Ultraviolet and infrared properties of the sine-Gordon field

General problem: define, for $x \in R^2$ and $F = c \cos(\alpha \varphi_x)$

$$P(d\varphi) = const \exp\left[-\frac{1}{2} \int (\vec{\partial}_x \varphi_x)^2 + m^2 \varphi_x^2 dx\right]$$

$$\cdot \exp\left[\lambda \int F(\varphi_x) dx\right]$$

Ultraviolet: $\int \to \int_{\Lambda}$, finite Λ , or $\Lambda = R^2, m > 0$ Infrared: $\int \to \int_{R^2} \& m = 0$

$$\lambda = 0: \text{ Gaussian field } \langle \varphi_x \varphi_y \rangle = \left(\frac{1}{-\Delta + m^2}\right)_{xy} = \begin{cases} \frac{1}{2\pi} \log \frac{1}{m|x-y|} & x-y \to 0, \quad e^{-m|x-y|} & x-y \to \infty & \text{UV} \\ const + \frac{1}{2\pi} \log \frac{1}{\mu|x-y|} & x-y \to \infty & \text{IR} \end{cases}$$

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Regularization

$$C_{xy} \rightarrow C_{xy}^{[\leq N]} = \left(\frac{1}{-\Delta+1} - \frac{1}{-\Delta+2^{2N}}\right)_{xy}$$
$$= \begin{cases} \frac{1}{2\pi} \log \frac{1}{|x-y|} & |x-y| > 2^{-N} \\ \text{smooth} & |x-y| < 2^{-N} \end{cases} C_{00}^{[\leq N]} = \frac{N}{2\pi}$$
$$c_N = e^{\frac{\alpha^2}{4\pi}N} \Rightarrow$$
$$c_N \cos(\alpha \varphi) \equiv \sum_{k=0}^{\infty} \frac{(-\alpha^2)^k}{(2k)!} : \varphi^{2k} : \stackrel{def}{=} : \cos(\alpha \varphi) :$$
$$: \varphi^k : \stackrel{def}{=} \sqrt{\frac{C_{00}^{(\leq N)}}{2}} H_k\left(\frac{\varphi}{\sqrt{2C_{00}^{(\leq N)}}}\right) \quad \text{``Wick's polynomials''}$$

$$P_{\alpha,\alpha,N}(d\varphi) = \frac{P_{0,N}(d\varphi) \exp[\lambda \int :\cos(\alpha \varphi_x^{(\leq N)}) : dx]}{Z(\alpha,\lambda,N)}$$

UV Theorem:

(1)
$$P_{\alpha,\lambda,N}(d\varphi) \xrightarrow[N \to \infty]{} P_{\alpha,\lambda}(d\varphi), \quad \alpha^{2} < 8\pi$$

(2)
$$Z(\alpha,\lambda,N) = Z_{0}(\alpha,\lambda,N)e^{\sum_{k=1}^{n}R_{k}(\alpha,N)\lambda^{k}},$$

$$8\pi(1-\frac{1}{2n}) \le \alpha^{2} < 8\pi(1-\frac{1}{2(n+1)})$$

(3)
$$Z_{0}(\alpha,\lambda,N) \xrightarrow[N \to \infty]{} Z_{0}(\alpha,\lambda), \quad R_{k}(\alpha,N) \xrightarrow[N \to \infty]{} \infty$$

(4)
$$: e^{\pm i\alpha\varphi_{x}} :, \quad :e^{\varepsilon\varphi_{x}} :, |\varepsilon|^{2} < 4\pi, \quad \text{good random v.}$$

Comment:
$$:e_x^{\varepsilon \varphi^{(\leq n)}} := e^{-\frac{(\varepsilon^2 n)}{2\pi}} e^{\varepsilon \varphi_x^{(\leq n)}} \ge 0 \Rightarrow \forall n \le \infty$$

 $\mu(dx) = \frac{:e^{\varepsilon \varphi_x^{(\leq n)}}:}{\int_{\Lambda} :e^{\varepsilon \varphi_v^{(\leq n)}}: dv} dx$ is a random measure on Λ



number errors= $|\Lambda|2^{2N} \Rightarrow \text{total} (2^{-2N}2^{\frac{\alpha^2}{4\pi}N})^2 2^{2N} |\Lambda|$ summable!

$$\int P^{(N)}(d\varphi^{(N)}) e^{-\lambda \int :\cos(\alpha \varphi_x^{(N)}) :dx}$$

= $e^{\int P^{(N)}(d\varphi^{(N)}) :\cos(\alpha \varphi_x^{(N)}) :dx} e^{\pm \lambda^2 (2^{-2N} 2^{\frac{\alpha^2}{4\pi}N})^2 2^{-2N}|\Lambda|}$

Key property of Wick's polynomials $\int P(d\zeta) : (\varphi + \zeta)^k :=: \varphi^k :$

$$\Rightarrow \int P^{(N)}(d\varphi^{(N)}) : \cos(\alpha(\varphi_x^{($$

Conclusion for $\alpha^2 < 4\pi$

$$\int P^{(N)}(d\varphi^{(N)})e^{-\lambda\int:\cos(\alpha\varphi_x^{(N)}):dx} = e^{\pm\sum_{k=0}^{\infty}\lambda^2(2^{-2k}2^{\frac{\alpha^2}{4\pi}k})^22^{-2k}|\Lambda|}$$

Beyond 4π ? expansion to order *n*: error becomes summable when $(2^{-2N}2^{\frac{\alpha^2}{4\pi}N})^n 2^{-2N} < 1$ i.e. if $\alpha^2 < 8\pi(1-\frac{1}{2n})$.

However calculations to order *n* are harder (yet possible)

Important that the free energy definition only requires constant counterms (the $R_k(\alpha, N)$): physical integretation possible (Yukawa-gas)

Infrared: Coulomb and dipoled gases

$$C_{xy}^{(-R)} = \frac{1}{(2\pi)^2} \int d^2 k \left(\frac{1}{k^2 + 2^{-2R}} - \frac{1}{k^2 + 1}\right) e^{ik(x-y)}$$
$$\lim_{R \to \infty} \left(C_{xy}^{(-R)} - C_{00}^{(-R)}\right) = W(x-y) \simeq \frac{1}{2\pi} \log|x-y|^{-1}$$
$$C_{00}^{(-R)} = \frac{R}{2\pi} \log 2,$$

$$\lim_{R\to\infty} Z^{(-R)}(I,\beta,\lambda) = \lim_{R\to\infty} \int P(d\psi^{(\geq -R)}) e^{\lambda \int_I \cos\alpha \psi_x^{(-R,0)} dx}$$

is the partition function of the neutral Coulomb gas ($\alpha^2 = \beta e^2$) Ceremade, 20/9/2012 6 The change of variables $\varphi_x^{(< R)} = \psi_{2^{R_x}}^{(\geq -R)}$ transforms the problem into an UV problem in a box $I2^{-R}$

$$Z^{(-R)}(I,\boldsymbol{\beta},\lambda) = \int P(d\boldsymbol{\varphi}^{(\leq R)} e^{\frac{\lambda}{2}2^{(2-\frac{\boldsymbol{\alpha}^2}{4\pi})R} \sum_{\sigma=\pm} \int_{I^{2-R}} e^{i\sigma\alpha \varphi_{\chi}^{(\leq -R)}} dx}$$

Theorem For $\alpha_n^2 \stackrel{def}{=} 8\pi (1 - \frac{1}{2n}) \le \alpha^2 < 8\pi (1 - \frac{1}{2(n+1)})$ the first 2*n* coefficients of the Mayer expansion are finite. Or (1) as the temperature decreases for $+\infty$ (i.e. $\alpha^2 = 0$) to $T_c = \frac{1}{\sqrt{8\pi}}$ the Mayer series coefficients become defined and (2) for $T < T_c$ all are defined (BGM). (3) For α^2 large enough $\alpha^2 > 24\pi$ the Mayer series converges (at small λ) (GawKup). (4) At high *T*: "plasma phase" (Debye screening, (Bry))

(5) At low temperature "multipole phase" or

"Kosterlitz-Thouless" phase

Hierarchical model

In this case it can be shown

(1) the Mayer series converges for $\alpha^2 > 8\pi$, i.e. in the entire multipole phase, (BGN)

(2) correlations have power law decay exponent $\frac{\alpha^2}{2\pi}$ (BGN)

Furthermore the "dipole gas" can be represented via the integral

$$\int P(d\psi^{(-R)}) = conste^{\lambda \int_I dx d\theta \cos(\alpha \partial \varphi_x \cdot \vec{n}(\theta))}$$

and in this case it has been proved convergence of the Mayer expansion (GawKup) at all α^2 with power law decay of the corelations. With our method we can only prove finiteness of all Mayer coefficients (BGN).

The hierarchical version is obtained by replacing z_{Δ} with $\pm z_{\Delta}$ depending on the position of the point $x \in \Delta$ in the left or right half. For this model: convergence of the Mayer expansion

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