J}$$

$$m_p \cdot c^2 = 1.6 \cdot 10^{-27} \text{ Kg} \cdot (3 \cdot 10^8 \text{ m/s})^2$$

$$= 938 \cdot \text{MeV}$$

$$[m] = \text{MeV}/c^2$$

perché l'energia è conservata?
 l'impulso " " ?
 \vec{L} " " ?

E: traslazione temporale

P: traslazione spaziale

\vec{L} : rotazione

TEOREMA DI NOETHER

$$L = -mc^2 \sqrt{1 - \frac{\dot{x}_1^2 + \dot{x}_2^2 + \dot{x}_3^2}{c^2}}$$

invarianza spaziale

$$\frac{\partial L}{\partial x_i} = 0$$

$$\frac{d}{dt} \left[\frac{\partial L}{\partial \dot{x}_i} \right] = \frac{\partial L}{\partial x_i} = 0$$

$$P_i = \frac{\partial L}{\partial \dot{x}_i} = \text{costante}$$

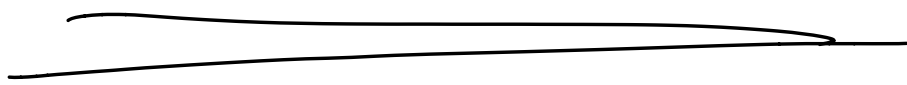
invarianza sotto trasl. temporali

$$\frac{\partial L}{\partial t} = 0$$

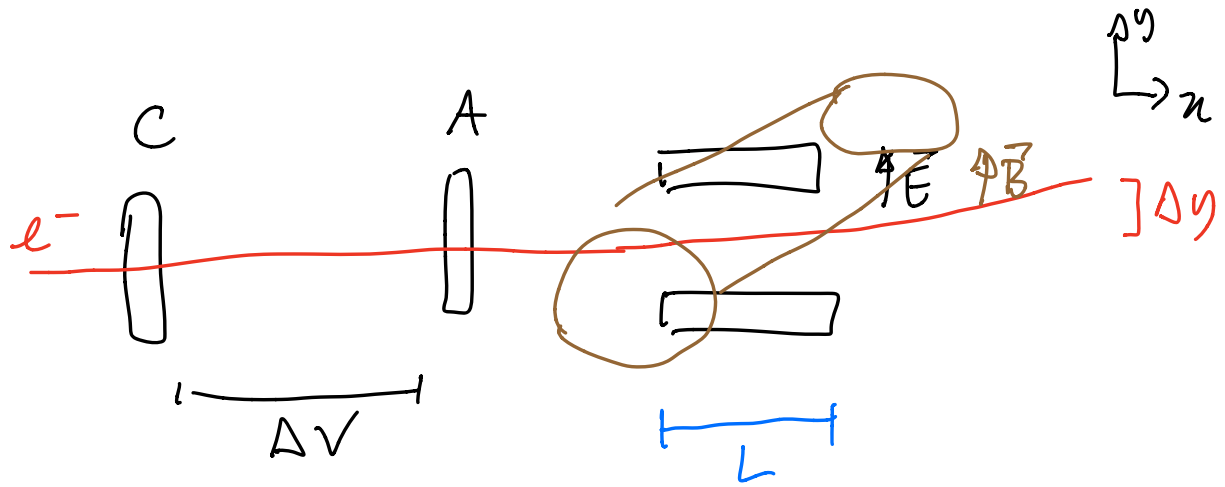
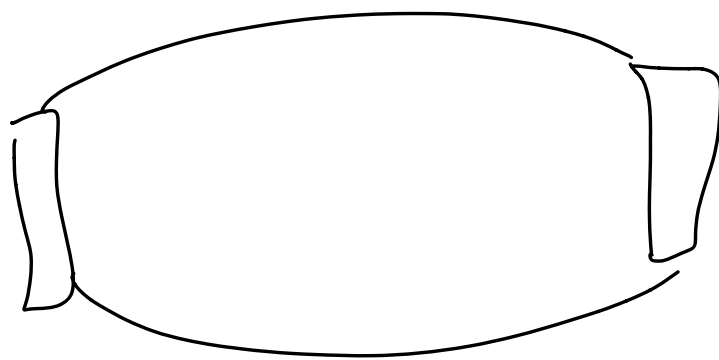
$$\begin{aligned} \frac{dL}{dt} &= \frac{\partial L}{\partial t} + \sum_{i=1}^3 \left[\frac{\partial L}{\partial x_i} \cdot \frac{\partial x_i}{\partial t} + \frac{\partial L}{\partial \dot{x}_i} \cdot \frac{\partial \dot{x}_i}{\partial t} \right] \\ &= \frac{\partial L}{\partial t} + \sum_{i=1}^3 \left[\left(\frac{d}{dt} \frac{\partial L}{\partial \dot{x}_i} \right) \cdot \frac{\partial x_i}{\partial t} + \frac{\partial L}{\partial \dot{x}_i} \cdot \frac{\partial \dot{x}_i}{\partial t} \right] \\ &= \frac{\partial L}{\partial t} + \frac{d}{dt} \left[\sum_{i=1}^3 \left[\frac{\partial L}{\partial \dot{x}_i} \cdot x_i \right] \right] \end{aligned}$$

$$\begin{aligned} - \frac{\partial L}{\partial t} &= \overset{\text{(inv.) sotto}}{\text{trasl. temporali}} 0 = - \frac{dL}{dt} + \frac{d}{dt} \sum_{i=1}^3 \left[\frac{\partial L}{\partial \dot{x}_i} \cdot \dot{x}_i \right] \\ &= - \frac{d}{dt} \left[L - \sum_{i=1}^3 \frac{\partial L}{\partial \dot{x}_i} \cdot \dot{x}_i \right] \\ &= - \frac{d}{dt} [-H] = \frac{d}{dt} [H] \end{aligned}$$

A OGNI SIMA, CONTINUA DI L
 COMISPOSURE UNA GRANDEZZA
 CONSERVATA



elettrone



$$\vec{F} = q\vec{E} = qE \hat{y}$$

$$\vec{F} = q\vec{v} \times \vec{B} = qvB \hat{y}$$

solos con \vec{E} :

$$\Delta y = \frac{1}{2} \frac{q}{m} \cdot E \cdot \frac{L^2}{v_x^2}$$

IMPONHO \vec{B} tal de

$$qE = qv_x B \rightarrow v_x$$
$$\rightarrow \frac{q}{m}$$