# Comments on Ruelle's analysis of the OK41 turbulence theory (P. Garrido & GG)

Brief survey of OK41 (Obukov-Kolmogorov) on N-S

$$\underline{\dot{u}}_{\underline{k}} = -i \sum_{\underline{k}_1 + \underline{k}_2 = \underline{k}} \mathcal{U}_{\underline{k}_1} \cdot \underline{k}_2 \Pi_{\underline{k}} \underline{u}_{\underline{k}_2} + \underline{g}_{\underline{k}} \delta_{|\underline{k}| = k_0} \qquad k_0 \le |\underline{k}| \le k_\nu$$
$$\underline{\dot{u}}_{\underline{k}} = -\nu \underline{k}^2 \underline{u}_{\underline{k}} - i \sum_{\underline{k}_1 + \underline{k}_2 = \underline{k}} \mathcal{U}_{\underline{k}_1} \cdot \underline{k}_2 \Pi_{\underline{k}} \underline{u}_{\underline{k}_2} \qquad k_\nu < |\underline{k}| < k'_\nu$$

Scale length  $(k_0^{-1})$  "injection scale" Scale length  $\geq k_{\nu}^{-1}$  "inertial range:  $\nu$  negligible, turbulence Scale length in  $[k'_{\nu}^{-1}, k_{\nu}^{-1}]$  "dissipation": laminar motion To determine  $k_{\nu}$ ? Define Energy variation rate in modes at scales  $< \sigma^{-1}$ :

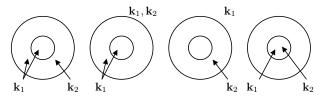
$$E^{\sigma} \stackrel{def}{=} \frac{1}{2} \frac{d}{dt} L^3 \sum_{|\underline{k}| < \sigma^{-1}} |\underline{u}_{\underline{k}}|^2 = \mathcal{L} - \mathcal{E}_{\sigma,\sigma'} - \mathcal{E}_{\sigma,\sigma',\infty} + 0$$

 $\mathcal{L} = \text{forcing work (on scale } k_0^{-1}); \ 2L^3 \operatorname{Re} \underline{\bar{g}}_{\underline{k}_0} \cdot \underline{u}_{\underline{k}_0} \\ \mathcal{E}_{\sigma,\sigma'} = \text{ work on modes } |\underline{k}| < \sigma \text{ due to } |\underline{k}| \in [\sigma, \sigma'):$ 

$$\mathcal{E}_{\sigma,\sigma'} = iL^3 \sum_{|\underline{k}_3| < \sigma} \left( \sum_{\substack{\underline{k}_1 + \underline{k}_2 + \underline{k}_3 = 0\\ |\underline{k}_1| \in [k_0,\sigma'), |\underline{k}_2| \in [\sigma,\sigma']}} \underline{u}_{\underline{k}_1} \cdot \underline{k}_2 \underline{u}_{\underline{k}_2} \right) \cdot \underline{u}_{\underline{k}_3}$$

 $\begin{array}{l} \displaystyle \boldsymbol{\mathcal{E}}_{\boldsymbol{\sigma},\boldsymbol{\sigma}',\infty} = \text{work "modes } |\underline{k}| < \boldsymbol{\sigma} \text{ from } |\underline{k}_2| \in [\boldsymbol{\sigma}',\infty) \text{ or } \\ \displaystyle |\underline{k}_2| \in [\boldsymbol{\sigma},\boldsymbol{\sigma}'), |\underline{k}_1| \in [\boldsymbol{\sigma}',\infty) \end{array}$ 

0=internal work inside  $|\underline{k}| < \sigma$ 



 $\mathcal{E}_{\sigma,\sigma'}$  first,  $\mathcal{E}_{\sigma,\sigma',\infty}$  second and third, 0 fourth.

The second and third involve pairs of  $\underline{u}_{\underline{k}}$  separated by the gap.

0) **Homogeneous turbulence**:  $\langle \mathcal{E}_{\sigma,\sigma',\infty} \rangle = 0$  if  $\sigma$  and  $\sigma' = \kappa \sigma$  are  $\langle k_{\nu}$ , in asymptotic regime.

Provided  $\kappa$  large enough: to neglect exchanges by noncontiguous scales. Hence  $\mathcal{E}_{\sigma,\sigma'} = L^3 \varepsilon$  constant.

Energy dissipated only on scales  $|\underline{k}| > k_{\nu}$ , "cascading" without dissipation from large scales to small ones.

 $\begin{array}{l} \varepsilon \text{ can only depend on velocity } v_{\sigma} \text{ characteristic of lenfgth} \\ \text{scale } \sigma \text{ and on } \sigma \text{: hence } \varepsilon = \frac{v_{\sigma}^3}{\sigma} \\ v_{\sigma}^2 \stackrel{def}{=} \langle \left(\frac{1}{\Delta_{\sigma}} \int_{\Delta_{\sigma}} (\underline{u}(\underline{x}) - \underline{u}(\underline{x}_0)) d\underline{x}\right)^2 \rangle \end{array}$ 

2) <u>k</u>-th FT of velocity var. assumed statistically indep, 3) Velocity var. in same scale boxes assumed stat. indep. Then OK41  $\Rightarrow$  energy is  $K(\sigma)d\sigma = const \varepsilon^{\frac{2}{3}} \sigma^{-\frac{5}{3}}$ 

Reynolds num. on scale  $\sigma$ :  $R_{\sigma} = \frac{v_{\sigma}\sigma}{\nu}$  is (as  $\frac{v_{\sigma}^3}{\sigma} = const$ !)

$$R_{\sigma} = \frac{v_{\sigma}\sigma}{\nu} = \frac{v_L L}{\nu} (\frac{\sigma}{L})^{\frac{4}{3}}$$

The Kolmogorv scale is then defined by the scale at which motion is laminar, *i.e.*  $R_{\sigma} = 1$ : this is

$$k_{\nu} = \sigma_{\nu}^{-1} = L^{-1} R^{\frac{3}{4}}$$

Dimension of attractor  $\propto R^{\frac{9}{4}}$ 

As 
$$\frac{v_{\sigma}^3}{\sigma} = \varepsilon \Rightarrow \tau_p = 0$$
  
 $\langle v_{\sigma}^p \rangle = \varepsilon \sigma^{\frac{p}{3}(1+\tau_p)} \equiv \varepsilon \kappa^{-\frac{n}{3}(1+\tau_p)}, \qquad \sigma = L\kappa^{-n}$ 

Of course  $v_{\sigma}$  fluct., observably,  $\Rightarrow \tau_p \not\equiv 0$ : except for p = 1. Why? Ruelle, [3, 4] : 0)  $v_{\sigma}^3/\sigma$  be a r.v. with constant average on scale  $\sigma = \kappa^{-n}$ 1) In a box  $\Delta_{\sigma}$  there are  $\kappa^3$  boxes of scales  $\frac{\sigma}{\kappa}$ 

$$\int v_n^3 P(v_n|v_{n-1}) dv_n \equiv \kappa^{-1} v_{n-1}^3, \qquad \text{constraint}$$

2) which one? "Boltzmannian": *i.e.* maximizing entropy under constraint  $\int (-\log P(v|w) - \lambda v^3) P(v|w) dv = \max \Rightarrow$ 

$$W_n \stackrel{def}{=} v_n^3, \qquad P(W_n | W_{n-1}) = e^{-\frac{\kappa W_n}{W_{n-1}}} \frac{\kappa dW_n}{W_{n-1}}$$

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$$\mathcal{P}_r(d\underline{W}) = \prod_{m=0}^n \prod_{i=1}^{\kappa^m} \frac{dW_{i,m+1}}{\kappa W_{i'm}} e^{-\frac{W_{i,m+1}}{\kappa W_{i',m}}}$$

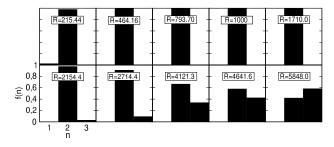
Predictions? if  $\sigma_n \stackrel{def}{=} \kappa^{-n}$ 

$$\langle v_n^p \rangle = \kappa^{-\frac{np}{3}} \kappa^{-n\tau_p}, \ \tau_p = -\frac{\log \Gamma(1+p)}{\log \kappa}$$

One free parameter:  $\kappa$ . Experimental data  $\exists$  for p < 18: Fit gives  $\kappa \sim 22$ , "very large" (!?).

Surprising:  $\Rightarrow$  at "moderate" *R*'s the number of scales to reach K-scale can be calculated and is at most 2 (*i.e.* 3, 4 exist but are very rare).

Recall that this all started from the new finding (Schumacher et al.) that universality starts being manifest already at small *Re*.



Distribution of events that reach the Kolmogorov scale  $\kappa^{-n}$  for different values of the Reynold's numbers R and  $\kappa = 22$ . The total number of events is  $10^{12}$ .

Another quantity measured (simulations) is "radial" velocity component distribution at Kolmogorov scale (or at any inertial scale where *Re* goes below a prefixed treshold):

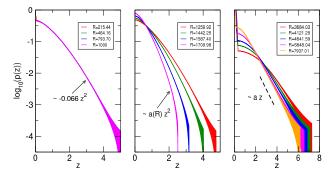
$$n(\underline{W}) = \text{first } i \text{ s.t.} R_i = \frac{W_i^{\frac{1}{3}}}{\kappa^{-i}} \equiv \frac{v_i}{\kappa^{-i}} < 1$$

 $\kappa^{-i}$  is the K. scale. Ask probability dens. that at the K. scale radial velocity  $\frac{v_i \cos \theta}{\langle v_i^2 \cos^2 \theta \rangle^{\frac{1}{2}}} \in [z, z + dz]$  $p(z) = \frac{1}{2} \mu_2^{\frac{1}{2}} P(\mu_2 |z|)$ 

In (Schumacher et al, 2014) it was found

- 1) the p(z) is Gaussiann at low Re.
- 2) but at moderate  $Re. \log p(z)$  develops a linear tail
- 3) achieved via impressive simulations

Ruelle's distribution tests the above result

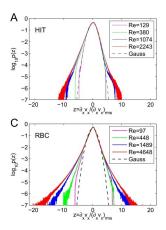


 $\begin{array}{l} \log_{10} p(z) \text{ distribution for different Reynold's numbers.} \\ \text{Center: } a(1259.92) = -0.08, \ a(1442.25) = -0.12, \ a(1587.40) = -0.24, \ a(1709.98) = -0.52. \\ \text{Right: } a(3684.03) = -0.51, \ a(4121.29) = -0.55, \ a(4621.59) = -0.58, \end{array}$ 

a(5848.04) = -0.61, a(7937.01) = -0.64.

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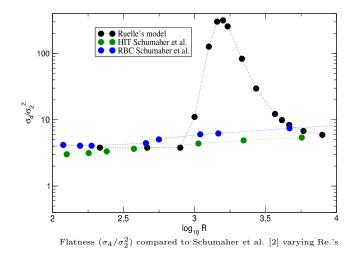
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p(z) by Schumaher et al. for varying Re.'s

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Computation of  $P^*(\xi)$  is a problem on extreme events

$$W_n^{\frac{1}{3}}\kappa^n \equiv W_0^{\frac{1}{3}}w_1w_2\dots w_n, \qquad w_k = \frac{\kappa^k W_k}{\kappa^{k-1}W_{k-1}}$$
$$P_n(dw_1,\dots dw_n) = \prod_{i=1}^n p(w_i)dw_i$$

for i.d. random variables; extreme event, \*, is

 $\begin{aligned} \frac{\log W_0}{3} + \varphi \sum_{\substack{i=1\\j\equiv 1}}^m \log w_i > 0 \ \forall m < n \\ & \underbrace{\log W_0}{3} + \sum_{i=1}^n \log w_i < 0 \end{aligned}$ with probability:  $P^*(\xi) = \sum_{n=1}^\infty \int^* P_n(d\underline{w}) \delta(\sum_{i=1}^n \log w_i - \xi) \end{aligned}$ 

If p(w) is not faster than exp., as in R, analysis of  $P^*$  involves the Gumbel distribution  $\Phi(t) = e^{3t-e^{3t}}$ 

It should also be true (?) that if values of n really matter (i.e. at much larger Re.)

no matter which p(w) is used provided with exponential tail

the result will be the same as with R's  $p(w) = e^{-w}dw$ .

If so the *"Boltzmannian prescription*" would be set in a conceptually general perspective,

And the universality of the tails of the dissipation-pdf, *i.e.* pdf of  $\frac{v_{\sigma}^3}{\sigma}$  in the inertial range may be perhaps more clear.

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