

From QCD to Quark-Gluon Plasma

Pasquale Di Nezza&Simonetta Gentile

Pasquale.DiNezza@Lnf.infn.it

Università Sapienza – Nuclei e Particelle, a.a. 2018-2019

History of the Universe



Big Bang



Today

Source: Nuclear Science Wall Chart

History of the Universe



Big Bang



Neutrons 1012K, 10-4s Low-mass Nuclei 109K, 3 min



Neutral Atoms 4000K, 105y Star Formation 109y Heavy Elements >10⁹y Today

Source: Nuclear Science Wall Chart

History of the Universe



Neutral Atoms 4000K, 105y

0

151

Star Formation 109y Heavy Elements >10⁹y Today

Source: Nuclear Science Wall Chart



Energy Scales

The beginning The universe is a hot plasma of fundamental particles ... quarks, leptons, force mediating particles (and other particles ?) 10^{-43} s Planck scale (quantum gravity ?) $10^{19} \, {\rm GeV}$ Grand unification scale (strong+electroweak) 10^{-35} s $10^{15} \,\mathrm{GeV}$ Inflationary period 10⁻³⁵-10⁻³³ s 10⁻¹¹ s Electroweak unification scale 200 GeV Micro-structure 10⁻⁵ s QCD scale - protons and neutrons form 200 MeV 3 mins Primordial nucleosynthesis 5 MeV 3×10^5 y Radiation and matter decouple - atoms form $1 \,\mathrm{eV}$

• Large scale structure

- 1 b yrs Proto-galaxies and the first stars
- 3 b yrs Quasars and galaxy spheroids
- 5 b yrs Galaxy disks
- Today Life !

Energy Scales

• The beginning

The universe is a hot plasma of fundamental particles ... quarks, leptons, force mediating particles (and other particles ?)

200 GeV

10-43 sPlanck scale (quantum gravity ?)1019 GeV10-35 sGrand unification scale (strong+electroweak)1015 GeVInflationary period 10-35-10-33 s

10⁻¹¹ s Electroweak unification scale

• Micro-structure

10 ⁻⁵ s	QCD scale - protons and neutrons form	200 MeV
3 mins	Primordial nucleosynthesis	5 MeV
3×10 ⁵ y	Radiation and matter decouple - atoms form	1 eV

• Large scale structure

- 1 b yrs Proto-galaxies and the first stars
- 3 b yrs Quasars and galaxy spheroids
- 5 b yrs Galaxy disks
- Today Life !

Standard Model



arXiv:hep-ph/9505231

arXiv:hep-ph/9505231



Strong Interaction:

- binds quarks into hadrons
- binds hadrons into nuclei

QCD describes:

- quark-gluon interactions
- gluon-gluon interactions

```
arXiv:hep-ph/9505231
```

Very successful theory valid over 25 orders of magnitude and up to the TeV scale!

Comparison pQCD with hadron jet production cross section



```
arXiv:hep-ph/9505231
```

Very successful theory valid over 25 orders of magnitude and up to the TeV scale!



The hadron mass

 A proton is composed out of uud quarks



- The proton mass is 938.3 MeV/c²
- The sum of bare quark masses is only ~10 MeV/c²

How is the extra-mass generated?



Confinement

- An isolated quark has never been observed
- The quarks seem confined within the hadrons
- Half of the fundamental fermions is not observable directly

Why?

(SU(3)_c invariant)

$$\mathcal{L}_{\text{QCD}} \equiv -\frac{1}{4} G^{\mu\nu}_a G^a_{\mu\nu} + \sum_f \bar{q}_f \left(i\gamma^\mu D_\mu - m_f \right) q_f$$

Structure: QED-like (generalised Maxwell (Yang-Mills) + Dirac)

(SU(3)_C invariant)

$$\mathcal{L}_{\text{QCD}} \equiv -\frac{1}{4} G^{\mu\nu}_{a} G^{a}_{\mu\nu} + \sum_{f} \bar{q}_{f} \left(i\gamma^{\mu} D_{\mu} - m_{f} \right) q_{f}$$
Elementary quark field
$$\begin{pmatrix} \text{color} & a = 1, \dots, 3 \\ \text{spin} & \alpha = 1, 2 \\ \text{flavor} & f = u, d, s, c, b, t \end{pmatrix}$$

Structure: QED-like (generalised Maxwell (Yang-Mills) + Dirac)

(SU(3)_C invariant)

$$\mathcal{L}_{\text{QCD}} \equiv -\frac{1}{4} G_{a}^{\mu\nu} G_{\mu\nu}^{a} + \sum_{f} \bar{q}_{f} \left(i\gamma^{\mu} D_{\mu} - m_{f} \right) q_{f}$$
Elementary quark field
$$(q_{\alpha})_{f}^{a} \begin{cases} \text{color} \quad a = 1, \dots, 3 \\ \text{spin} \quad \alpha = 1, 2 \\ \text{flavor} \quad f = u, d, s, c, b, t \end{cases}$$
Tensorial part
$$G_{\mu\nu}^{a} = \partial_{\mu} A_{\nu}^{a} - \partial_{\nu} A_{\mu}^{a} + g f^{abc} A_{\mu}^{b} A_{\nu}^{c} \end{pmatrix}$$

$$i \not D q = \gamma^{\mu} \left(i \partial_{\mu} + g A_{\mu}^{a} t^{a} \right) q$$
Elementary gluon field
$$\left(A_{\mu}^{a} \right) \begin{cases} \text{color} \quad a = 1, \dots, 8 \\ \text{spin} \quad \epsilon_{\mu}^{\pm} \end{cases}$$

Contrary to photons, gluons carry (color) charge and can interact among themselves creating complicated structures.

(SU(3)_c invariant)

$$\mathcal{L}_{\text{QCD}} \equiv -\frac{1}{4} G_{a}^{\mu\nu} G_{\mu\nu}^{a} + \sum_{f} \bar{q}_{f} (i\gamma^{\mu}D_{\mu} - m_{f}) q_{f}$$
Elementary quark field
$$(q_{\alpha})_{f}^{a} \begin{cases} \text{color} \quad a = 1, \dots, 3 \\ \text{spin} \quad \alpha = 1, 2 \\ \text{flavor} \quad f = u, d, s, c, b, t \end{cases}$$
Tensorial part
$$G_{\mu\nu}^{a} = \partial_{\mu}A_{\nu}^{a} - \partial_{\nu}A_{\mu}^{a} + gf^{ab}A_{\nu}^{b}A_{\nu}^{c}$$

$$i \not D q = \gamma^{\mu} (i\partial_{\mu} + gA_{\mu}^{a}t^{a}) q$$
Elementary gluon field
$$A_{\mu}^{a} \begin{cases} \text{color} \quad a = 1, \dots, 8 \\ \text{spin} \quad \epsilon_{\mu}^{\pm} \end{cases}$$
Ex. gluon-gluon interaction

Running coupling

Consider the interaction of 2 elementary particles as a function of Q²



Because of Heisenberg U.P.: small $Q^2 \rightarrow$ large distances large $Q^2 \rightarrow$ small distances

Virtual pairs screen the bare interaction resulting in momentum-transfer dependent interaction strength $\rightarrow \alpha(Q)$

Running coupling: α vs α_s



Running coupling: α vs α_s

QED
$$\alpha(Q^2) \approx \frac{\alpha(\mu^2)}{1 - \frac{1}{3\pi} \alpha(\mu^2) \log \frac{|Q^2|}{\mu^2}}$$

Negative
Small Q² (large distances) \rightarrow weaker α
(similar to screening of charge in dielectric materials)
QCD $\alpha(Q^2) \approx \frac{\alpha(\mu^2)}{1 + \frac{11N_{color} - 2n_{flavor}}{12\pi} \alpha(\mu^2) \log \frac{|Q^2|}{\mu^2}}{\alpha(\mu^2) \log \frac{|Q^2|}{\mu^2}}$
 $(33-12)/12\pi \rightarrow Positive$
Small Q² (large distances) \rightarrow stronger α
(anti-screening larger than screening)









Confinement and Asymptotic freedom: a toy model

Let's parameterise the increase of the potential for a $q\overline{q}$ pair by a potential "a la Cornell" + linear term (flux tube) which considers the confinement (semi-classic, non relativistic)



- potential vanishes for large distance
- $\circ \quad \begin{array}{l} \text{deconfinement of quarks} \\ \rightarrow \text{QGP} \end{array}$

in vacuum:

- linear increase with distance, strong attractive force
- \circ confinement of quarks to hadrons

Confinement and Asymptotic freedom: a toy model

Let's parameterise the increase of the potential for a $q\overline{q}$ pair by a potential "a la Cornell" + linear term (flux tube) which considers the confinement (semi-classic, non relativistic)

Coulomb potential

$$V(r) = -\frac{\alpha}{r} + kr \quad Confinement$$



Increasing "r" it become energetically favorable to convert the stored energy into a new $q\overline{q}$ pair

Confinement and Asymptotic freedom: a toy model

Let's parameterise the increase of the potential for a $q\overline{q}$ pair by a potential "a la Cornell" + linear term (flux tube) which considers the confinement (semi-classic, non relativistic)

Coulomb potential

$$V(r) = -\frac{\alpha}{r} + kr$$
Confinement

(colorless)

Increasing "r" it become energetically favorable to convert the stored energy into a new $q\overline{q}$ pair

The confinement cannot be described perturbatively. At scales of the hadron size (~1fm) the perturbative methods lose validity. Calculations rely on approximate methods (lattice theory, effective theories). Ex: MIT Bag Model, simple QCD inspired model

Superdense Matter: Neutrons or Asymptotically Free Quarks?

J. C. Collins and M. J. Perry

Department of Applied Mathematics and Theoretical Physics, University of Cambridge, Cambridge CB3 9EW, England (Received 6 January 1975)

We note the following: The quark model implies that superdense matter (found in neutron-star cores, exploding black holes, and the early big-bang universe) consists of quarks rather than of hadrons. Bjorken scaling implies that the quarks interact weakly. An asymptotically free gauge theory allows realistic calculations taking full account of strong interactions.

A neutron has a radius¹⁰ of about 0.5-1 fm, and so has a density of about 8×10^{14} g cm⁻³, whereas the central density of a neutron star^{1,2} can be as much as $10^{16}-10^{17}$ g cm⁻³. In this case, one must expect the hadrons to overlap, and their individuality to be confused. Therefore, we suggest that matter at such high densities is a quark soup. How does QCD matter behave under extreme conditions of temperature and energy density?

A question rooted in the QCD, with cosmological and astrophysical implications

MIT Bag Model - (confinement)

The Model assumes that the quarks are confined within bags of perturbative (empty) vacuum of radius R, in which they are free to move

The QCD vacuum creates a confining bag with pressure B

The bag constant is obtained by balancing the vacuum with the kinetic pressure of the quarks.

By minimizing:

$$E \approx \frac{2N}{R} + \frac{4}{3}\pi R^3 B$$

With N=3 (quarks) and R=0.8 fm

 $B \approx (200 \text{ MeV})^4 = 0.2 \text{ GeV/fm}^3$



At the end, 0.2 GeV/ fm³ are sufficient to confine 3 quarks within the proton volume

pressure = **B**

MIT Bag Model - (confinement)

The Model assumes that the quarks are confined within bags of perturbative (empty) vacuum of radius R, in which they are free to move

The QCD vacuum creates a confining bag with pressure B

The bag constant is obtained by balancing the vacuum with the kinetic pressure of the quarks. By minimizing:

With N=3 (quart

Th N=3 (quere sure exceeds bag pressure $3\pi R^3 B$ If kinetic pressure exceeds fm B \approx (200)

A toy model - (deconfinement)

We can heat matter so much that individual hadrons start to overlap

From statistical mechanics, for an ideal gas

$$p = \frac{\epsilon}{3} = \left(g_B + \frac{7}{8}g_F\right)\frac{\pi^2 T^4}{90}$$



A toy model - (deconfinement)

We can heat matter so much that individual hadrons start to overlap

From statistical mechanics, for an ideal gas ... let's add the compression too

$$p = \frac{\epsilon}{3} = \left(g_B + \frac{7}{8}g_F\right)\frac{\pi^2 T^4}{90} + g_F\left(\frac{\mu_F^2 T^2}{12} + \frac{\mu_F^4}{24\pi^2}\right)$$



A toy model - (deconfinement)

We can heat matter so much that individual hadrons start to overlap

From statistical mechanics, for an ideal gas ... let's add the compression too

.. let's add the compression too

$$p = \frac{\epsilon}{3} = \left(g_B + \frac{7}{8}g_F\right)\frac{\pi^2 T^4}{90} + \frac{g_F}{\left(\frac{\mu_F^2 T^2}{12} + \frac{\mu_F^4}{24\pi^2}\right)} + \frac{\mu_F^4}{Phase transition} \qquad \text{QGP}$$

If the pression>B and/or T>T_c we have the conditions for Quark Gluon Plasma (QGP)

 $g_B = 2x8$ (spin x colors) = 16 $g_F = 2x2x3x3$ ($q\overline{q}$ x spin x flavor x colors) = 24

 $g_{\rm B}=0, g_{\rm F}=2$

Compression

P •

π

A phase transition has brought the system^{*} to a deconfined stage with release of degrees of freedom

*of a non negligible dimensions

Matter in extreme conditions

How does matter behave in such extreme conditions?

What are the properties of the Quark-Gluon Plasma and the early Universe? Remember that even with the most powerful telescopes, we cannot go back in time to less than ~ 400,000 years after the Big Bang (except GW)

Matter in extreme conditions

How does matter behave in such extreme conditions?

What are the properties of the Quark-Gluon Plasma and the early Universe? Remember that even with the most powerful telescopes, we cannot go back in time to less than ~ 400,000 years after the Big Bang (except GW)

Fermi Notes on Thermodynamics 35 QGP 12 Electron proton gas 10 Non deg. 8 6 4 2 18 32 2 26 28 30 'n 10 127 14 20 V.Greco





Fig. 1. Schematic phase diagram of hadronic matter. ρ_B is the density of baryonic number. Quarks are confined in phase I and unconfined in phase II.

Experimental hadronic spectrum and quark liberation Cabibbo and Parisi, PLB59 (1975) 67
QCD phase diagram



QCD phase diagram



First order phase transition (lattice calculation) Including quark masses (not at the first order)

QCD phase diagram



Several different phases found (present status)

 T_{room} ~300 K ~ 25 meV (milli-eV !)

 $n_c^B = 0.72 \text{fm}^{-3}$ (net-baryon density of about 5 x nucleus)

The confinement cannot be described perturbatively. At scales of the hadron size (~1fm) the perturbative methods lose validity.

QCD can be solved numerically by putting fields on a space-time lattice. It is a rigorous way of doing calculations in non-perturbative regime of QCD.

Computationally demanding: farm with 300.000 cores, petaFLOPS

The confinement cannot be described perturbatively. At scales of the hadron size (~1fm) the perturbative methods lose validity.



Snapshot of fluctuating quark and gluon fields on a discrete space-time lattice

Fluctuating quark and gluon fields on a discrete space-time lattice







We discretize the space-time and, on this lattice, we solve the QCD equations



We discretize the space-time and, on this lattice, we solve the QCD equations

In order to go back from the lattice to the real physics we have to apply the:

- Continuum limit a ->0, infinite momenta
- Infinite volume limit $\lor \rightarrow \infty$
- Set scales using data (e.g. hadron masses)

Problems of approach:

- 2xfermions then the real world
- Small masses ask huge CPU time, large masses are needed.
- Very difficult for finite $\boldsymbol{\mu}$



Excellent agreement between Lattice (2 flavors) and experimental data

Temperature dependence of the heavy quark free energy (static potential) in 3-flavour QCD



Increasing T there is the creation of spontaneous qqbar-pairs in the "heat bath" → exhibits screening of long range confining potential with increasing temperature "Quasi free interaction"



Temporal extension (space-time at large volume)



 $(T_c \sim 170 \text{ MeV}, \varepsilon_c \sim 1 \text{ GeV/fm}^3)$



Chiral symmetry





proton neutron

source: http://de.wikipedia.org



 $2 m_{\rm u} + m_{\rm d} = 9.6 \text{ MeV/}c^2$ $m_{\rm proton} = 938.27 \text{ MeV/}c^2 !!!$

positive pion

- Hadron mass scale set by constituent quarks masses ($m_{u,d,const} \approx 300 \text{ MeV}/c^2$)
- QCD responsible for 99% of the mass of your body!
- Related to breaking of chiral symmetry

Chiral symmetry



In the absence of quark mass the QCD Lagrangian splits into two independent quark terms:

$$\mathcal{L}_{\rm QCD} = \mathcal{L}_{\rm gluons} + i\bar{q}_L\gamma^\mu D_\mu q_L + i\bar{q}_R\gamma^\mu D_\mu q_R$$

For two flavors (i=u,d) the Lagrangian is symmetric under $SU(2)_L \times SU(2)_R$

Symmetry NOT observed \rightarrow solution: the vacuum is not invariant. The "empty" vacuum is unstable. There is a state of lower energy that consists of cells, each containing a gluon pair \rightarrow "Liquid" vacuum

 $\langle 0 | \bar{q}_L q_R | 0 \rangle \neq 0$ Chiral condensate

Chiral symmetry: Fermions and anti-fermions have opposite helicity

Spontaneous symmetry breaking (pseudo-goldstone bosons: pions)



Restoration of bare masses

Quarks have very small masses generated by the coupling to Higgs (light q< 10 MeV)

Confined quarks (i.e. in the proton) require ~350 MeV generated dynamically through the confining effects of the strong interaction

Deconfinement must be accompanied by a restoration of the masses to the bare mass values they have in the Lagrangian:

- m(u,d): ~350 MeV → few MeV
- m(s): ~500 MeV → 150 MeV



*Partial because the symmetry is exact only for massless particles, therefore its restoration here is only partial

Satz, arXiv:0803.1611

QCD , a successful theory with some fundamental problem

Is there a regime were the symmetry is restored?

QCD phase transition

Where?

At the Big Bang

we think that in the first instants of life of the Universe, quarks and gluons were not trapped inside hadrons (protons, neutrons, ...) but could move freely in a "deconfined" state: the Quark-Gluon Plasma

10 µs: the birth of hadrons

after about 10 μ s from the Big Bang, the Universe cooled down to less than 2 x 10¹² degrees

at that point, the QCD phase transition took place: quarks and gluons were confined inside hadrons

the familiar particles, such as pions, kaons, protons and neutrons appeared on the stage of the Universe



... and -in the core of the neutron stars -in the heavy-ion collision experiments

Phase transitions of the Universe

The early Universe (Kolb, Turner) Schwarz, astro-ph/0303574



Fig. 3.5: The evolution of $g_*(T)$ as a function of temperature in the $SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$ theory.



T.D.Lee, Rev.Mod.Phys. 47 (1975) 267

In high energy physics we have concentrated on experiments, in which we distribute a higher and higher amount of energy into a region with smaller and smaller dimensions. In order to study the question of "vacuum", we must turn to a different direction; we should investigate some "bulk" phenomena by distributing high energy over a relatively large volume.

Nucleus-Nucleus Collisions

- We need a small system so that it can be accelerated to ultrarelativistic speed (99.9% c)
- That system (i.e. a chunk of matter and not just a single particle) must follow simple rules of thermodynamics and form a new state of matter in a particular phase
- We can use heavy ions (e.g. Pb). They are tiny (~10⁻¹⁴ m) and have a finite volume that can be exposed to pressure and temperature (the system is more than 1 order of magnitude larger than the pp)

Nucleus-Nucleus Collisions



A strong and critic difference is the time scale evolution of the system

We need Heavy-Ions



Colliding Heavy Ions



Lorentz-contracted nuclei (Δz~R/γ) Hard Collisions

Parton Dynamics Hadron Dynamics

pQCD

QCD Matter quarks and gluons are the relevant degrees of freedom Mesons and Baryons are always the final degrees of freedom

Simulation "VNI" (Geiger, Longacre, Srivastava)





Run:244918 Timestamp:2015-11-25 11:25:36(UTC) System: Pb-Pb Energy: 5.02 TeV



CMS Experiment at LHC, CERN Data recorded: Wed Nov 25 12:21:51 2015 CET Run/Event: 282548 / 14582169 Lumi section: 309



Run: 286665 Event: 419161 2015-11-25 11:12:50 CEST

first stable beams heavy-ion collisions

ATLAS



Event 2598326 Run 168486 Wed, 25 Nov 2015 12:51:53





final detected



Soft processes:

- High cross section
- Decouple late
- \rightarrow Indirect signals for QGP

<u>EM probes (real and</u> virtual photons): insensitive to the hadronization phase Hard processes:

Low cross section

• Probe the whole evolution of the collision



Various observables will probe different stages of the collision



Hard probes (jets, heavy flavor, EW bosons)

Transverse flow Thermal photons Multiplicity, HBT Particle yields + spectra

Atom not ατομοσ





Ernest Rutherford (r) and Hans Geiger (I) in Manchester





Atom not ατομοσ



... going deeper





Jerome I. Friedman Prize share: 1/3



Photo: T. Nakashima Richard E. Taylor

The discovery of quarks (still point-like objects) 1990 Nobel Prize

Same idea for exploring the QGP?

"Calibrated probes" can be used to probe the QGP

The idea is to measure how QGP can modify the probes?



Same idea for exploring the QGP?

"Calibrated probes" can be used to probe the QGP

The idea is to measure how QGP can modify the probes?



The path to the Heavy lons LHC

Brookhaven National Laboratory (BNL)

- AGS (1986-2000) Si and Au beams, \sqrt{s} ~5 GeV (only hadronic variables)
- RHIC (2000-...) ³He, Cu, Au beams, up to $\sqrt{s}=200$ GeV (4 experiments, 2 left)



• CERN-LHC (2009-...) Pb beams, √s~5000 GeV (ALICE, ATLAS, CMS, LHCb)
RHIC @ BNL



Nucleus-Nucleus Collisions at the LHC

Fully ionised ²⁰⁸Pb nucleus accelerated in the LHC (configuration magnetically identical to that for pp)

$$p_{Pb} = Zp_p = 82 \cdot 6.5 = 533TeV$$
$$\sqrt{s_{PbPb}} = 1066TeV$$

The relevant figure is √s per nucleon-nucleon collision (latest configuration):

$$\sqrt{s_{NN}} = \frac{2E_{Pb}}{A} = \frac{Z}{A}\sqrt{s_{pp}} = \frac{82}{208}\sqrt{s_{pp}} = 5.1TeV$$

		SPS	RHIC	LHC
√s _{NN}	[GeV]	17.3	200	5500
dN _{ch} /dy		450	800	1600
3	[GeV/fm ³]	3	5.5	~ 10



Nucleus-nucleus collisions at the LHC



Nucleus-Nucleus Collisions at the LHC



ALICE (A Large Ion Collider Experiment) HI dedicated experiment: -Low-p_T tracking, PID, mid-rapidity + forward muons

ATLAS and **CMS**, multipurpose experiments. Large capabilities for HI collisions. Large acceptance, full calorimetry, high-p_T tracking

LHCb, complementary phase space for HI collisions Forward tracking, PID, calorimetry (pPb in 2013 and 2016, PbPb since 2015)

Nucleus-Nucleus Collisions at the LHC



Production and acceleration of Pb ions



▶ p (proton) ▶ ion ▶ neutrons ▶ p̄ (antiproton) →+> proton/antiproton conversion ▶ neutrinos ▶ electron

Production and acceleration of Pb ions

- ECR source: Pb²⁷⁺ (80 mA)
- RFQ: Pb²⁷⁺ to 250 A keV
- Linac3: Pb²⁷⁺ to 4.2 A MeV
- Stripper: Pb⁵³⁺
- PS Booster: Pb⁵³⁺ to 95 A MeV
- PS: Pb^{53+} to 4.25 A GeV
- Stripper: Pb⁸²⁺ (full ionisation)
- SPS: Pb⁸²⁺ to 158 A GeV
- LHC: Pb⁸²⁺ to 2.76 A TeV

Huge differences in the delivered luminosity between PbPb (~ 10^{27} cm⁻²s⁻¹) and pp (~ 10^{34} cm⁻²s⁻¹) collisions



External control parameters

Kinematical variables

Hadronic collisions are characterized by limited transfer of transverse momentum



The kinematical distribution of the produced particles are usually expressed as a function of rapidity (y) and transverse momentum (p_T)

$$y = \frac{1}{2} \ln \left(\frac{E + p_z}{E - p_z} \right) \qquad p_T = \sqrt{p_x^2 + p_y^2}$$

p_T: Lorentz-invariant with respect to a boost in the beam direction y: no Lorentz-invariant but additive transformation law \rightarrow y'=y-y_β (where y_β is the rapidity of the ref. system boosted by a velocity β)

y measurement needs particle ID (measure momentum and energy) Practical alternative: pseudorapidity (η)

$$\eta = \frac{1}{2} \log \left(\frac{|\vec{p}| + p_z}{|\vec{p}| - p_z} \right) = -\log \left[\tan \left(\frac{\theta}{2} \right) \right]$$



Geometry of a Pb-Pb collision



- <u>central collisions</u>
 - small impact parameter b
 - high number of participants \rightarrow high multiplicity
- peripheral collisions
 - large impact parameter b
 - low number of participants \rightarrow low multiplicity

Many nucleons involved

- Many nucleon-nucleon collisions
- Large interaction volume
- Many produced particles
- Few nucleons involved
- Few nucleon-nucleon collisions
- Small interaction volume
- Few produced particles

N.B. In pp there are always 2 participants



Centrality

• How do measure the impact parameter b?



Striking relation between b and multiplicity

Glauber model

Glauber model

Nuclear cross-section classes (by slicing in bins of multiplicity)







Realistic Example



light nucleons: have not participated (spectators) dark nucleons: have participated

Figure: nucl-ex/0701025

Number of participants vs b

- ❑ With respect to N_{coll}, the dependence on the nucleon-nucleon cross section is much weaker
- □ When σ_{inel} > 30 mb, practically all the nucleons in the overlap region have at least one interaction and therefore participate in the collisions



Centrality: how to access experimentally

- Two main strategies to evaluate the impact parameter in heavy-ion collisions
 - □ Measure observables related to the energy deposited in the interaction region \rightarrow charged particle multiplicity, transverse energy ($\propto N_{part}$)
 - □ Measure energy of hadrons emitted in the beam direction
 - \rightarrow zero degree energy ($\propto N_{spect}$)





Multiplicity and transverse energy

(Estimate of energy density and related to entropy)

Particle multiplicity

Most central collisions at LHC: up to 1600 charged particles per unit of η



Results normalised to pp (vacuum)

- $\sqrt{s_{NN}}=2.76$ TeV Pb+Pb, 0-5% central, | η | <0.5
- $dN_{ch}/d\eta / (\langle N_{part} \rangle / 2) = 8.3 \pm 0.4$ (sys.)

Bjorken's formula

Let's evaluate the energy density reached in the collision:

 $\frac{1}{Sc\tau_0}\frac{dE_T}{dy}$

 $\mathcal{E} =$

S=transverse dimension of the nucleus

 $\tau_{0}\text{=}$ formation time, from the hard scattering to a neutral color object ~1 fm/c



Energy Density

Let's evaluate the energy density reached in the collision:

 $\varepsilon = \frac{1}{Sc\tau_0} \frac{dE_T}{dy} \bigg|_{y=0}$

S=transverse dimension of the nucleus

 $\tau_{0}\text{=}$ formation time, from the hard scattering to a neutral color object ~1 fm/c

• experimentally, for central collisions at the LHC: $\frac{dE_T}{dy}\Big|_{y=o} \approx 1800 \,\text{GeV}$

• transverse dimension: $S \approx 160 \,\mathrm{fm}^2$ $(R_A \approx 1.2 A^{1/3} \,\mathrm{fm})$

 $\varepsilon \sim (1800/160) \, \text{GeV/fm}^3 \sim 10 \, \text{GeV/fm}^3$

More than enough for deconfinement!

N.B. only necessary, not sufficient condition ... pp collisions

Centrality dependence of $dN/d\eta$



The shapes between RHIC and LHC are very similar! Factorization in energy and centrality

Centrality dependence of $dN/d\eta$





Heavy-Ion Environment

• Measurements in an environment with $dN_{ch}/d\eta$ up to 1600 ($\sqrt{s_{NN}} = 2.76$ TeV) = 400 pp MB collisions = 1 event with 399 pile-up events (ATLAS/CMS reconstruct up to 100)



- In one collision, there are in the tracker acceptances
 - 3200 tracks in ALICE | 8000 tracks in CMS/ATLAS





Start of with qq



A gluon gets emitted at small angles



It radiates a further gluon



And so forth



Meanwhile the same happened on other side of event



And then a non-perturbative transition occurs



Giving a pattern of hadrons that "remembers" the gluon branching Hadrons mostly produced at small angle wrt $q\bar{q}$ directions or with low energy

Jets: collimated, energetic bunches of particles

PbPb, where are the jets?



Underlying events cause locally, fluctuating, high background



A Back-to-Back Jet



ATLAS, PRL105:252303,2010 Drawing: A. Mischke





Dijet Asymmetry

- How often do jets lose lot of energy?
- Quantify by dijet asymmetry
- 2 highest energy jets with $\Delta \phi > 2\pi/3$

$$A_{J} = \frac{\left| p_{T1} - p_{T2} \right|}{p_{T1} + p_{T2}} \qquad \stackrel{\mathbf{p_{T1}} = \mathbf{p_{T2}} \rightarrow \mathbf{A_{J}} = \mathbf{0}}{\underbrace{\mathbf{p_{T1}} = \mathbf{p_{T2}} \rightarrow \mathbf{A_{J}} = \mathbf{0}}_{\mathbf{1/3} \mathbf{p_{T1}} = \mathbf{p_{T2}} \rightarrow \mathbf{A_{J}} = \mathbf{0}.\mathbf{\xi}}$$

- Peripheral collisions: Pb-Pb ~ Pythia
- Central collisions: Significant difference

PRC 84 (2011) 024906 PRL105:252303,2010





Dijet Asymmetry

- How often do jets lose lot of energy?
- Quantify by dijet asymmetry
- 2 highest energy jets with $\Delta \phi > 2\pi/3$



C



Probes Traverse the QGP



The nuclear modification factor

Nuclear Modification Factor

$$R_{AA}(p_T) = \frac{Yield(Pb + Pb)}{Yield(p + p) \times \langle N_{coll} \rangle}$$

Average number of NN collisions in PbPb



In case of "No Effect": -R<1 at small momenta, production from thermal bath -R=1 at higher momenta where hard processes dominate

but if R<1 at high momenta, hot and dense medium is affecting the parton propagation 108


- Evidence for a strong parton energy loss and a large medium density at the LHC
- Behaviour reproduced by all models/calculations. R_{AA} alone is not highly discriminating

<u>Hadrons</u> constrain the parton kinematics very loosely \rightarrow <u>Jets</u> can capture the modified fragmentation process of partons: high-p_T partons interact strongly with QCD medium prior to fragmentation 109



$R_{AA}\,for\,vector\,bosons$

Fundamental check:

Electroweak probes are unmodified \rightarrow the interacting medium is colored, interacting strongly







Dead Cone Effect

- Due to kinematical constraints, gluon radiation in vacuum suppressed for angles $\theta < m/E = 1/\gamma$ by $\left(1 + \frac{m/E}{\theta}\right)^2$ Due to kinematical constraints, gluon radiation
 - Massless parton m = $0 \rightarrow$ no suppression



- Similar effect in the medium
 - Significant for charm and beauty
 - Radiative energy loss reduced by 25% (c) and 75% (b) $[\mu = 1 \text{ GeV/c}^2]$
- Implies quark mass dependence

$$R^{\pi}_{AA} < R^{D}_{AA} < R^{B}_{AA}$$



PLB519:199-206,2001 Lect. Notes Phys. 785,285 (2010)



D is stronger suppressed than $B \downarrow \rightarrow$ hint of quark mass dependence



J/ψ Suppression

• Observed at SPS in Pb-Pb collisions ($\sqrt{s_{NN}} = 17 \text{ GeV}$)



J/ψ Suppression (2)

• ... and at RHIC ($\sqrt{s_{NN}} = 200 \text{ GeV}$)

ERN



Wouldn't we expect a stronger suppression at larger $\sqrt{s_{NN}}$?

J/ψ Suppression (3)



FRN

LHC \rightarrow RHIC : $\sqrt{s_{NN}}$ 14 times larger ... but the suppression is smaller !



J/ψ Regeneration





J/ψ Regeneration





Other quarkonia states melt at different temperatures \rightarrow QGP thermometer

Particle correlations: Elliptic Flow

Non-central collisions are azimuthally asymmetric



The transfer of this asymmetry to momentum space provides a measure of the strength of collective phenomena

- Large mean free path
 - particles stream out isotropically, no memory of the asymmetry
 - extreme: ideal gas (infinite mean free path)
- Small mean free path
 - larger density gradient -> larger pressure gradient -> larger momentum
 - extreme: ideal liquid (zero mean free path, hydrodynamic limit)



Elliptic Flow

• Particles as a function of ϕ - Ψ_{RP}





- Define $v_2 = \langle \cos 2 (\phi \Psi_{RP}) \rangle$ - Second coefficient of Fourier expansion
- Ψ_{RP} common symmetry plane (for all particles)
- What if there were no correlations with Ψ_{RP} ?





Centrality Dependence

- Strong centrality dependence
- v₂ largest for 40-50%
- Spatial anisotropy very small in central collisions
- Largest anisotropy in midcentral collisions
- Small overlap region in peripheral collisions



CMS, PRC 87(2013) 014902

$v_2(p_T)$ very similar at LHC and RHIC



- system still have low viscosity
- similar behaviour





We have a liquid of quarks and gluons!

The QGP is the perfect liquid!

(not the gas of "free" quarks and gluons we expected)



Particles through Pb-Pb and p-Pb



• Strong suppression of charged hadrons in Pb-Pb (wrt pp) up to very high momenta

•Direct photons, W and Z are not quenched ... reference particles

• p-Pb results (consistent with unity up to 50 GeV) confirm that strong suppression in Pb-Pb is due to hot nuclear matter effects

The ridge in A+A collisions



In (central) A+A, the ridge is commonly interpreted as hydrodynamic "hubble" flow of initial "stringy" structures in rapidity

The structures in the $\Delta \, \phi \,$ direction are decomposed and studied by the v_n Fourier "Flow moments"



The discovery



Distinct long range correlation in η collimated around $\Delta \Phi \approx 0$





Similar for pPb (high mult), pp (high mult) and PbPb (peripheral)

Hydrodynamic flow in pp and pPb collisions?

Correlations: double ridge in p-Pb



∆φ (**rad**)

(B)

Let's go back to "fundamentals"

Different energy scales offer information on different aspects of proton internal structure



momentum fraction that the scattering particle would carry if the proton were made of ...

A point-like particle	
	and an and a second sec

momentum fraction

momentum fraction that the scattering particle would carry if the proton were made of ...



momentum fraction that the scattering particle would carry if the proton were made of ...



momentum fraction that the scattering particle would carry if the proton were made of ...



What have we learned in terms of this picture by now?

 Up and down quark "valence" distributions peaked ~1/3



What have we learned in terms of this picture by now?

- Up and down quark "valence" distributions peaked ~1/3
- Lots of sea quark-antiquark pairs and even more gluons!







$$\sigma(pp \to \pi^0 X) \propto q(x_1) \otimes g(x_2) \otimes \hat{\sigma}^{qg \to qg}(\hat{s}) \otimes D_q^{\pi^0}(z)$$

Particle production rates can be calculated using pQCD from:

- Parton distribution functions (from experiment)
- pQCD partonic scattering rates (from theory)
- "Fragmentation functions" (from experiment)

We can use *factorized* perturbative QCD (pQCD) to calculate particle production at high-energy facilities

... but then something strange happened

Charged pions produced preferentially on one or the other side with respect to the transversely polarized beam direction ... by up to 40%!!



Spin-momentum correlations: 1976 discovery in p+p collisions



... but then something strange happened

Charged pions produced preferentially on one or the other side with respect to the transversely polarized beam direction ... by up to 40%!!



Had to wait more than a decade for the birth of a new subfield:

 In 1990 D.W. Sivers departs from traditional collinear factorization assumption in pQCD and proposes correlation between the intrinsic transverse motion of the quarks and gluons and the proton's spin Spin-momentum correlations: 1976 discovery in p+p collisions



First quark distribution function describing a spin-momentum correlation in the proton

New frontier! Quark dynamics inside QCD bound states, and in their formation process

The Proton Spin Crisis



A proton has a total spin +1/2 along some axis. Most naively, you'd expect it to contain two quarks with spin +1/2 and one with spin -1/2. 1/2 + 1/2 - 1/2 = +1/2



Hence ~12% of the proton spin is carried by the spin of the quarks, the remaining spin must be carried by gluons or orbital angular momentum





In QCD bound states we need to include Spin-spin and spin-momentum correlations ζ_2

Unpolarized

$$\mathbf{f}_1 = \mathbf{O}$$

In QCD bound states we need to include Spin-spin and spin-momentum correlations $\zeta_{\rm c}$



In QCD bound states we need to include Spin-spin and spin-momentum correlations 4






Hadron tomography



Hadron tomography







Transverse Momentum Distribution Functions (TMDs)

PDFs involving transversely polarized quarks are **chiral-odd** -can only be observed experimentally in conjunction with a second chiral-odd function



Merging 3 worlds



Merging 3 worlds

For the first time we will have an experiment with 2 Interaction Points: pp + p-target (working in synergy)

Most advanced detectors Fixed Target (pol, unpol)

LHC

New perspectives in QCD and soft QCD for Cosmic Ray Physics



CERNCOURIER | International journal of high-energy physics



Home | About | News | Features | Community | Viewpoint | Reviews | Archive | Past Issues

LHCb brings cosmic collisions down to Earth



Collision and scattering events (expand for full image)

In an effort to improve our understanding of cosmic rays, the LHCb collaboration has generated high-energy collisions between protons and helium nuclei similar to those that take place when cosmic rays strike the interstellar medium. Such collisions are expected to produce a certain number of antiprotons, and are currently one of the possible explanations for the small fraction of antiprotons (about one per 10,000 protons) observed in cosmic rays outside of the Earth's atmosphere. By

measuring the antimatter compor can potentially unveil new high-er notably a possible contribution fro decay of dark-matter particles. In the last few years, space-borne study of cosmic rays have dramati

knowledge of the antimatter comp

Alpha Magnetic Spectrometer (AM



About CERN	Students & Educators		Scie	ntists	CERN community		
Accelerators	Experiments	Physics	Computing	Engineering	Updates	Opinion	

Cosmic collisions at the LHCb experiment

by Stefania Pandolfi

Posted by Stefania Pandolfi on 27 Mar 2017. Last updated 27 Mar 2017, 16.00. *Voir en français* This content is archived on the CERN Document Server

Х



An example of a fully reconstructed proton-helium collision event in the LHCb detector. The particle



... where I am the responsible

I phase: 2019-2020 II phase: 2024-...



... where I am the responsible

Conclusions

The QGP is a great tool to investigate the early Universe and to unveil the deepest "secrets" of the QCD



In November 2010, the field of ultrarelativistic nuclear collisions has entered a new era with the start of heavy-ion collisions at the LHC ... an ideal place where to study the QGP

Exiting results already achieved, the future looks bright!

Wealth of new intriguing phenomena in the medium!



Conclusions (2)

Fixed target collisions at the LHC represent a unique possibility for a laboratory for QCD in unexplored kinematic regions ... in a realistic time schedule

The \Box_{spin}^{+} project represents a fantastic challenge both for its physics potentialities and for the technology involved



Additional literature

- C.Y. Wong, Introduction to High-Energy Heavy-Ion Collisions, World Scientific, 1994 <u>http://books.google.de/books?id=Fnxvrdj2NOQC&printsec=frontcover</u>
- L. P. Csernai, Introduction to Relativistic Heavy-Ion Collisions, 1994 (free as pdf) <u>http://www.csernai.no/Csernai-textbook.pdf</u>
- E. Shuryak, The QCD vacuum, hadrons, and superdense matter, World Scientific, 2004 <u>http://books.google.de/books?id=rbcQMK6a6ekC&printsec=frontcover</u>
- Yagi, Hatsuda, Miake, Quark-Gluon Plasma, Cambridge University Press, 2005 <u>http://books.google.de/books?id=C2bpxwUXJngC&printsec=frontcover</u>
- R. Vogt, Ultrarelativistic Heavy-ion Collisions, Elsevier, 2007 <u>http://books.google.de/books?id=F1P8WMESgkMC&printsec=frontcover</u>
- W. Florkowski, Phenomenology of Ultra-Relativistic Heavy-Ion Collisions, World Scientific, 2010

http://books.google.de/books?id=4glp05n9lz4C&printsec=frontcover

 S. Sarkar, H. Satz and B. Sinha, The physics of the quark-gluon plasma, Lecture notes in physics, Volume 785, 2010 (free within CERN/university network) https://link.springer.com/book/10.1007%2F978-3-642-02286-9

or email to: Pasquale.Di.Nezza@cern.ch