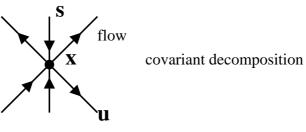
Chaotic motions and developed turbulence

Non equilibrium: Chaotic hypothesis

The attractor of a chaotic evolution can be regarded as an Anosov system $\mathcal{F} = \text{phase space}$



 $\Rightarrow \exists$ SRB distribution, *i.e.* the "statistics"

$$\frac{1}{T} \int_0^T F(S_t x) dt \xrightarrow[T \to \infty]{} \langle F \rangle \equiv \int \mu(dy) F(y) \qquad a.e.$$

 μ is singular (in general): $\dot{x} = f_E(x), \quad x \in \mathcal{F}$

$$\sigma(x) \stackrel{def}{=} - \text{divergence} f_E(x), \qquad \sigma_+ \stackrel{def}{=} \langle \sigma \rangle_{\mu} \ (\geq 0 \text{ Ruelle})$$

Idea: Chaotic hypothesis + symmetries \rightarrow "predictions"

Symmetries: time reversal, symplectic, ...

Time rev. = isometry,
$$I^2 = 1$$
 and anticommuting: $IS_t = S_{-t}I$ or $IS = S^{-1}I$

For time reversible Anosov maps or flows:

Fluctuation theorem (Cohen, G, Gentile) if $\sigma_{+} > 0$

$$p \stackrel{def}{=} \frac{1}{\tau} \int_0^\tau \frac{\sigma(S_t x)}{\sigma_+} dt, \qquad \text{probab.} (p \in \Delta) = c \, e^{\tau \max_{\Delta} \zeta(p) + O(1)} \ \, \text{and}$$

$$\zeta(-p) = \zeta(p) - p\sigma_+, \quad \text{for all } |p| < p^*, \quad p^* \ge 1$$

with no free parameters (symmetry relation).

Symplectic symmetry \Rightarrow pairing rule (Dettman, Morriss).

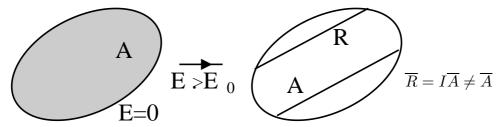
Applications

(1) Not transitive, (attractor not dense)

(2) Not reversible, (viscosity models)

(3) No pairing, (lack of symplectic structure)

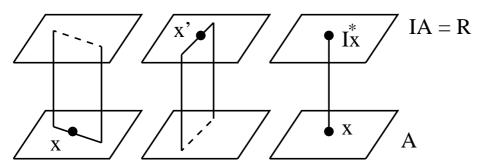
Not transitive: Chaotic hypothesis \rightarrow attracting set is a manifold.



Unstable manifold of a point on A is in A

Stable mnf "sticks out" reaching repeller R

If transversal: line from R to A is defines λ defines ix = x' commuting with S and $I^* \stackrel{def}{=} Ii$ is local time reversal on A (Bonetto, G.).



FT holds for the phase space contraction on the attractor (quite useless but ...)

Pairing symmetry: (if holding) implies

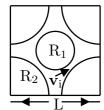
$$\zeta(-p) = \zeta(p) - \sigma_{+} \frac{M}{N}$$

2M = dimension of attractor, 2N dimension of phase space.

Lack of reversibility?

Idea: an irreversible dynamics can be equivalent to a reversible one: same statistics (G.)

Example (Drude's theory of conductivity)



$$\ddot{\mathbf{q}}_{i} = E\mathbf{u} + collisions - \boldsymbol{\vartheta}_{i}: \text{ and}$$

$$speed \ renormalized \ to \ \sqrt{3k_{B}\Theta}$$

$$or \ keep \ constant \ speed: \ \boldsymbol{\vartheta}_{i} = \alpha(\dot{\mathbf{q}})\dot{\mathbf{q}}_{i}, \ \alpha \equiv \frac{\mathbf{E}\cdot\sum\dot{\mathbf{x}}_{i}}{\sum\dot{\mathbf{x}}_{i}^{2}}$$

$$or \ use \ viscosity: \ \boldsymbol{\vartheta}_{i} = -\nu\dot{\mathbf{q}}_{i}$$

If
$$F(\mathbf{p}, \mathbf{q}) > 0$$
 is a local observable: $\frac{\mu_{\nu}(F)}{\mu_{\mathcal{E}}(F)} \xrightarrow[\mathcal{E} = \frac{3}{2}NkT]{L \to \infty, N/L = \rho} 1$

provided ν is tuned so that $\langle \mathcal{E} \rangle_{\mu_{\nu}} = \mathcal{E}_0$.

This is "equivalence of ensembles": analogy

 ν = canonical temperature and \mathcal{E} = microcanonical energy.

Application to **NS** (incompressible $\partial \cdot \mathbf{u} = 0$)

$$\dot{\mathbf{u}} + \mathbf{\underline{u}} \cdot \partial_{\mathbf{u}} \mathbf{u} = \nu \Delta \mathbf{u} - \partial p + f \mathbf{g}, \qquad R = \frac{\sqrt{fL}}{\nu}$$

Actually think of: cut off at $|\mathbf{k}| \leq K_k = L^1 R^{\frac{3}{4}}$, $N \simeq R^{\frac{9}{4}}$, i.e. OK41 is assumed. To apply the chaotic hyp. need

- (1) chaos (yes, if R large).
- (2) reversibility
- (3) pairing (because the attractor is very small)
- (1) Equivalence with reversible equations "Gaussian NS eq."

$$\dot{\mathbf{u}} + \underline{\mathbf{u}} \cdot \partial_{\alpha} \mathbf{u} = \alpha(\mathbf{u}) \Delta \mathbf{u} - \partial p + f \mathbf{g}, \qquad \alpha = \frac{\int \mathbf{u} \cdot f \mathbf{g}}{\int (\partial \mathbf{u})^2} \Rightarrow \int \mathbf{u}^2 = \mathcal{E} = const$$

Same statistics for "local observables": F local \Rightarrow F depends on finitely many Fourier components of \mathbf{u} . Same statistics as $R \to \infty$ if \mathcal{E} is chosen $= \langle \int \mathbf{u}^2 \rangle_{\mu_{\nu}}$ (equivalence)

Consequence $\langle \alpha \rangle / \nu \to 1$: only numerical tests in strongly cut off equations and d=2 (Rondoni, Segre).

Earlier She, Jackson: large numerical simulations (different reversible equation)

Other tests: are Lyapunov spectra also identical? (Rondoni, Segre, G.). Here are a few graphs in highly truncated equations (d=2)

Also the linear FR relation comes out within the precision: the approximate pairing that can be observed leads to test the slope $(1 - \frac{2M}{2N})\sigma_+$ in the GNS equations: from the theory it is expected a slope $< \sigma_+$ by the ratio of the number of negative pairs to the nuber of total pairs.

Barometric formula:

Consider the equations (incompressible NS and ED)

$$\dot{\mathbf{u}} + \mathbf{u} \cdot \partial \mathbf{u} = \nu \Delta \mathbf{u} - \partial p + f \mathbf{g}, \qquad \dot{\mathbf{u}} + \mathbf{u} \cdot \partial \mathbf{u} = -\chi \mathbf{u} - \partial p + f \mathbf{g},$$

here
$$\mathbf{u} = \sum_{\mathbf{k}} \gamma_{\mathbf{k}} e^{i\mathbf{k}\cdot\mathbf{x}}, \, \mathcal{E} = L^3 \sum_{\mathbf{k}} |\gamma_{\mathbf{k}}|^2.$$

The equivalence idea leads to think that although the statistics of the two equations are certainly different nevertheless they might coincide on an appropriate scale. The friction in NS varies with the scale \mathbf{k} and at some scale it might match that of ED.

By OK41 $v_k^3 k = constant = \eta \nu$ in NS: OK41 does not hold for ED: to fix ideas asume that at fixed cut off k_{χ} there is equipartition between the modes. Then $\langle |\gamma_{\bf k}|^2 \rangle \equiv \gamma^2$

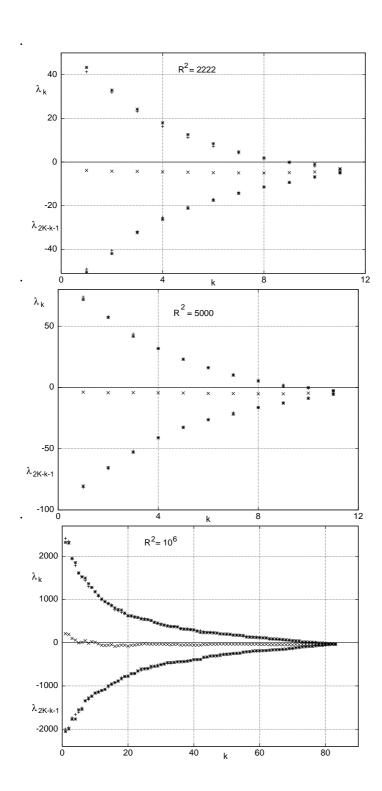
$$\frac{4\pi}{3}\gamma^2 (k_\chi \frac{L}{2\pi})^3 = \varepsilon,$$
 energy density at equipartition

$$K_E(k) = \frac{3\varepsilon}{4\pi} \frac{k^2}{k_{\chi}^3},$$
 energy density between k and $k + dk$

$$v_k^3 k = \left((kL)^3 \gamma^2 \right)^{\frac{3}{2}} k = \varepsilon^{\frac{3}{2}} k_\chi \left(\frac{k}{k_\chi} \right)^{\frac{11}{2}}, \qquad dimensionless \ dissipation \ on \ scale \ k$$

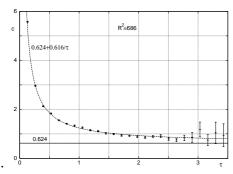
If $\varepsilon^{\frac{3}{2}}k_{\chi}\left(\frac{k}{k_{\chi}}\right)^{\frac{11}{2}}=\eta\nu$ then NS equation and ED equations have the same statistics on scale k: doubling the dissipation in NS the statistics of the two equations agree on scale 1.3 higher.

The k, or better $\log \frac{k}{k_{\chi}}$, is the analog of the height and the dim.less dissipation is type analogue of the pressure.



R^2	$\delta Q_0/\langle Q_0 \rangle_{NS}$	$\triangle \alpha$	$\triangle Q_1$	o(M)/M
800	0.005	0.030	0.053	0.068
1250	0.020	0.018	0.062	0.057
2222	0.002	0.039	0.058	0.077
4444	0.050	0.021	0.093	0.059
5000	0.010	0.008	0.058	0.033

Equivalence NS-GNS dynamics at different Reynolds numbers, column $\Delta \alpha$ to be compared with 1, cfr. [RS99]).



Evolution towards limiting slope as τ increases